

Částice částice částice

částice částice částice částice ...



Co jsem slíbil...

V přednášce se podíváme na objevy prvních částic a antičástic, podivných částic a neutrin, a jak můžeme určit, která částice okolo nás právě "letí", čím se liší a čím jsou zajímavé. Také si připomeneme výročí 50 let od objevu jedné z nich, která znamenala listopadovou revoluci v částicové fyzice. A nakonec se podíváme kritickým okem na částicové experimenty ze seriálu Problém tří těles:)

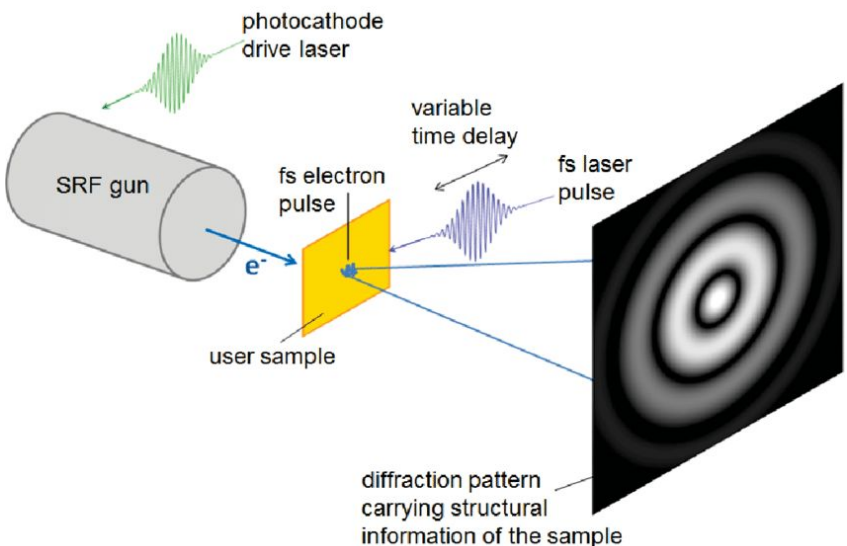
When you incorporate all the feedback from every critique.

In your Research Paper

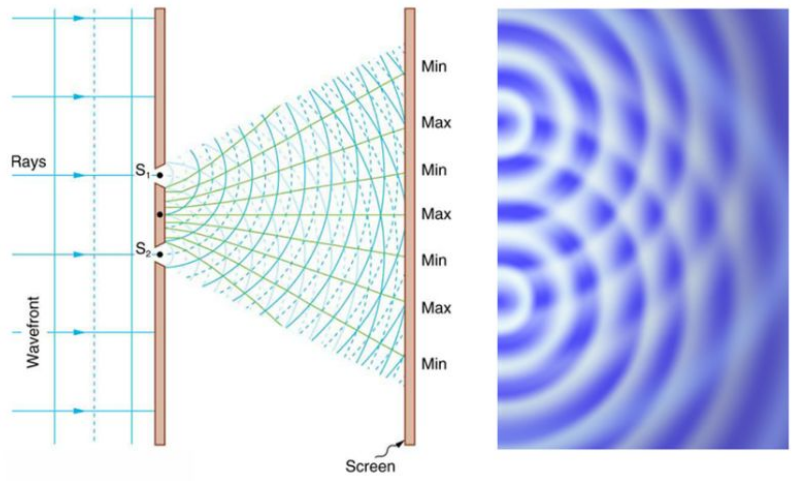


Částice nebo vlna?

Elektrony

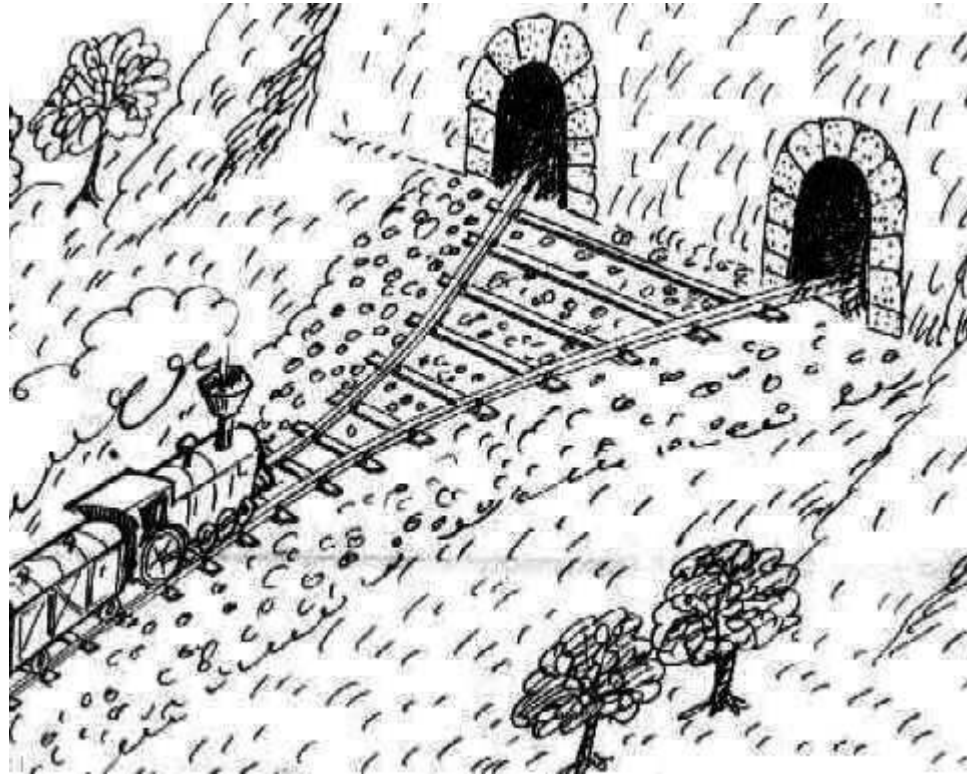


Světlo



https://www.researchgate.net/figure/The-setup-of-an-ultrafast-electron-diffraction-experiment-41_fig5_333639953
<https://courses.lumenlearning.com/suny-physics/chapter/27-3-youngs-double-slit-experiment/>

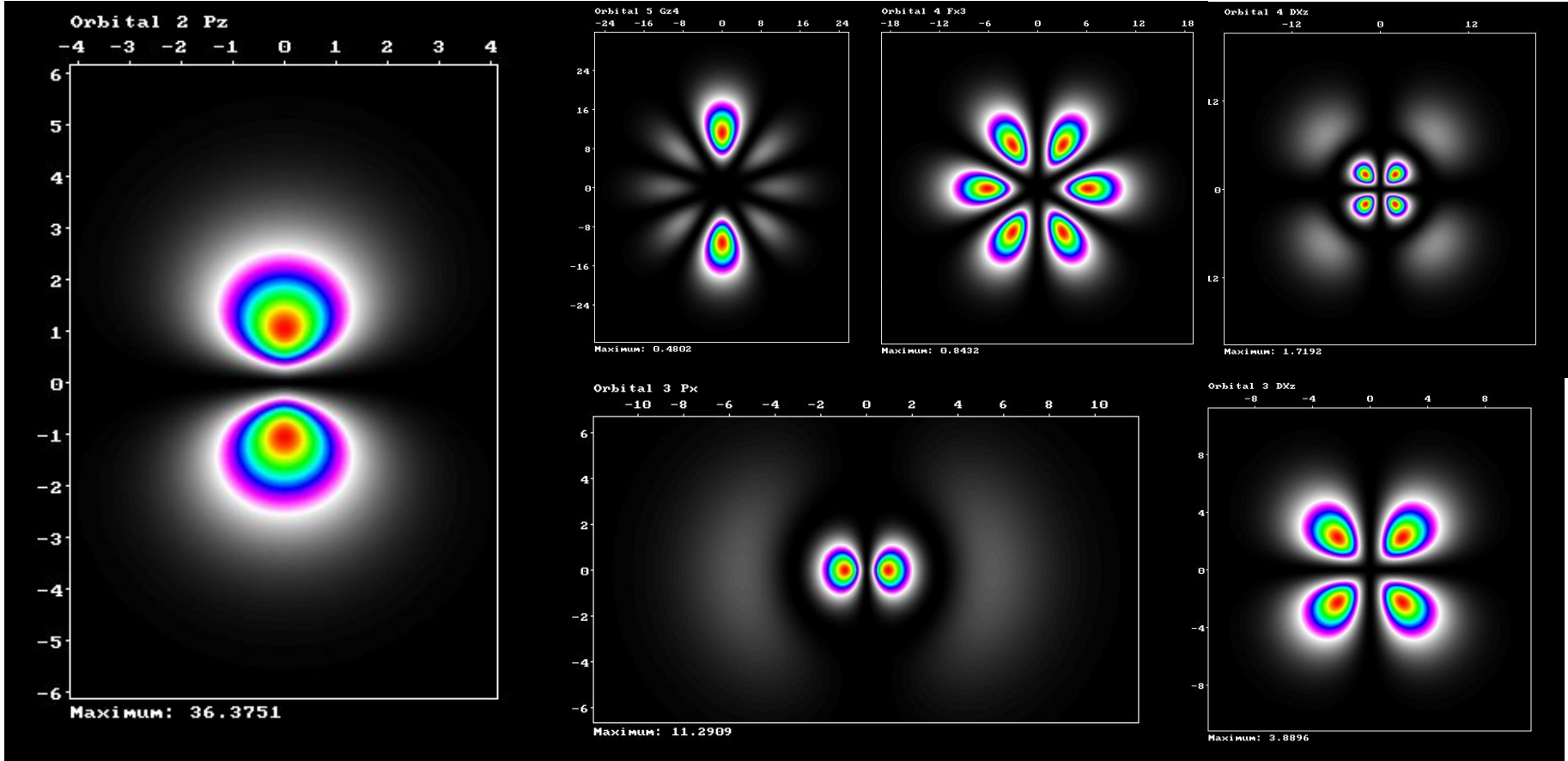
Částice nebo vlna?



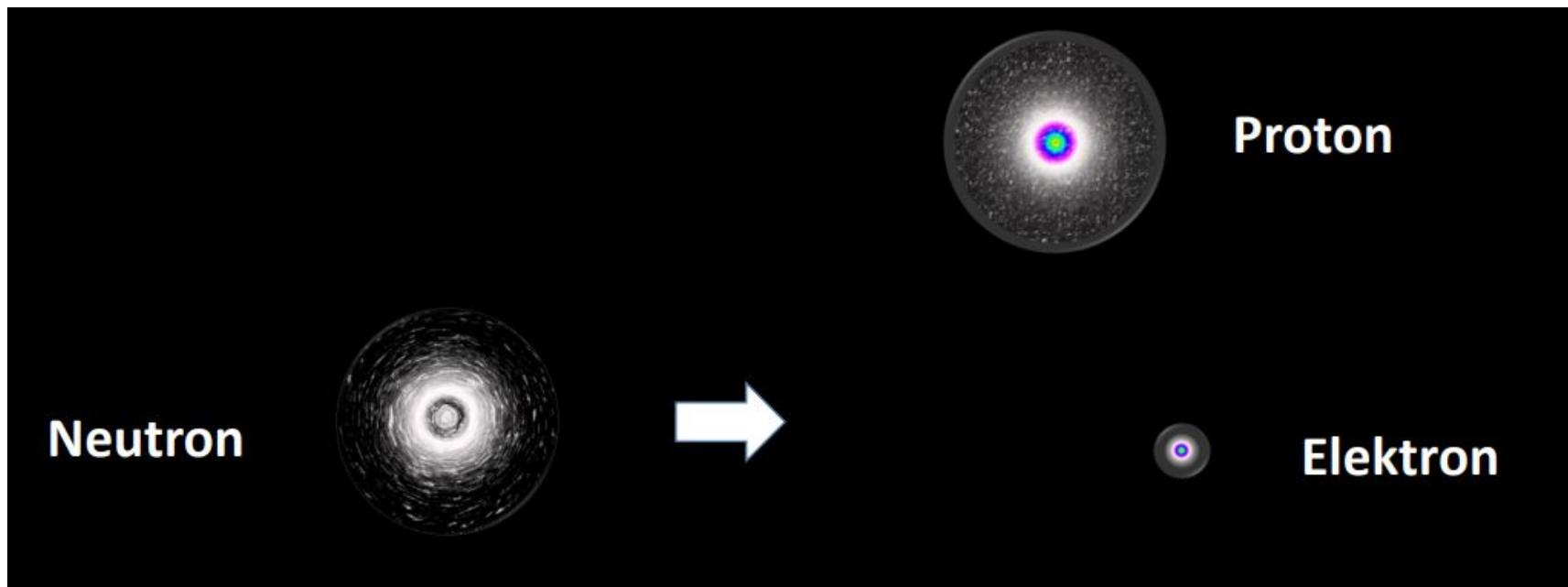
Jaké znáte částice?

Jaké znáte antičástice?

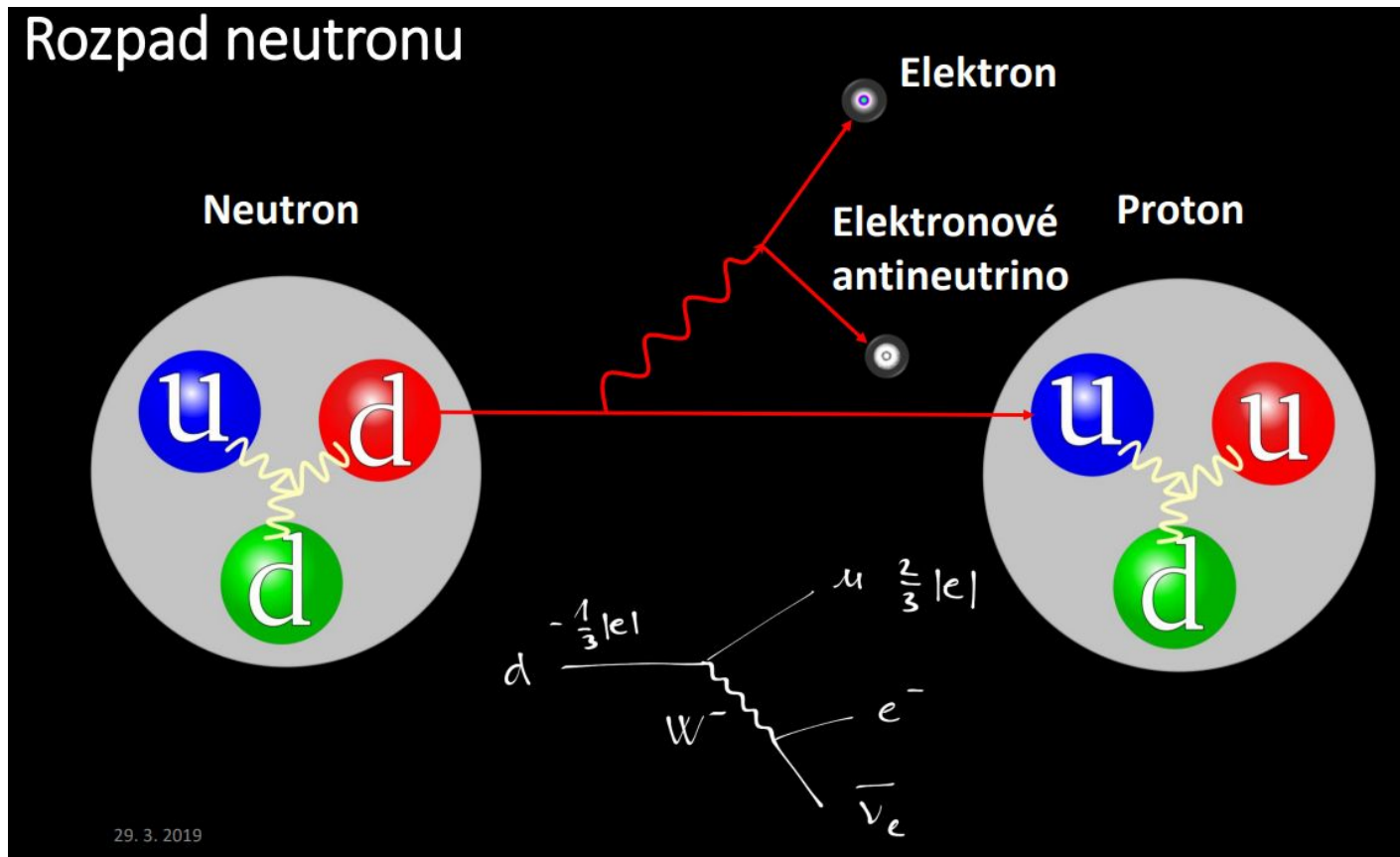
Atom – stojaté vlny elektronu okolo jadra



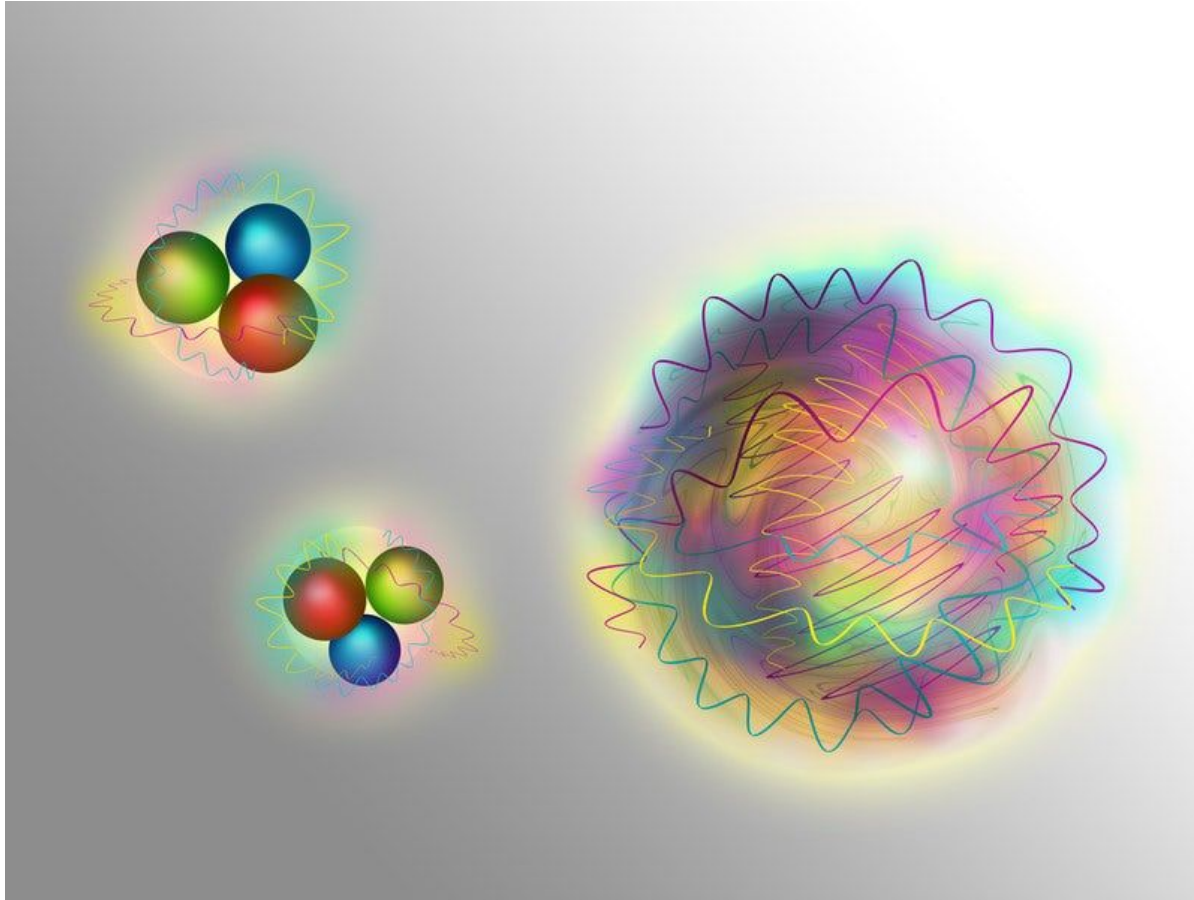
Rozpad neutronu





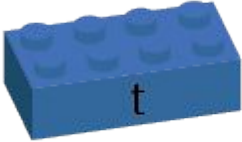



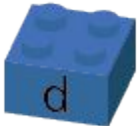
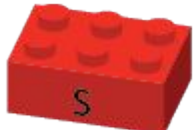
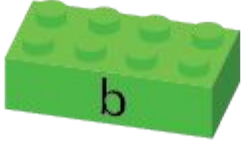
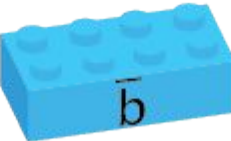


Rozpad neutronu



Proton

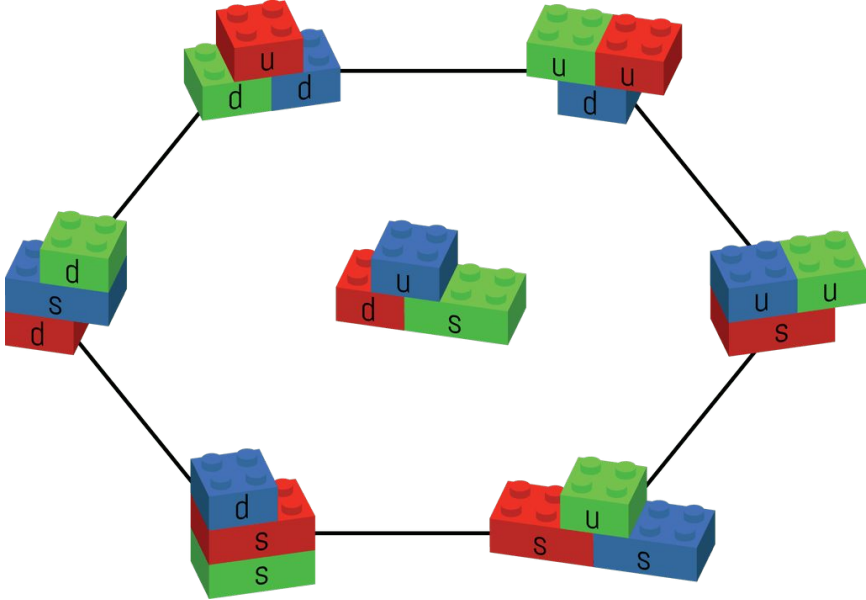


Kvarky a antikvarky

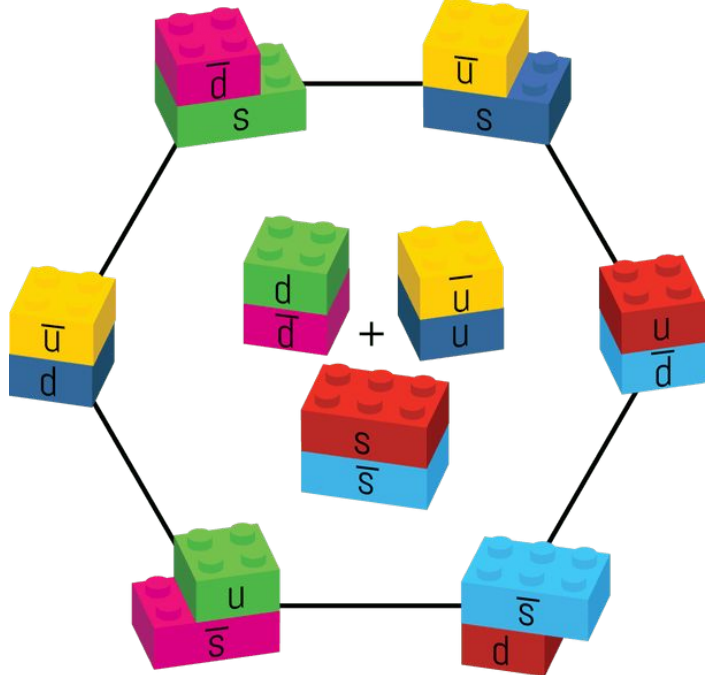
Quarks			Antiquarks		
1	2	3	3	2	1
Up 	Charm 	Top 	Antitop 	Anticharm 	Antiup 
Down 	Strange 	Bottom 	Antibottom 	Antistrange 	Antidown 

Hadrony, skládání kvarků u, d, s

Baryony = qqq

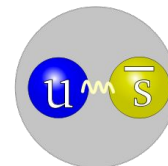
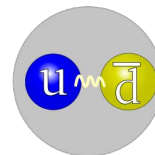
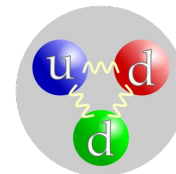
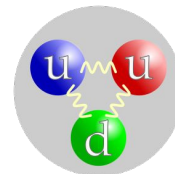
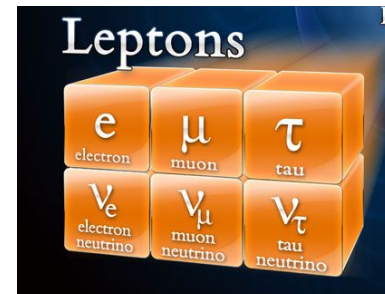


Mesony = q + anti-q

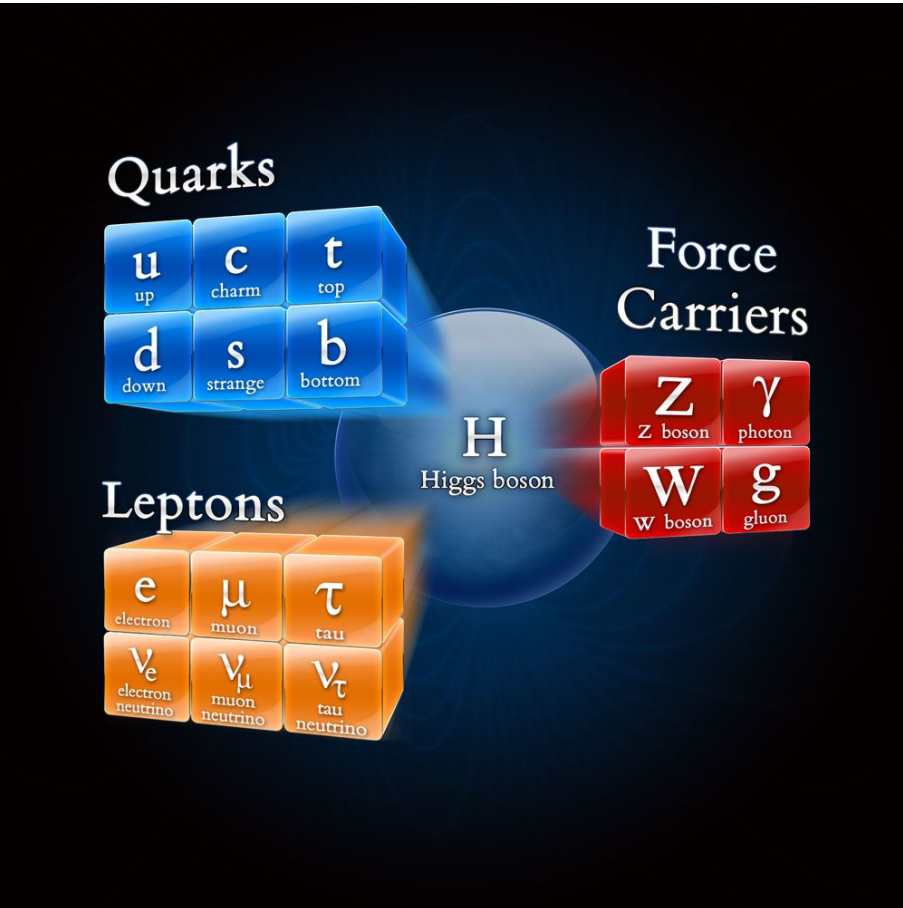
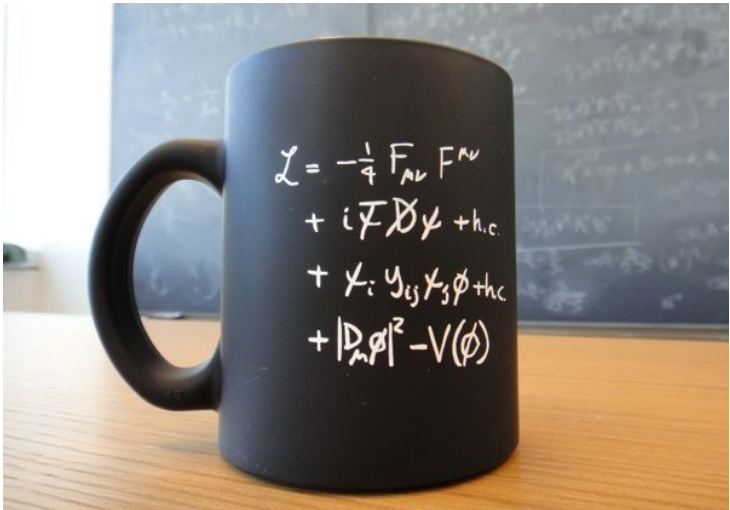


“Stabilní” částice: pozorovatelné v detektoru

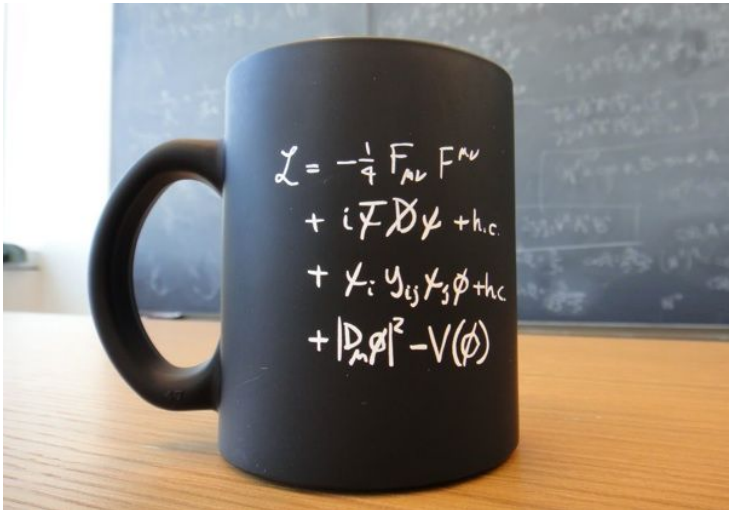
- Fotony, $m=0$
- Elektrony e^- , $mc^2 = 0.511 \text{ MeV}$
 - Pozitrony e^+
- Protony p^+ , 938.3 MeV
- Neutrony n^0 , 939.6 MeV
- Neutrína – málo interagující
- Miony μ^\pm – doba života $2\mu\text{s}$, $c\tau = 600\text{m}$. $mc^2 = 105.6 \text{ MeV}$
 - 200x těžší než elektrony, pronikavější, kosmické záření.
- Piony, 140 MeV
 - π^\pm – doba života 26ns , $c\tau = 7.8\text{m}$
 - π^0 – rychle se rozpadá
- Kaony
 - K^+ , K^- , $c\tau = 3.7\text{m}$
 - K^0 , anti- K^0



Quarks, Leptons, Gauge Bosons and the BEH boson



Quarks, Leptons, Gauge Bosons and the BEH boson



WHAT PART OF

$$-\frac{1}{2} \partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_a f^{abc} \partial_\mu g_\nu^b \partial_\nu g_\mu^c - \frac{1}{2} g_a^2 f^{abc} f^{ade} g_\nu^b g_\nu^c g_\mu^d g_\mu^e + \frac{1}{2} i g_a^2 (\bar{\psi} \gamma^\mu \psi) g_\mu^a$$

$$\bar{C}^a \partial^\mu C^a + g_a f^{abc} \partial_\mu C^a g_\nu^b g_\nu^c - \partial_\mu W_\nu^+ \partial_\mu W_\nu^- - M^2 W_\mu^+ W_\mu^- - \frac{1}{2} \partial_\mu Z_\nu^0 \partial_\nu Z_\mu^0 - \frac{1}{2} M^2 Z_\mu^0 Z_\mu^0 -$$

$$\frac{1}{2} \partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2} \partial_\mu H \partial_\mu H - \frac{1}{2} m_H^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2} \partial_\mu \phi^0 \partial_\mu \phi^0$$

$$+ \frac{1}{2} M^2 \phi^+ \phi^- - \beta_h [2\lambda \phi^4 + 2\lambda H^4 + \frac{1}{2} (H^2 + \phi^+ \phi^- + 2\phi^+ \phi^-)] + 2\lambda \phi^4 - \alpha_h - i g_{cW} [\partial_\mu Z_\nu^0 (W_\mu^+ W_\nu^- -$$

$$W_\mu^- W_\nu^+) - Z_\nu^0 (\partial_\mu W_\nu^+ - W_\mu^+ \partial_\nu W_\nu^+) + Z_\nu^0 (W_\mu^+ \partial_\nu W_\nu^- - W_\nu^- \partial_\mu W_\mu^+)] - i g_{sW} \partial_\nu A_\mu (W_\mu^+ W_\nu^- -$$

$$W_\mu^- W_\nu^+) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\mu W_\mu^+) + A_\nu (W_\mu^- \partial_\nu W_\mu^+ - W_\nu^+ \partial_\mu W_\mu^-)] - \frac{1}{2} g^2 W_\mu^+ W_\nu^- W_\nu^+ W_\mu^- +$$

$$\frac{1}{2} g^2 W_\mu^+ W_\nu^- W_\nu^+ W_\mu^- + g^2 Z_\nu^0 (Z_\mu^0 W_\nu^+ W_\mu^- - Z_\mu^0 W_\nu^- W_\mu^+) + g^2 Z_\nu^0 (A_\mu W_\mu^+ A_\nu W_\nu^- -$$

$$A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_{\theta W} A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - W_\nu^- W_\mu^+) - 2 A_\mu Z_\nu^0 W_\mu^+ W_\nu^- - g \alpha [H^2 +$$

$$H \phi^0 \phi^0 + 2H \phi^+ \phi^-] - \frac{1}{2} g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- +$$

$$2(\phi^0)^2 H^2] - g M W_\mu^+ W_\nu^- H - \frac{1}{2} g \frac{M}{\Lambda} Z_\nu^0 Z_\mu^0 H - \frac{1}{2} i g [W_\mu^+ (\phi^0 \partial_\nu \phi^- - \phi^- \partial_\nu \phi^0) - W_\nu^- (\phi^0 \partial_\mu \phi^+ -$$

$$\phi^+ \partial_\mu \phi^0)] + \frac{1}{2} i g [W_\mu^+ (H \partial_\nu \phi^- - \phi^- \partial_\nu H) - W_\nu^- (H \partial_\mu \phi^+ - \phi^+ \partial_\mu H)] + \frac{1}{2} g^2 [Z_\nu^0 (H \partial_\mu \phi^0 -$$

$$\phi^0 \partial_\mu H) - i g \frac{M}{\Lambda} Z_\nu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + i g s_{\theta W} M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - i g \frac{1-2c_{2\theta}}{2c_{2\theta}} Z_\nu^0 \phi^+ \partial_\nu \phi^- -$$

$$\frac{1}{2} g^2 \frac{1}{2c_{2\theta}} Z_\nu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_{2\theta}^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2} g^2 \frac{1}{2c_{2\theta}} Z_\nu^0 Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) -$$

$$\frac{1}{2} i g^2 \frac{2s_{2\theta}}{c_{2\theta}} Z_\nu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2} g^2 s_{\theta W} A_\mu \phi^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) + \frac{1}{2} i g^2 s_{\theta W} A_\mu H (W_\mu^+ \phi^-$$

$$- W_\mu^- \phi^+) - g^2 \frac{2s_{2\theta}}{c_{2\theta}} (2c_{2\theta}^2 - 1) Z_\nu^0 A_\mu \phi^+ \phi^- - g^2 s_{\theta W}^2 A_\mu \phi^+ \phi^- - \tilde{e}^\dagger (\gamma^0 + m_e) e^+ -$$

$$\tilde{\nu}^\dagger \gamma^0 \nu^+ - \tilde{u}^\dagger (\gamma^0 + m_u) u^+ - \tilde{d}^\dagger (\gamma^0 + m_d) d^+ + i g s_{\theta W} A_\mu [-(e^+ \gamma^\mu e^+) + \frac{2}{3} (\tilde{u}^\dagger \gamma^\mu \tilde{u}^\dagger) -$$

$$\frac{1}{3} (\tilde{d}^\dagger \gamma^\mu \tilde{d}^\dagger)] + \frac{2g}{\Lambda} Z_\nu^0 [(\tilde{\nu}^\dagger \gamma^\mu \nu^+ + \tilde{u}^\dagger \gamma^\mu u^+) + (e^+ \gamma^\mu (4s_{2\theta}^2 - 1 - \gamma^5) e^+) - (\tilde{u}^\dagger \gamma^\mu (\frac{2}{3} s_{2\theta}^2 -$$

$$1 - \gamma^5) \tilde{u}^\dagger) + (\tilde{d}^\dagger \gamma^\mu (1 - \frac{2}{3} s_{2\theta}^2 - \gamma^5) \tilde{d}^\dagger)] + \frac{1}{\Lambda} W_\mu^+ [(\nu^\dagger \gamma^\mu (1 + \gamma^5) e^+) - (\tilde{u}^\dagger \gamma^\mu (1 +$$

$$\gamma^5) C_{\nu u} d^+)] + \frac{g}{\sqrt{2}} W_\mu^- [(e^+ \gamma^\mu (1 + \gamma^5) \nu^+) + (\tilde{d}^\dagger C_{\nu d} \gamma^\mu (1 + \gamma^5) \tilde{u}^\dagger)] + \frac{1}{\sqrt{2}} M [-\phi^+ (\tilde{\nu}^\dagger (1 -$$

$$\gamma^5) e^+) + \phi^- (e^+ (1 + \gamma^5) \nu^+) - \frac{2}{\Lambda} \tilde{u}^\dagger [H (e^+ e^+) + i \phi^0 (e^+ \gamma^5 e^+)] + \frac{1}{\Lambda} \tilde{d}^\dagger \phi^+ [-m_d (\tilde{u}^\dagger C_{\nu u} (1 -$$

$$\gamma^5) d^+) + m_u (\tilde{u}^\dagger C_{\nu u} (1 + \gamma^5) d^+)] + \frac{1}{\Lambda} \tilde{d}^\dagger \phi^- [m_d (\tilde{d}^\dagger C_{\nu d}^+ (1 + \gamma^5) \tilde{u}^\dagger) - m_u (\tilde{d}^\dagger C_{\nu d}^+ (1 -$$

$$\gamma^5) \tilde{u}^\dagger) - \frac{2}{\Lambda} \tilde{u}^\dagger H (u^+ u^+) - \frac{2}{\Lambda} \tilde{d}^\dagger H (d^+ d^+) + \frac{g}{\Lambda} m_u \phi^0 (\tilde{u}^\dagger \gamma^5 u^+) - \frac{g}{\Lambda} m_d \phi^0 (\tilde{d}^\dagger \gamma^5 d^+)]$$

$$+ X^+ (\partial^2 - M^2) X^+ + X^- (\partial^2 - M^2) X^- + X^0 (\partial^2 - M^2) X^0 + Y \partial^2 Y + i g_{cW} W_\mu^+ (\partial_\nu X^0 X^\nu -$$

$$\partial_\nu X^+ X^\nu) + i g_{sW} W_\mu^+ (\partial_\nu \bar{X} X^\nu - \partial_\nu X^+ X^\nu) + i g_{cW} W_\mu^- (\partial_\nu X^0 X^\nu - \partial_\nu X^- X^\nu) +$$

$$i g_{sW} W_\mu^- (\partial_\nu X^- X^\nu - \partial_\nu \bar{X} X^\nu) + i g_{cW} Z_\nu^0 (\partial_\mu X^+ X^\mu - \partial_\mu X^- X^\mu) + i g_{sW} A_\mu (\partial_\nu X^+ X^\nu -$$

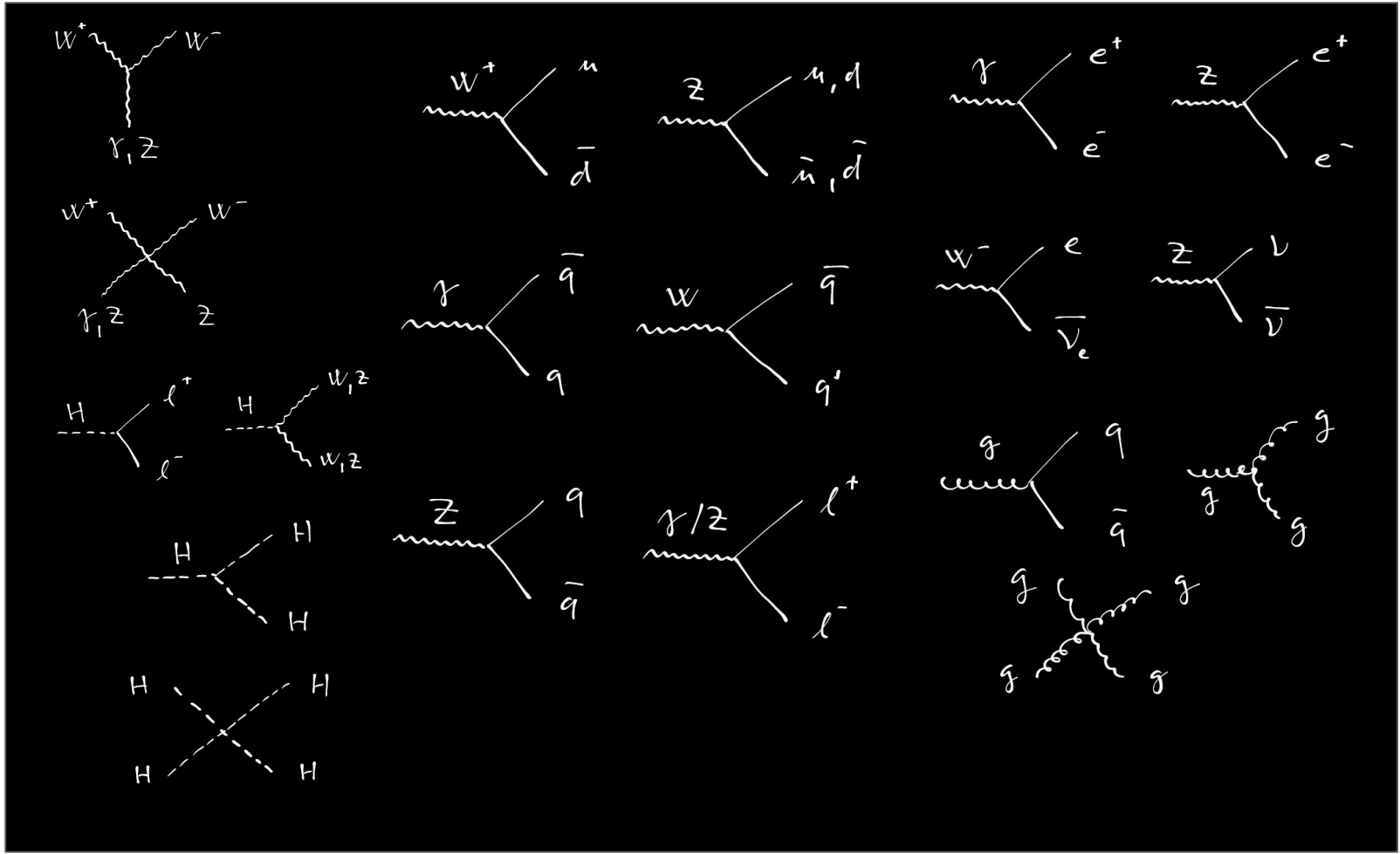
$$\partial_\nu X^- X^\nu) - \frac{1}{2} g M [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{2} \bar{X}^0 X^0 H] + \frac{1-2c_{2\theta}}{2c_{2\theta}} i g M [\bar{X}^+ X^0 \phi^+ -$$

$$X^- X^0 \phi^-] + \frac{1}{2c_{2\theta}} i g M [X^0 X^- \phi^+ - X^0 X^+ \phi^-] + i g M s_{\theta W} [X^0 X^- \phi^+ - X^0 X^+ \phi^-] +$$

$$\frac{1}{2} i g M \bar{X}^+ X^+ \phi^0 - X^- X^- \phi^0]$$

DO YOU NOT UNDERSTAND?

Quarks, Leptons, Gauge Bosons and the BEH boson

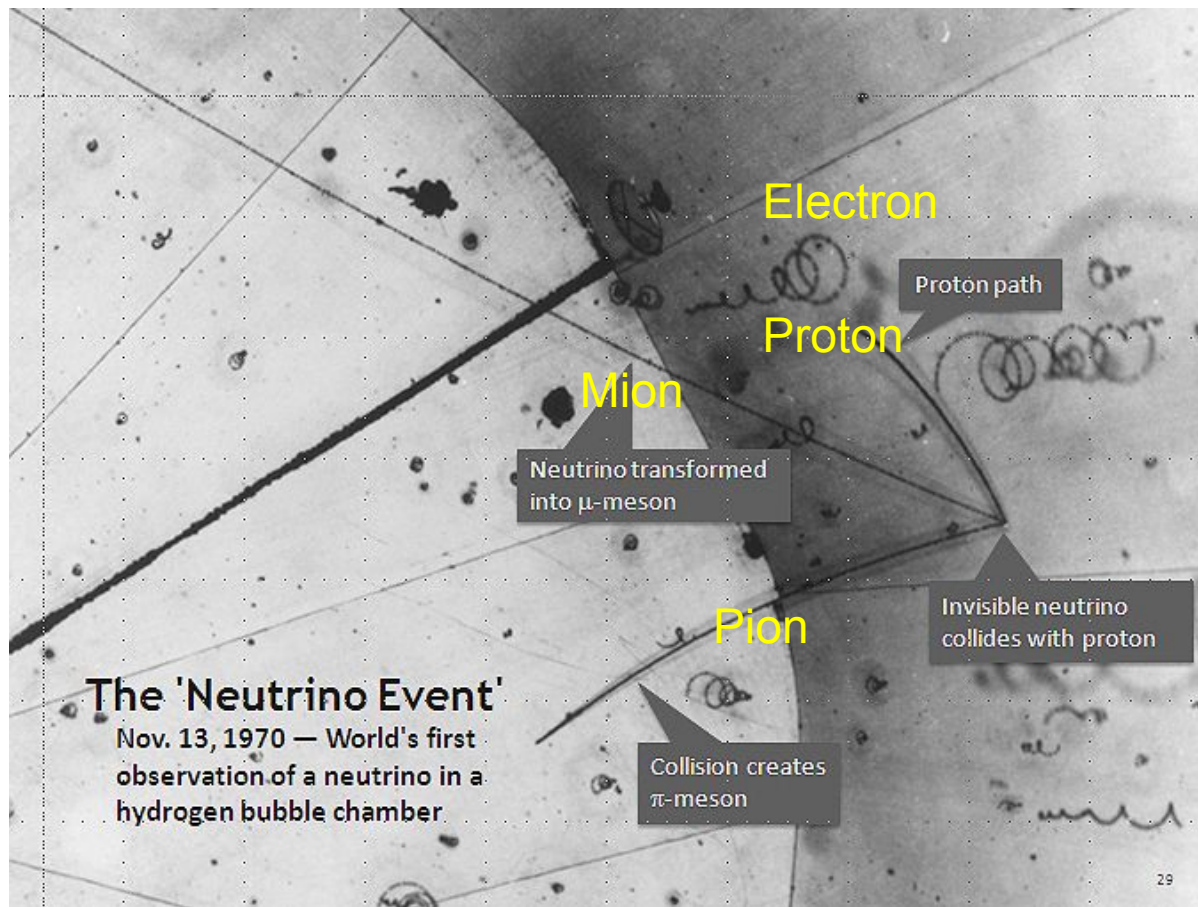


Jaké vidíme částice? :: Mlžná komora



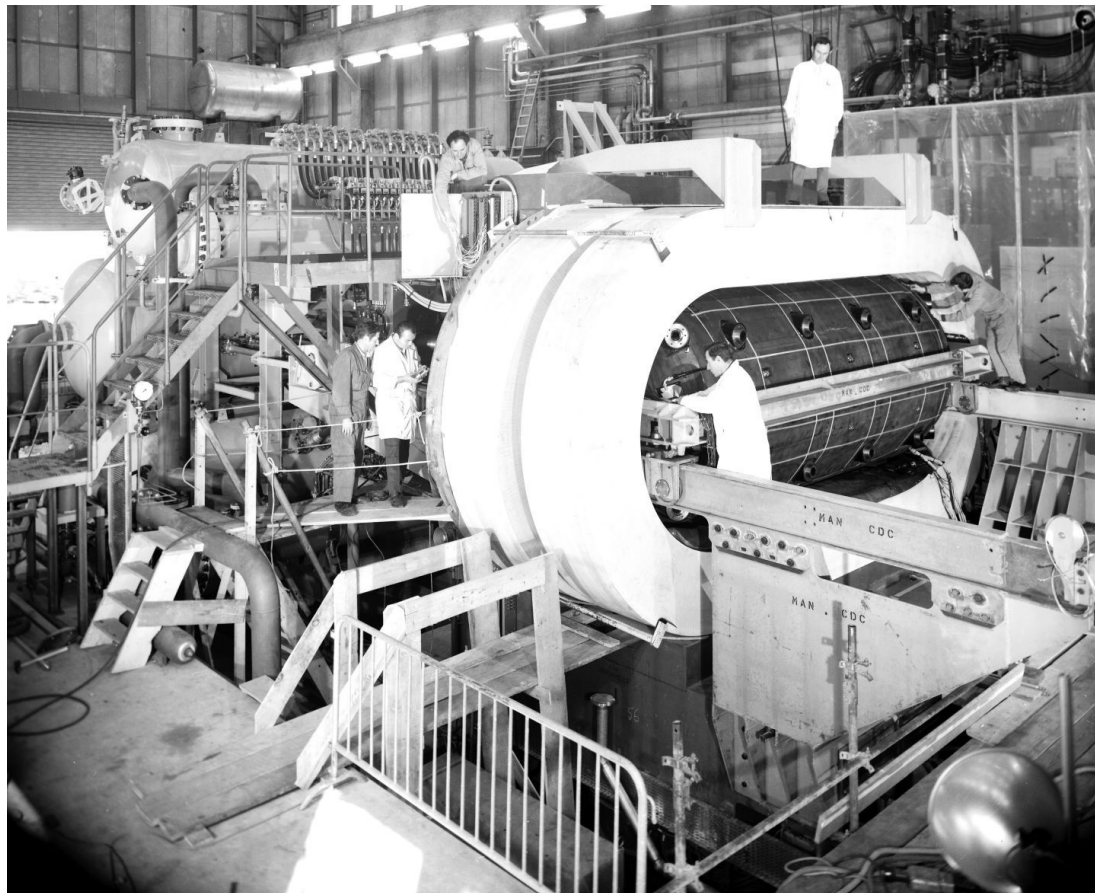
Jaké vidíme částice? :: Bublinová komora

<https://en.wikipedia.org/wiki/Neutrino>



Bublinová komora Gargamelle

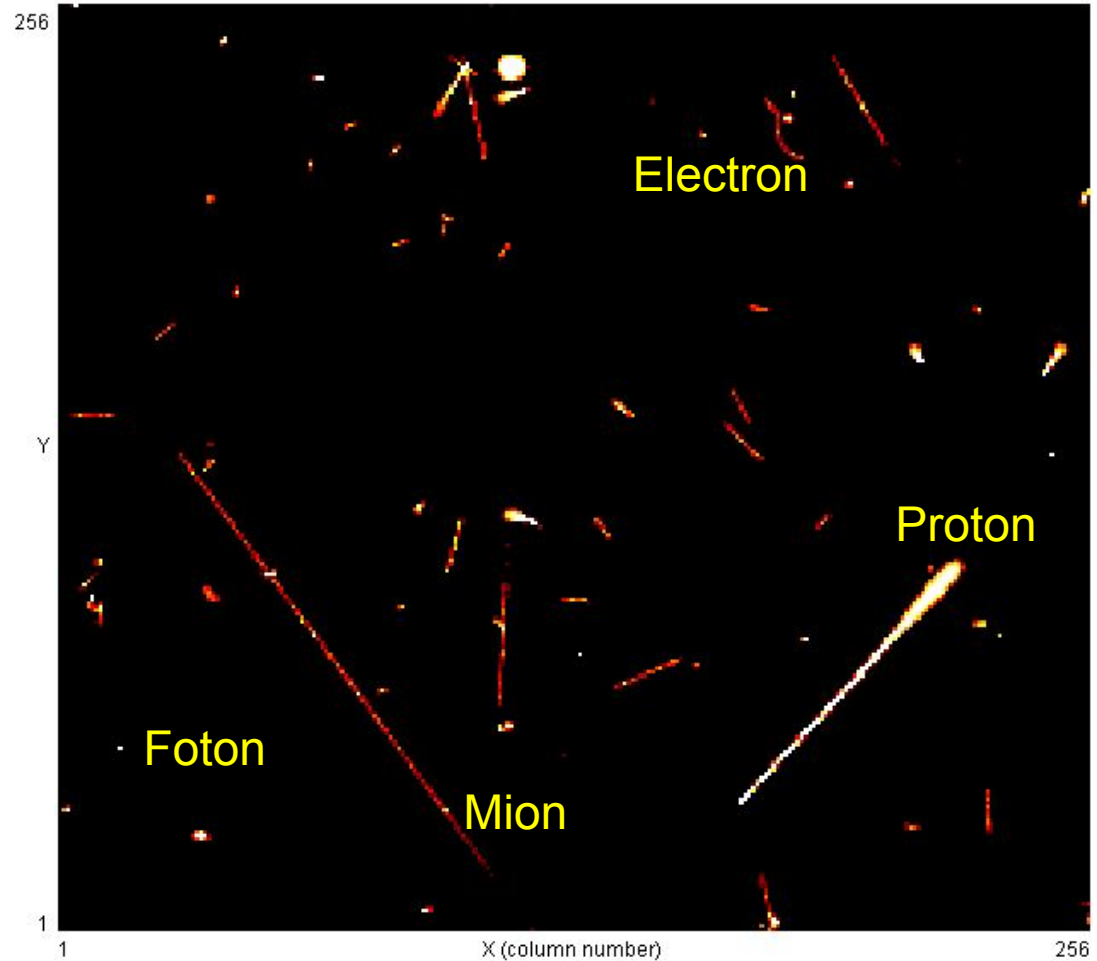
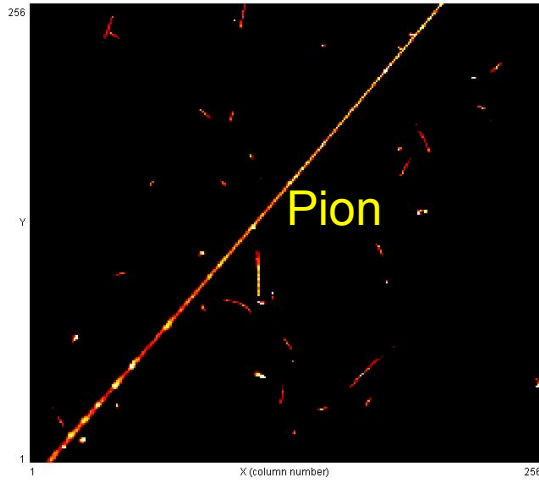
<https://cds.cern.ch/record/39648>



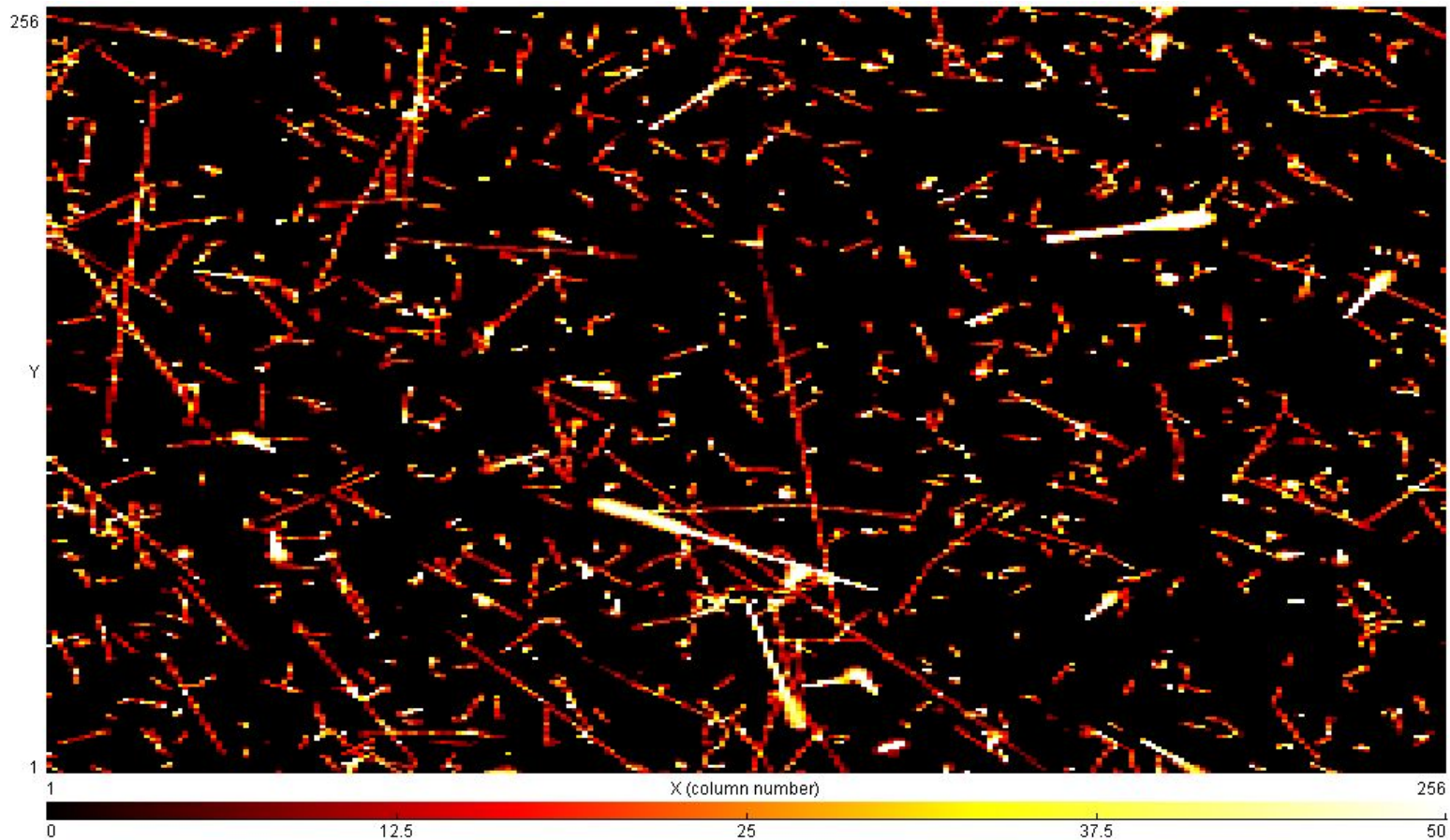
Bublinová komora Gargamelle



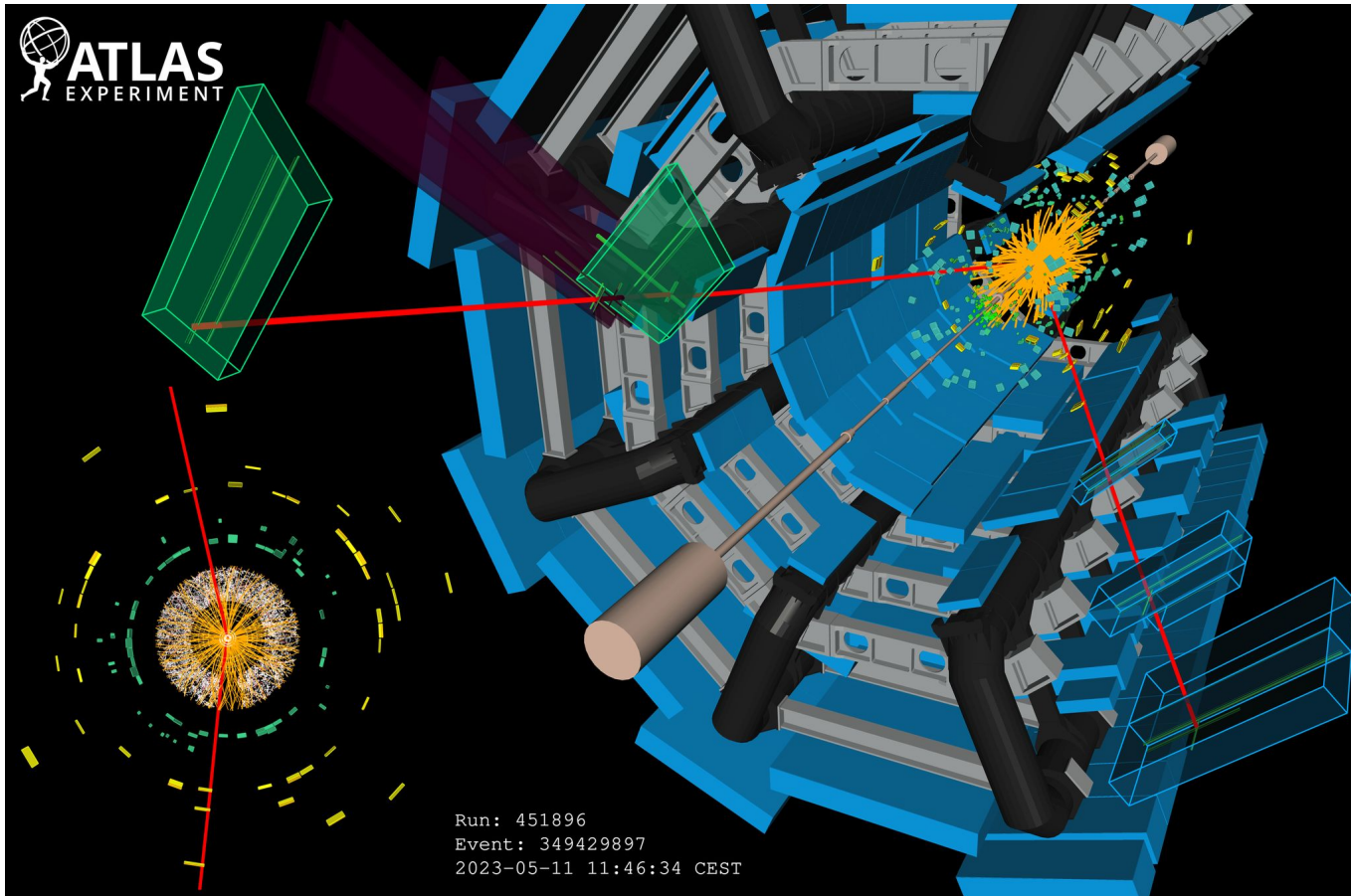
Částicová kamera



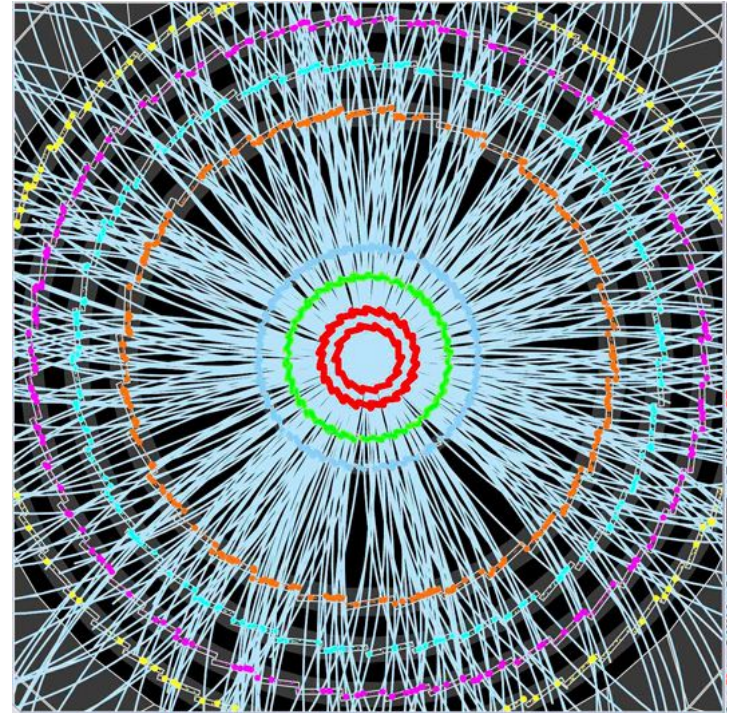
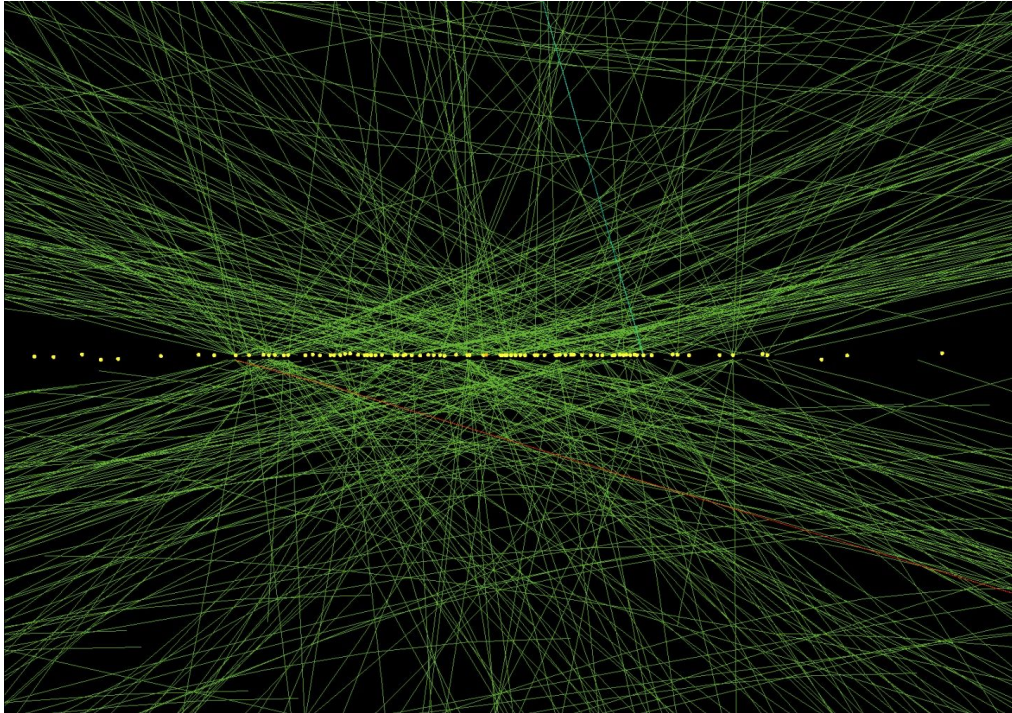
Částicová kamera :: Částice v letadle:)



ATLAS Event Display



Many collisions at the same time!

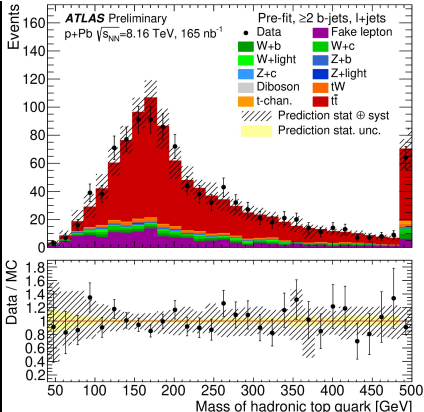
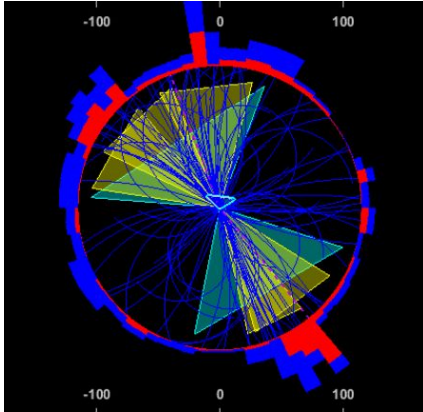


Data data data data data data data data data data...

```

414 cout << "*** Number of events in nominal and Alternative trees: " << nentries
415
416 // number of bytes read from TTrees
417 Long64_t nbytes = 0;
418 Long64_t nb_nominal = 0;
419 Long64_t nb_nominalAlt = 0;
420 Long64_t nb_ptcl = 0;
421 Long64_t nb_truth = 0;
422 m_ToRunOver = nentriesNominal;
423 if (m_isMCsignal && m_runOverAllLevels) {
424     // we need to go through all generated parton events
425     // and find matching events in the detector nominal and particle trees
426     m_ToRunOver = nentriesTruth;
427 }
428 if (m_isMCsignal && m_runOverPtclOnly) {
429     m_ToRunOver = nentriesParticle;
430 }
431 // HACK!
432 m_ToRunOver = 200000;
433
434 this -> InitLoop();
435
436 // +-----+
437 // | LOOP! |
438 // +-----+
439
440 if (m_isData) {
441     cout << "Will run over " << m_ToRunOver << " entries." << endl;
442 } else {
443     cout << "Will run over " << m_ToRunOver << " entries while sumWeights " << endl;
444     m_h_sumWeights -> SetBinContent(1, m_sumWeights);
445     m_h_sumWeightsSq -> SetBinContent(1, m_sumWeightsSq);
446 }
447
448 for (Long64_t jentry = 0; jentry < m_ToRunOver; jentry++) {
449     if (jentry % verbose == 0) {
450         cout << "Processing " << jentry << "/" << m_ToRunOver << endl;
451     }
452     nb_nominal = 0;

```



```

198 try:
199     m = ms[particle]
200 except:
201     print('Failed getting particle mass and correct the beta!')
202 if debug: print('f' m=[m:1.1f] A=[A:1.1f] I=[I:1.1f] Krel=[Krel:1.3f] C=[C:1.4f] Conv=[Conv:1.4f]')
203 if m > 0.:
204     dXScinti = dXScale*gScintiThickness
205     dXMylar = dXScale*gMylarThickness
206     # evaluate the fitted losses in T50
207     dE = -999.
208     if 'TS1' in region:
209         # beam window AND the T50 passage;
210         # assume losses will be similar, after all
211         # let's avoid fitting I in plastics and Mylar as I has a high unc. anyway
212         dE = (dXScinti + dXMylar) * Conv * getBaseMeanLossesFitVal(beta, A, I, C, debug)
213         #dE = dXScinti * Conv * getBaseMeanLossesFitVal(beta, A, I, C, debug)
214     else:
215         # just the beam window;-)
216         dE = dXMylar * Conv * getBaseMeanLossesFitVal(beta, A, I, C, debug)
217         pass
218
219 if dE > 0 and beta > 0. and beta < 1.:
220     gamma0 = 1./sqrt(1. - pow(beta,2))
221     E0 = m*gamma0
222     p0 = sqrt(E0*E0 - m*m)
223     gamma1 = (E0 - dE) / m
224     fitFracLoss = dE / E0
225
226 useHigherCorrs = False
227 # ADD ALSO BEAM WINDOW FOR BOTH T50 AND TS1!
228 fullName = 'e'
229 if particle == 'p':
230     fullName = 'proton'
231 if particle == 'D':

```

Princip objevu nových částic: invariantní hmota

$$E = mc^2$$

$$E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$E^2 = p^2 c^2 + m^2 c^4$$

$$m^2 = E^2 - p^2 \quad \text{v jednotkách } c=1$$

$$m^2 = \left(\sum_i E_i \right)^2 - \left(\sum_i \vec{p}_i \right)^2$$

$$E_i^2 = p_{ix}^2 + p_{iy}^2 + p_{iz}^2$$

$$E = \gamma mc^2$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \quad \beta \equiv \frac{v}{c}$$

$$E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$E = E_1 + E_2$$

$$\vec{p} = \vec{p}_1 + \vec{p}_2$$

$$\vec{p}^2 \equiv \vec{p} \cdot \vec{p} = \vec{p}_1^2 + 2\vec{p}_1 \cdot \vec{p}_2 + \vec{p}_2^2$$

$$m_1 \approx m_2 \approx 0 \Rightarrow$$

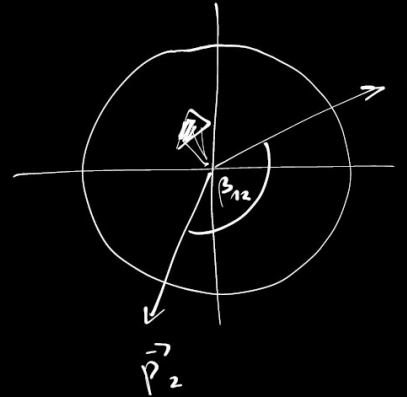
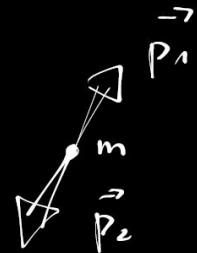
$$\doteq E_1^2 + 2\vec{p}_1 \cdot \vec{p}_2 + E_2^2$$

$$= E_1^2 + E_2^2 + 2E_1 E_2 \cos \beta_{12}$$

$$m^2 = (E_1 + E_2)^2 - (\vec{p}_1 + \vec{p}_2)^2$$

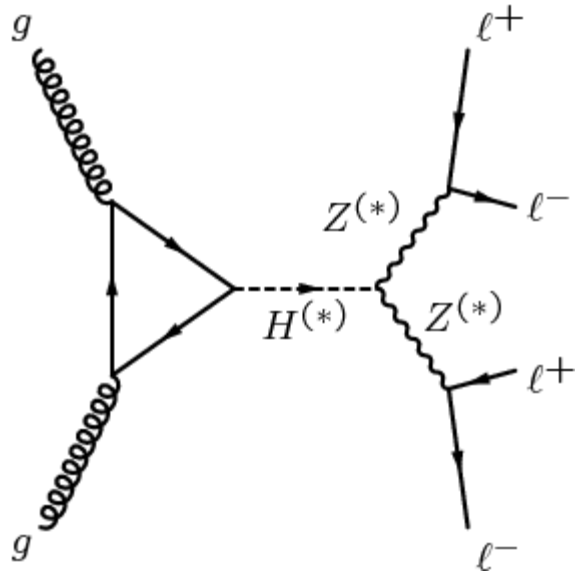
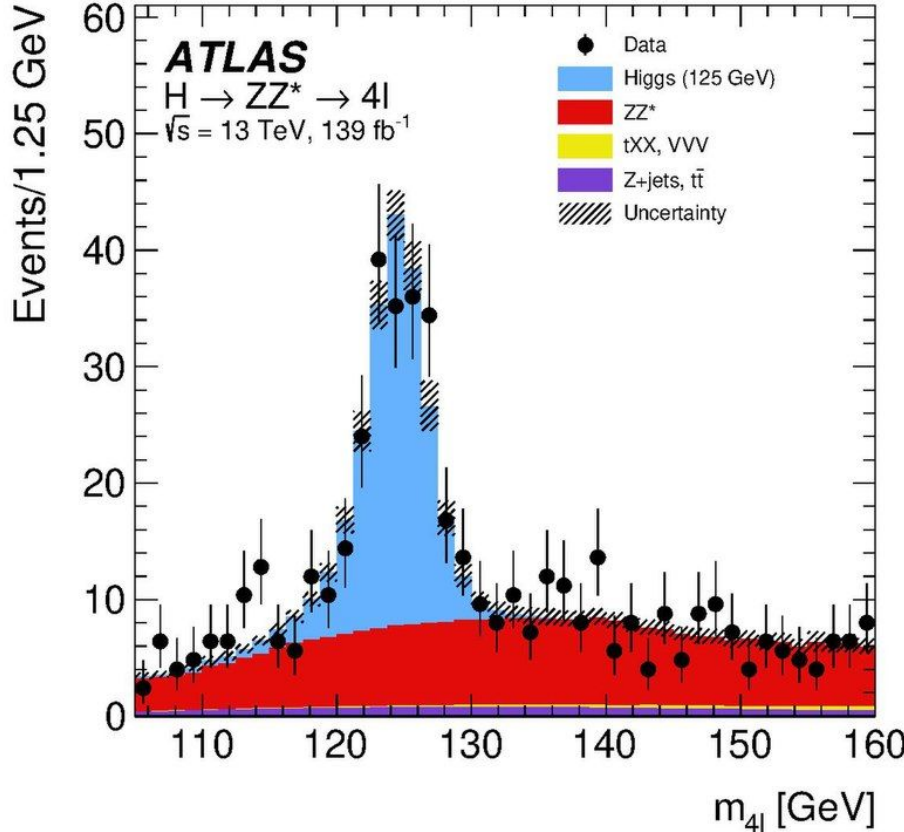
$$= \cancel{E_1^2} + 2E_1 E_2 + \cancel{E_2^2}$$

$$- \cancel{E_1^2} - \cancel{E_2^2} - 2E_1 E_2 \cos \beta_{12}$$



$$m^2 \approx 2E_1 E_2 (1 - \cos \beta_{12})$$

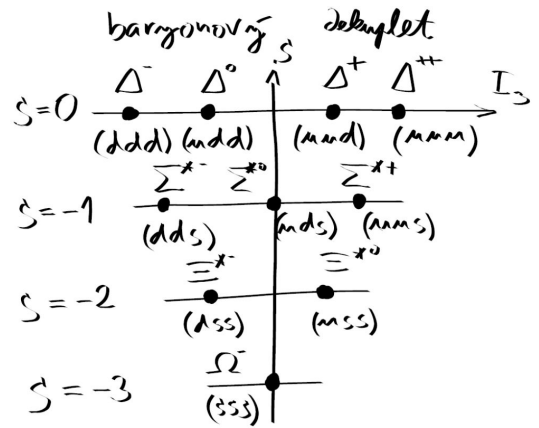
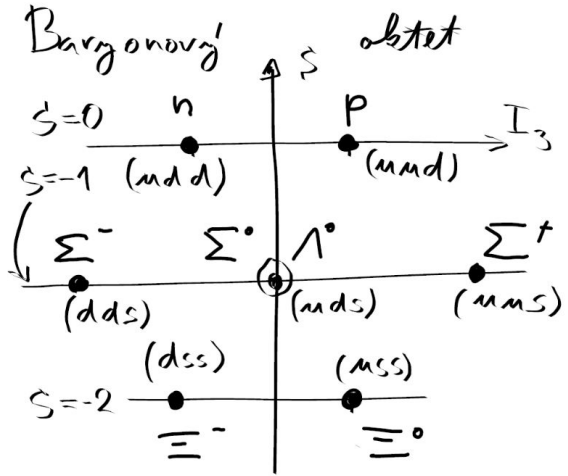
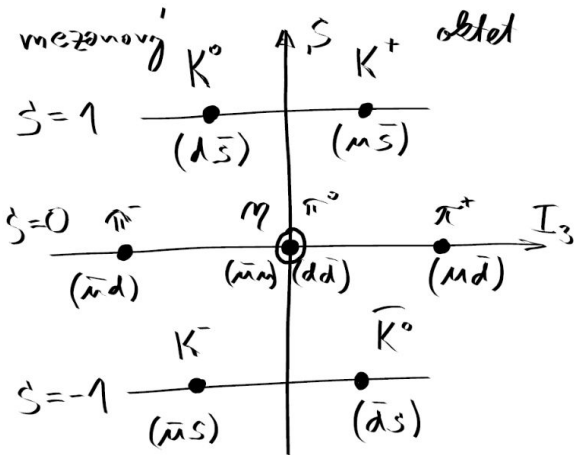
Takhle “vidíme” nestabilní částice – peak!



**Listopadová revoluce
...ve fyzice částic
1974**

Listopad 1974

- Tehdy známé kvarky, tj. stavební kameny hmoty: u, d, s
- Podivné částice: obsahují s-kvark (strange)
- Rodily se v párech, rozpadaly se "pomalu" a různě, bylo jich mnoho, ale podařilo se je systematizovat



Listopad 1974

- Existovaly teoretické důvody pro další, 4. kvark.
- Přesto byl jeho objev překvapivý, dobrodružný, spektakulární!
- Objeven současně ve dvou zcela odlišných experimentech!

Objev 4. kvarku a částice J/Psi :: Listopad 1974

- Skupina na urychlovači protonů Brookhaven National Laboratory.
- Urychlovač Alternating Gradient Synchrotron, AGS, 30 GeV protony.
- Srážky protonů s beryliovým terčem.
- Studium produkce párů elektron-pozitron.

VOLUME 33, NUMBER 23

PHYSICAL REVIEW LETTERS

2 DECEMBER 1974

Experimental Observation of a Heavy Particle J^{ψ}

J. J. Aubert, U. Becker, P. J. Biggs, J. Burger, M. Chen, G. Everhart, P. Goldhagen, J. Leong, T. McCorriston, T. G. Rhoades, M. Rohde, Samuel C. C. Ting, and Sau Lan Wu
Laboratory for Nuclear Science and Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139

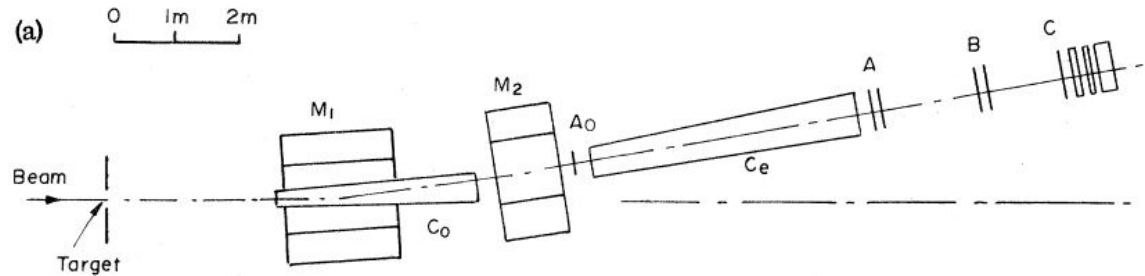
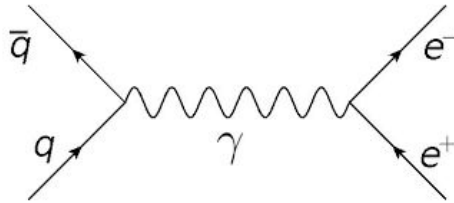
and

Y. Y. Lee

Brookhaven National Laboratory, Upton, New York 11973

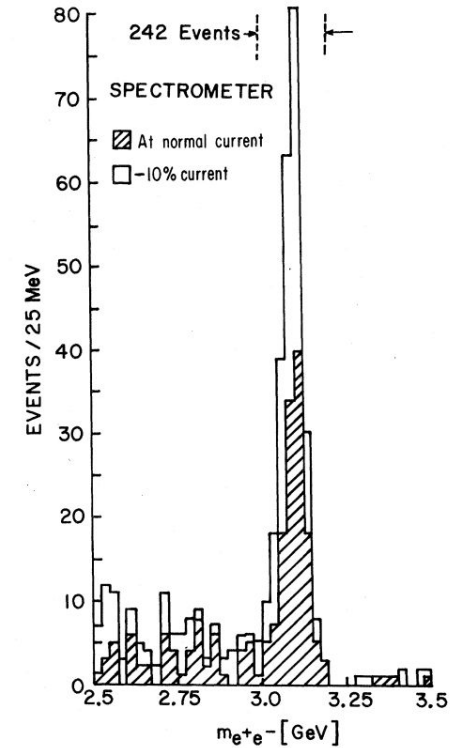
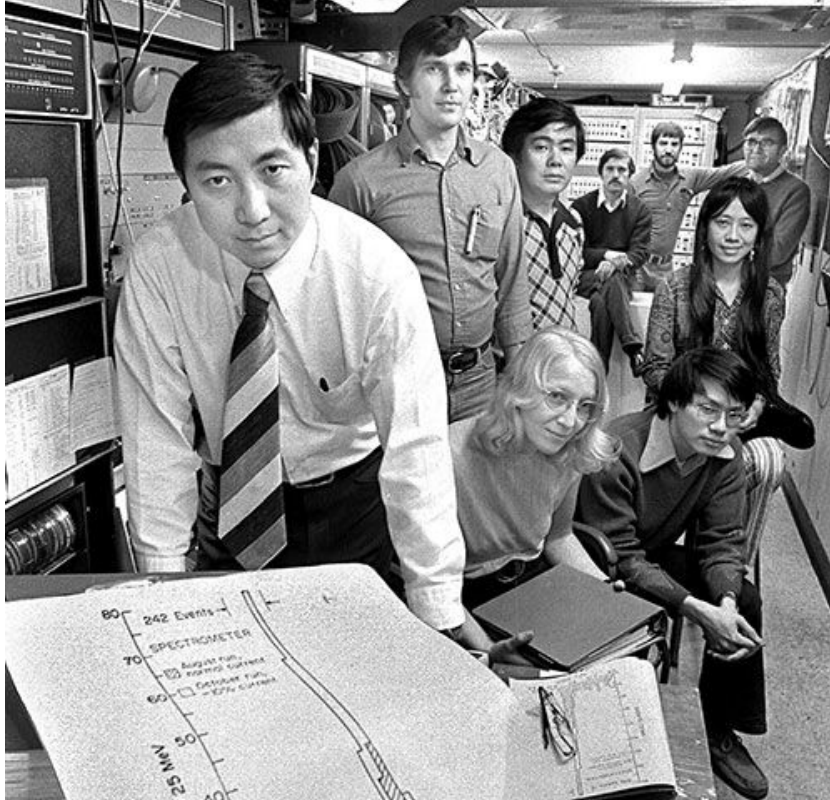
(Received 12 November 1974)

We report the observation of a heavy particle J , with mass $m = 3.1$ GeV and width approximately zero. The observation was made from the reaction $p + \text{Be} \rightarrow e^+ + e^- + x$ by measuring the e^+e^- mass spectrum with a precise pair spectrometer at the Brookhaven National Laboratory's 30-GeV alternating-gradient synchrotron.



Objev 4. kvarku a částice J/Psi :: Listopad 1974

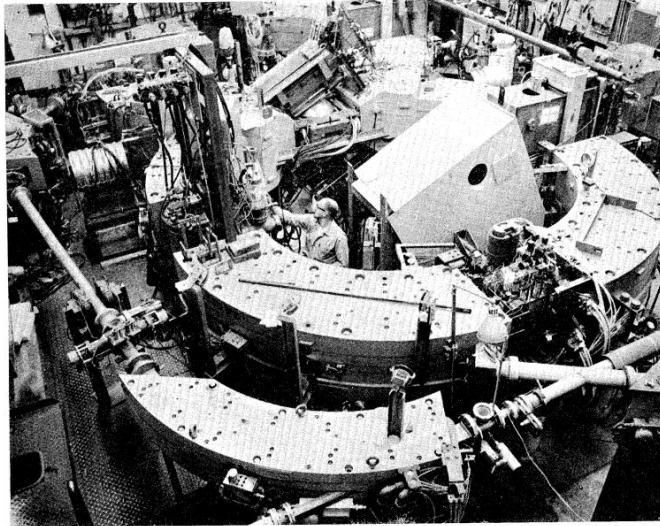
- Překvapivý peak v invariantní hmotě!



<https://www.bnl.gov/bnlweb/history/nobel/1976.php>

Objev 4. kvarku a částice J/Psi :: Listopad 1974

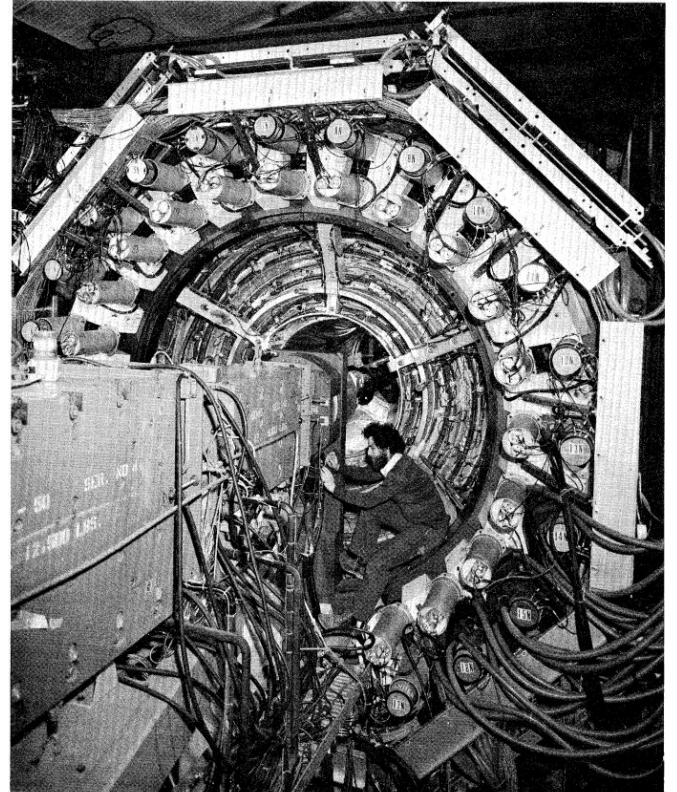
- Skupina v laboratoři SLAC, Stanford.
- Urychlovač elektronů a pozitronů SPEARS
- Detektor Mark I



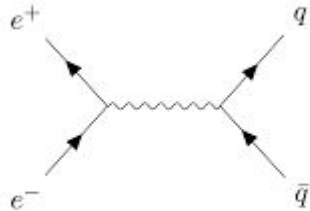
This photo shows the figure-8-shaped 500 MeV electron-electron storage rings that were built as a collaborative Stanford-Princeton project at the High Energy Physics Lab at Stanford. The orientation of this double-ring machine in the photo is about like this: ϕ . The odd-shaped magnet in the foreground is a part of the beam-injection system. This was the first colliding-beam machine to carry out a successful program of physics experiments, its first results being obtained in 1965. Burt Richter was one of four principal collaborators in the construction and research use of these rings. (Stanford News & Publications photo.)

SLAC Beam Line, November 1976

S-7

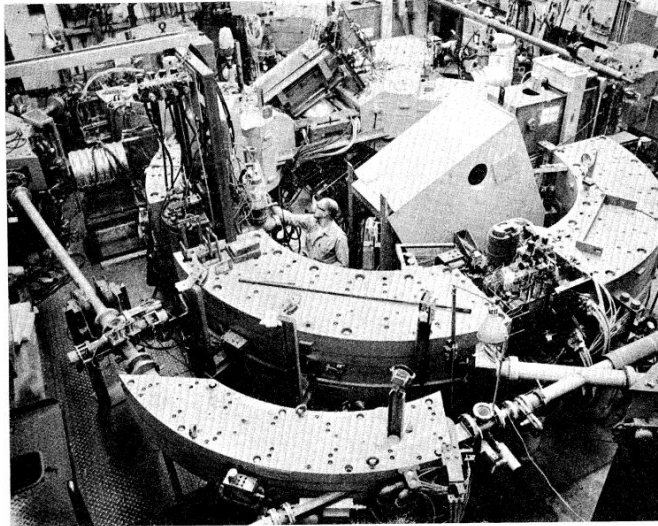


This photograph gives a better sense of the size and complexity of the SPEAR Mark I magnetic detector than the previous words. Former LBL experimentalist Carl Friedberg is shown working on the device with its end caps removed. The Mark I is connected to the outside world by water



Objev 4. kvarku a částice J/Psi :: Listopad 1974

- Skupina v laboratoři SLAC, Stanford.
- Urychlovač elektronů a pozitronů SPEARS
- Detektor Mark I
- Burton Richter



This photo shows the figure-8-shaped 500 MeV electron-electron storage rings that were built as a collaborative Stanford-Princeton project at the High Energy Physics Lab at Stanford. The orientation of this double-ring machine in the photo is about like this: \oslash . The odd-shaped magnet in the foreground is a part of the beam-injection system. This was the first colliding-beam machine to carry out a successful program of physics experiments, its first results being obtained in 1965. Burt Richter was one of four principal collaborators in the construction and research use of these rings. (Stanford News & Publications photo.)

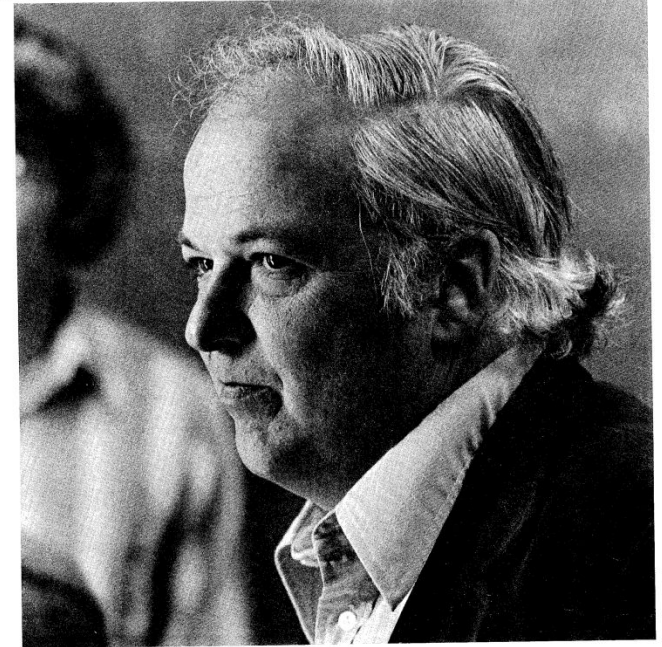


SLAC BEAM LINE

"There are therefore Agents in Nature able to make the Particles of Bodies stick together by very strong Attractions. And it is the Business of experimental Philosophy to find them out." - Isaac Newton Opticks (1704)

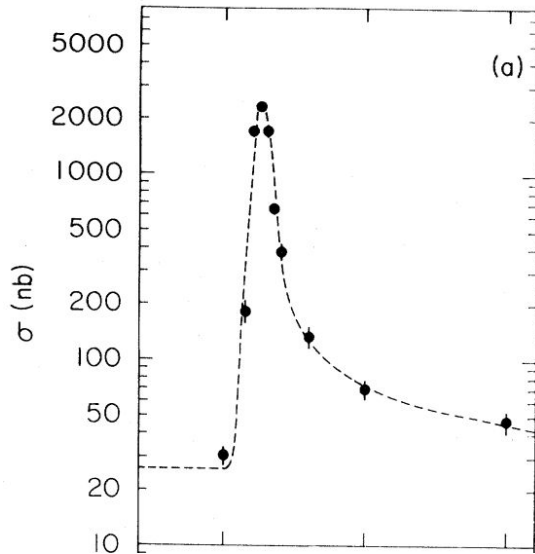
Volume 7, Number 11

November 1976



Objev 4. kvarku a částice J/Psi :: Listopad 1974

- Překvapivé navýšení počtu pozorovaných částic v úzké oblasti energií urychlovače!



Discovery of a Narrow Resonance in e^+e^- Annihilation*

J.-E. Augustin,† A. M. Boyarski, M. Breidenbach, F. Bulos, J. T. Dakin, G. J. Feldman,
G. E. Fischer, D. Fryberger, G. Hanson, B. Jean-Marie,† R. R. Larsen, V. Lüth,
H. L. Lynch, D. Lyon, C. C. Morehouse, J. M. Paterson, M. L. Perl,
B. Richter, P. Rapidis, R. F. Schwitters, W. M. Tanenbaum,
and F. Vannucci‡

Stanford Linear Accelerator Center, Stanford University, Stanford, California 94305

and

G. S. Abrams, D. Briggs, W. Chinowsky, C. E. Friedberg, G. Goldhaber, R. J. Hollebeek,
J. A. Kadyk, B. Lulu, F. Pierre,§ G. H. Trilling, J. S. Whitaker,
J. Wiss, and J. E. Zipse

Lawrence Berkeley Laboratory and Department of Physics, University of California, Berkeley, California 94720

(Received 13 November 1974)

We have observed a very sharp peak in the cross section for $e^+e^- \rightarrow \text{hadrons}$, e^+e^- , and possibly $\mu^+\mu^-$ at a center-of-mass energy of 3.105 ± 0.003 GeV. The upper limit to the full width at half-maximum is 1.3 MeV.

Objev 4. kvarku a částice J/Psi :: Meanwhile in Frascati, IT...

VOLUME 33, NUMBER 23

PHYSICAL REV

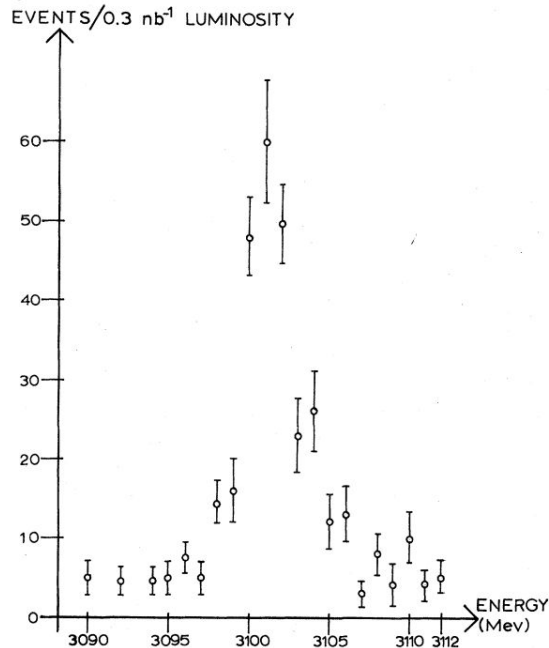


FIG. 1. Result from the Gamma-Gamma Group, total of 446 events. The number of events per 0.3 nb^{-1} luminosity is plotted versus the total c.m. energy of the machine.

Preliminary Result of Frascati (ADONE) on the Nature of a New 3.1-GeV Particle Produced in e^+e^- Annihilation*

C. Bacci, R. Balbini Celio, M. Berna-Rodini, G. Caton, R. Del Fabbro, M. Grilli, E. Iarocci, M. Locci, C. Mencuccini, G. P. Murtas, G. Penso, G. S. M. Spinetti, M. Spano, B. Stella, and V. Valente

The Gamma-Gamma Group, Laboratori Nazionali di Frascati, Frascati, Italy

and

B. Bartoli, D. Bisello, B. Esposito, F. Felicetti, P. Monacelli, M. Nigro, L. Paolufi, I. Peruzzi, G. Piano Mortemi, M. Piccolo, F. Ronga, F. Sebastiani, L. Trasatti, and F. Vanoli
The Magnet Experimental Group for ADONE, Laboratori Nazionali di Frascati, Frascati, Italy

and

G. Barbarino, G. Barbiellini, C. Bemporad, R. Biancastelli, F. Cevenini, M. Celvetti, F. Costantini, P. Lariccia, P. Parascandolo, E. Sassi, C. Spencer, L. Tortora, U. Troya, and S. Vitale

The Baryon-Antibaryon Group, Laboratori Nazionali di Frascati, Frascati, Italy

(Received 18 November 1974)

We report on the results at ADONE to study the properties of the newly found 3.1-BeV particle.

Objev 4. kvarku a částice J/Psi :: NP

The Nobel Prize
in Physics 1976

Summary **The Nobel Prize in Physics** Laureates **1976**

Burton Richter

Samuel C.C. Ting

Press release

Award ceremony
speech

Share this



Photo from the Nobel
Foundation archive.

Burton Richter

Prize share: 1/2



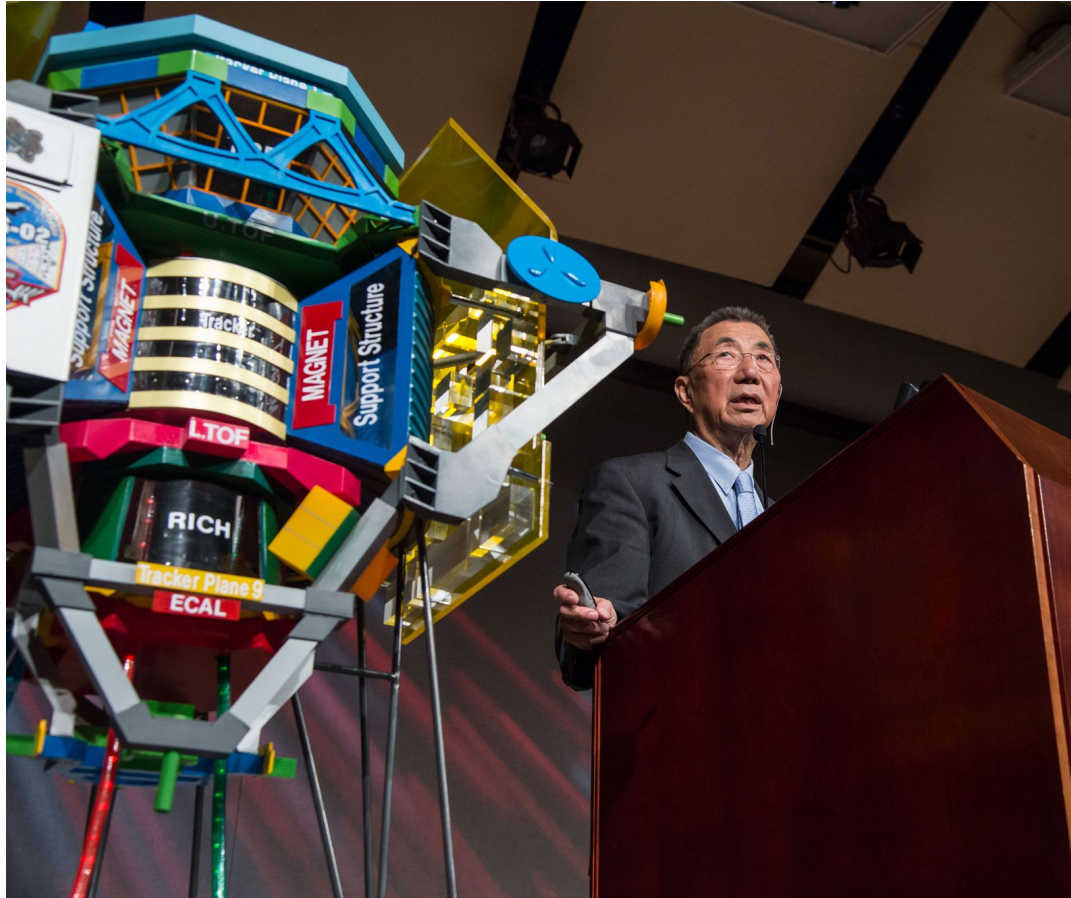
Photo from the Nobel
Foundation archive.

**Samuel Chao
Chung Ting**

Prize share: 1/2

The Nobel Prize in Physics 1976 was awarded jointly to Burton Richter and Samuel Chao Chung Ting "for their pioneering work in the discovery of a heavy elementary particle of a new kind"

Objev 4. kvarku a částice J/Psi :: Listopad 1974



Objev 4. kvarku a částice J/Psi :: Listopad 1974



<https://www6.slac.stanford.edu/news/2014-11-20-president-obama-bestows-national-medal-science-slac-director-emeritus-and-nobelist>

Objev 4. kvarku a částice J/Psi :: Listopad 1974

Meeting prof. Ting with our students, March 2022.



Burton Richter (1931–2018)

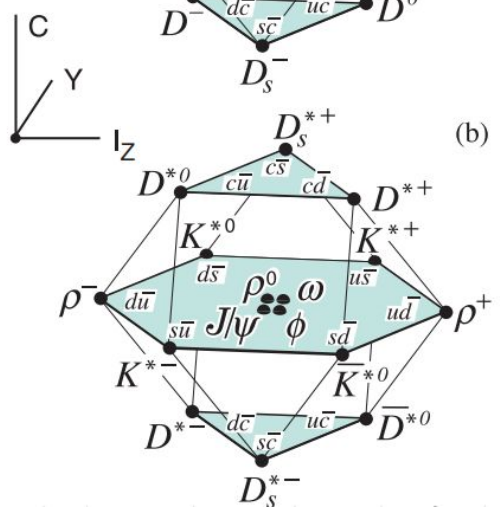
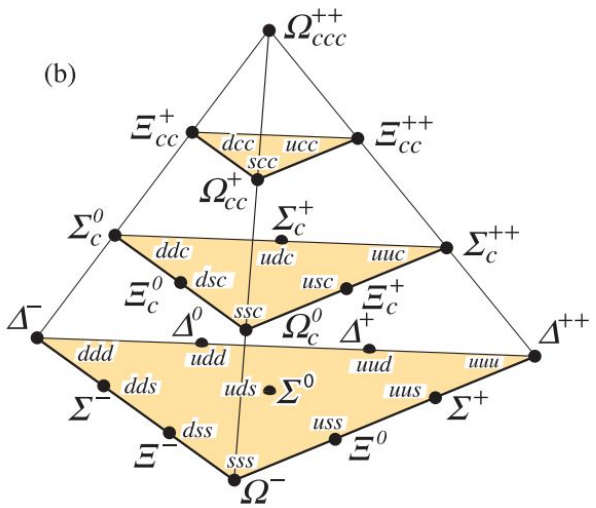
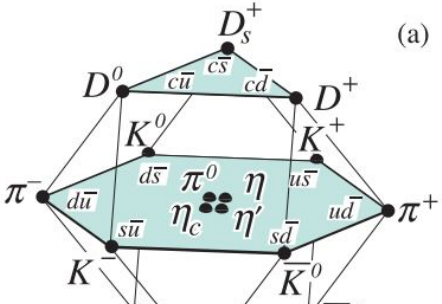
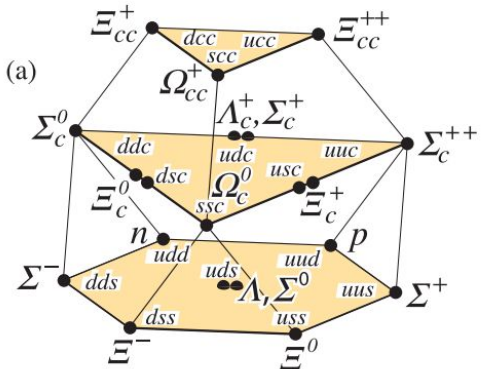
Physicist who helped to discover the first particle containing a charm quark.

By [Helen Quinn](#)

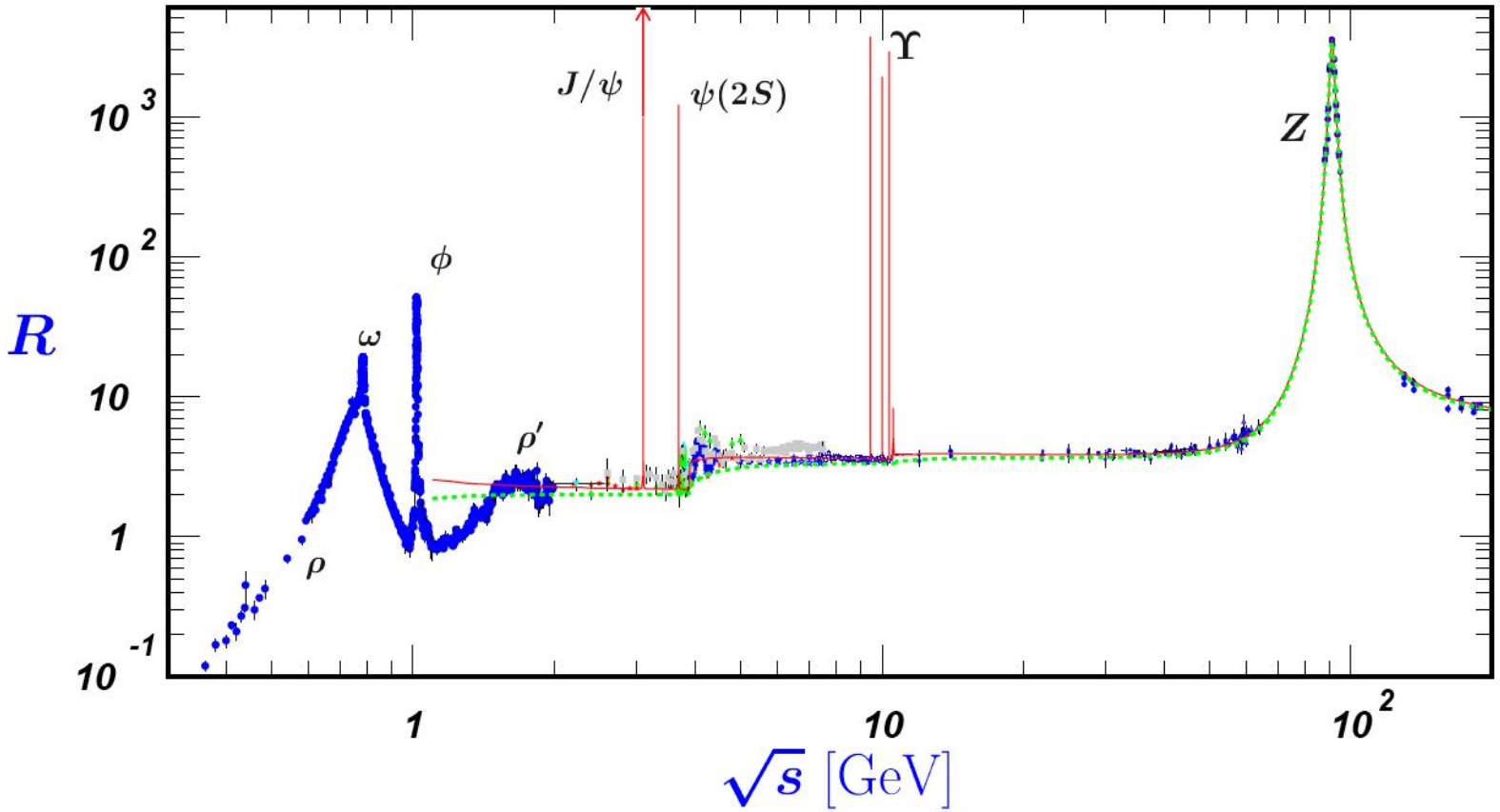


Credit: Eddie Adams/AP/REX/Shutterstock

Skládání kvarků u, d, s, c



Krajina peaků a částic:)



The Three Body Problem :: Problém tří těles

Novel by Cixin Liu

Movie by China

Movie by Netflix



The Three-Body Problem (film)

5 languages

Article Talk Read Edit View history Tools

From Wikipedia, the free encyclopedia

The Three-Body Problem (Chinese: 三体) is an unreleased Chinese science fiction 3D film,^[2] adapted from the book of the same name by Liu Cixin, directed by Zhang Fanfan, and starring Feng Shaofeng and Zhang Jingchu.^{[3][4][5]}

In March 2018, Amazon was rumored to be negotiating for the rights to the project.^{[6][7]} However, Youzu Pictures released a statement in response stating that it was the "sole owner of the rights for film and TV series adaptations." Although it "was originally scheduled to be released in 2017", the project "was postponed indefinitely due to the company's internal shuffling and the rumored 'bad quality' of the film's first cut."^[7]

Production

The Three-Body Problem



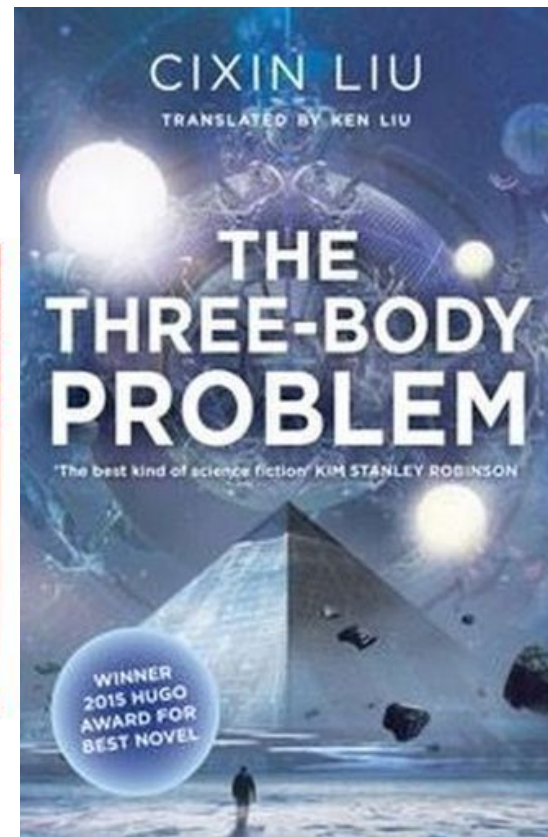
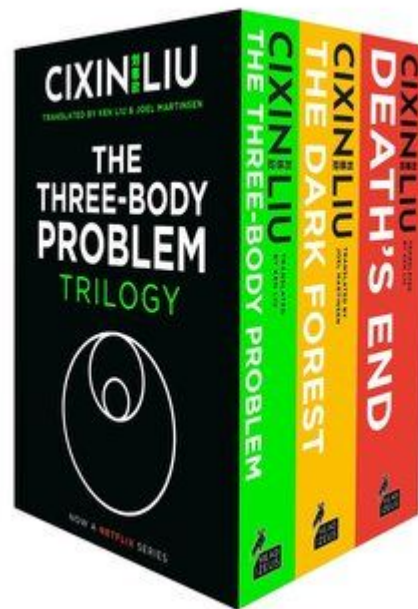
Poster

Traditional Chinese 三體

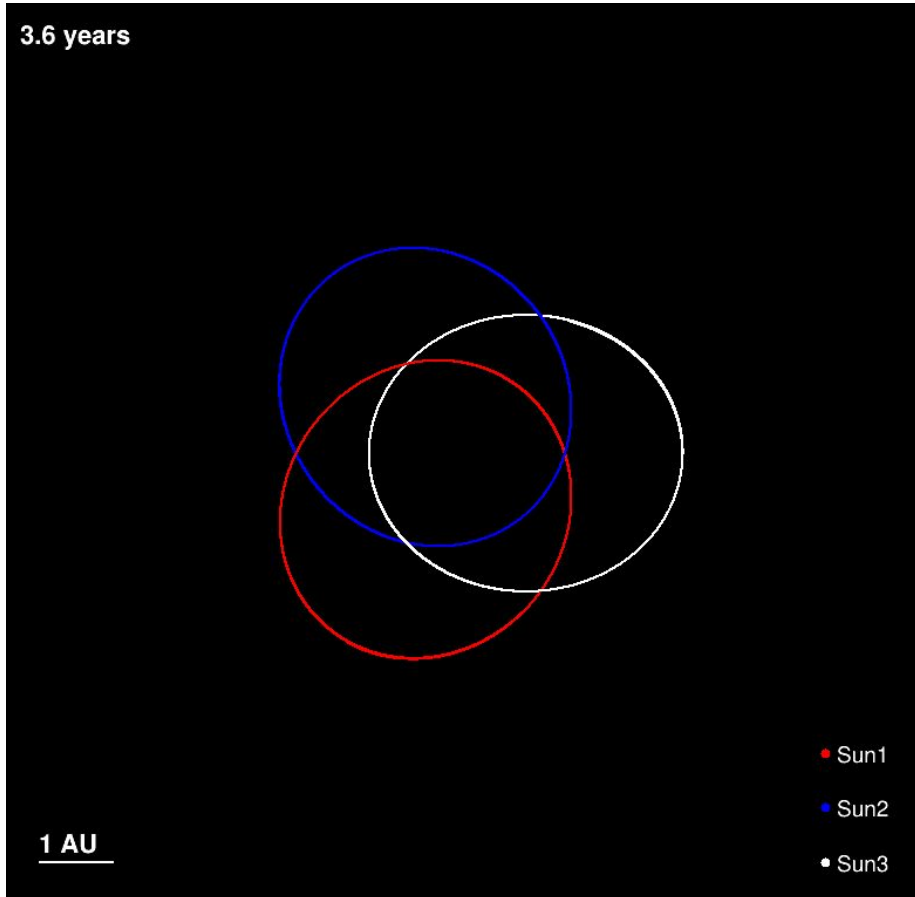
Simplified Chinese 三体

Literal meaning three bodies

Hanyu Pinyin sān tǐ



The Three Body Problem :: Problém tří těles

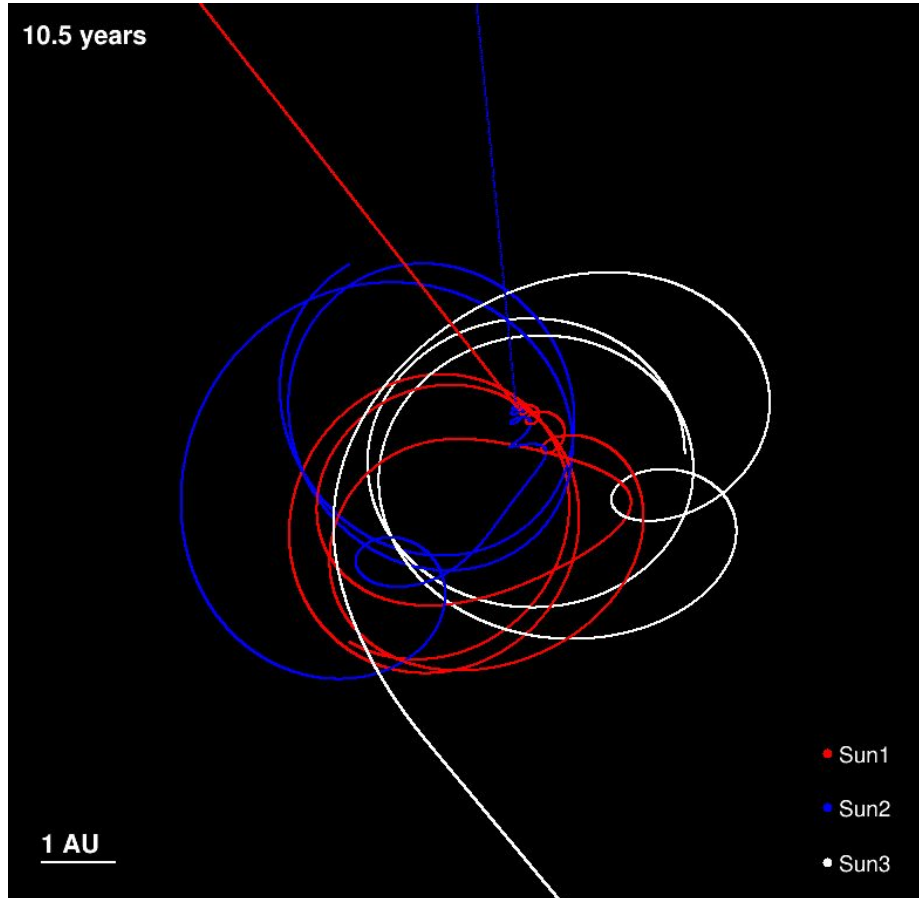


Same solar masses, quasi-stability.

More:

https://youtu.be/mm3l4m8YsnM?si=Ajz0tH_7xfPyW74o

The Three Body Problem :: Problém tří těles

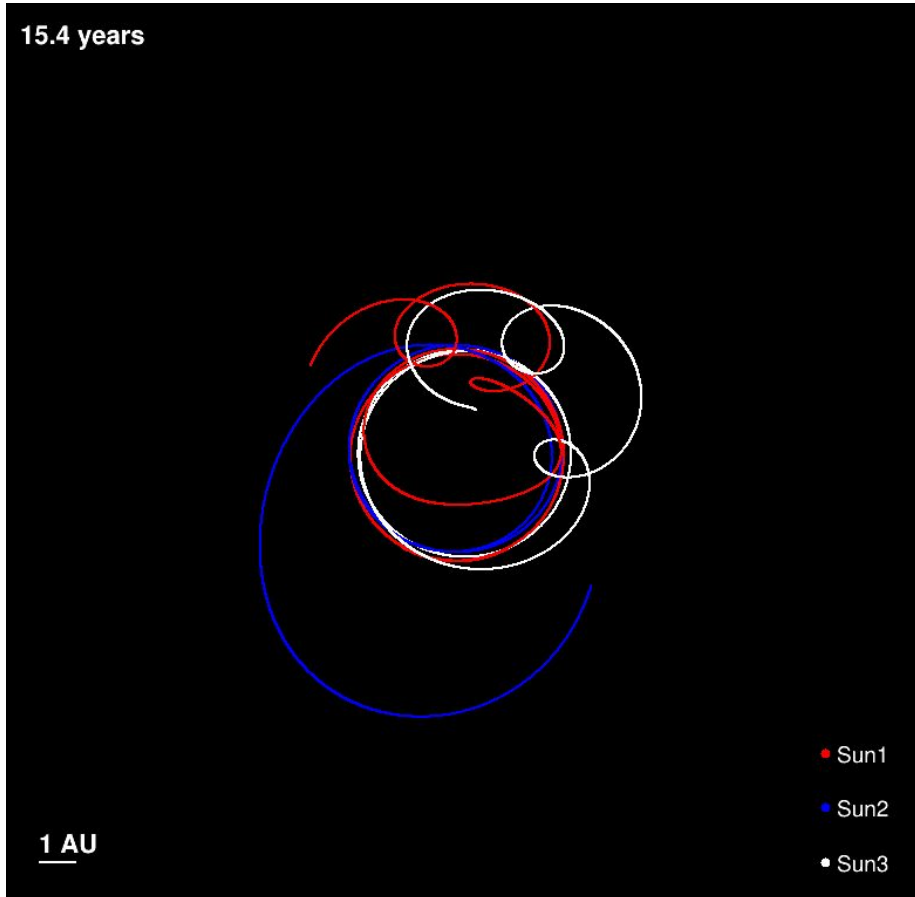


5% change in Solar masses

More:

https://youtu.be/mm3l4m8YsnM?si=Ajz0tH_7xfPyW74o

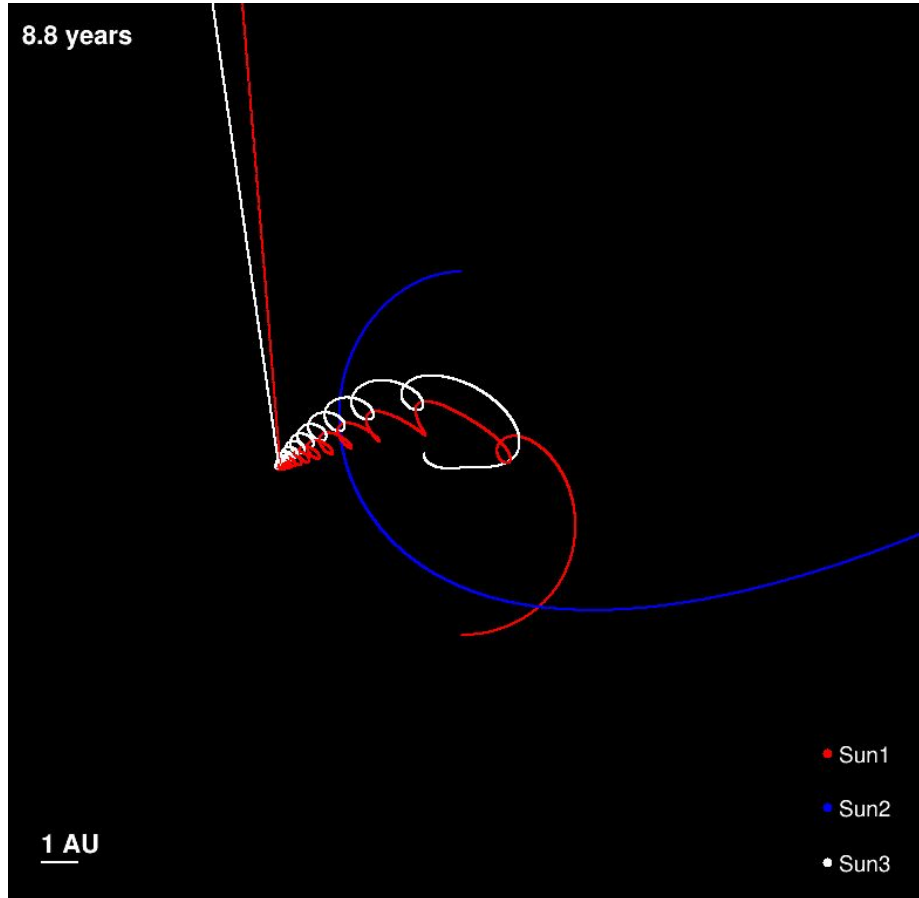
The Three Body Problem :: Problém tří těles



More:

https://youtu.be/mm3I4m8YsnM?si=Ajz0tH_7xfPyW74o

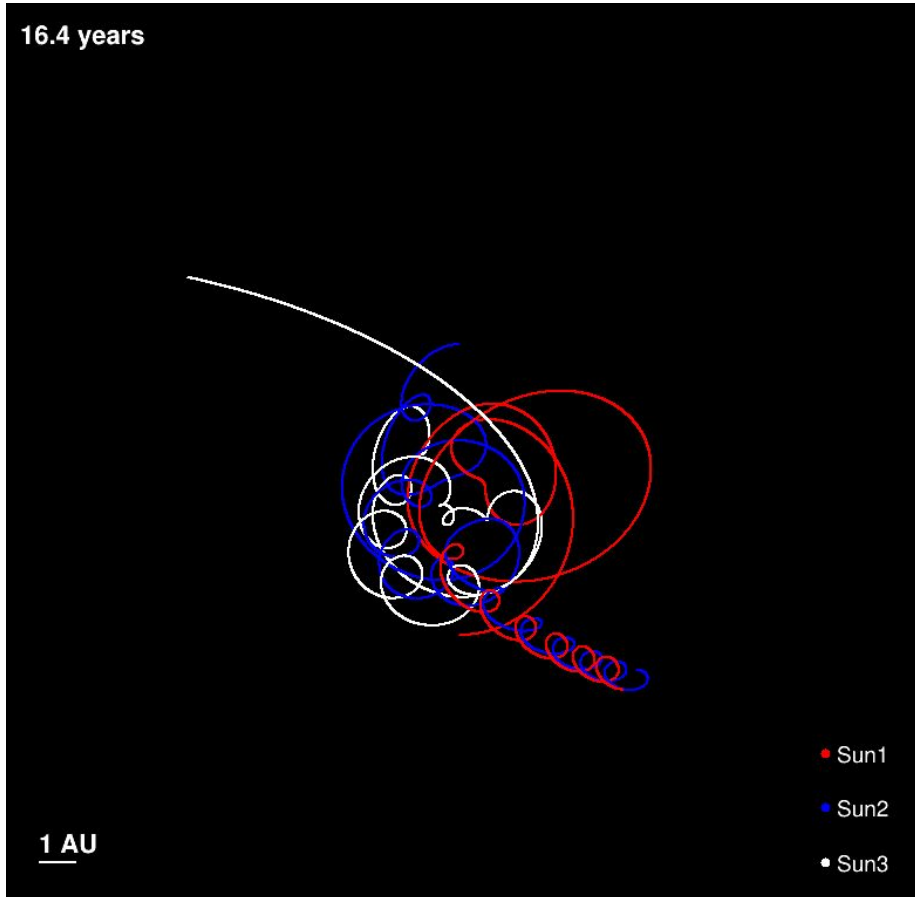
The Three Body Problem :: Problém tří těles



More:

https://youtu.be/mm3I4m8YsnM?si=Ajz0tH_7xfPyW74o

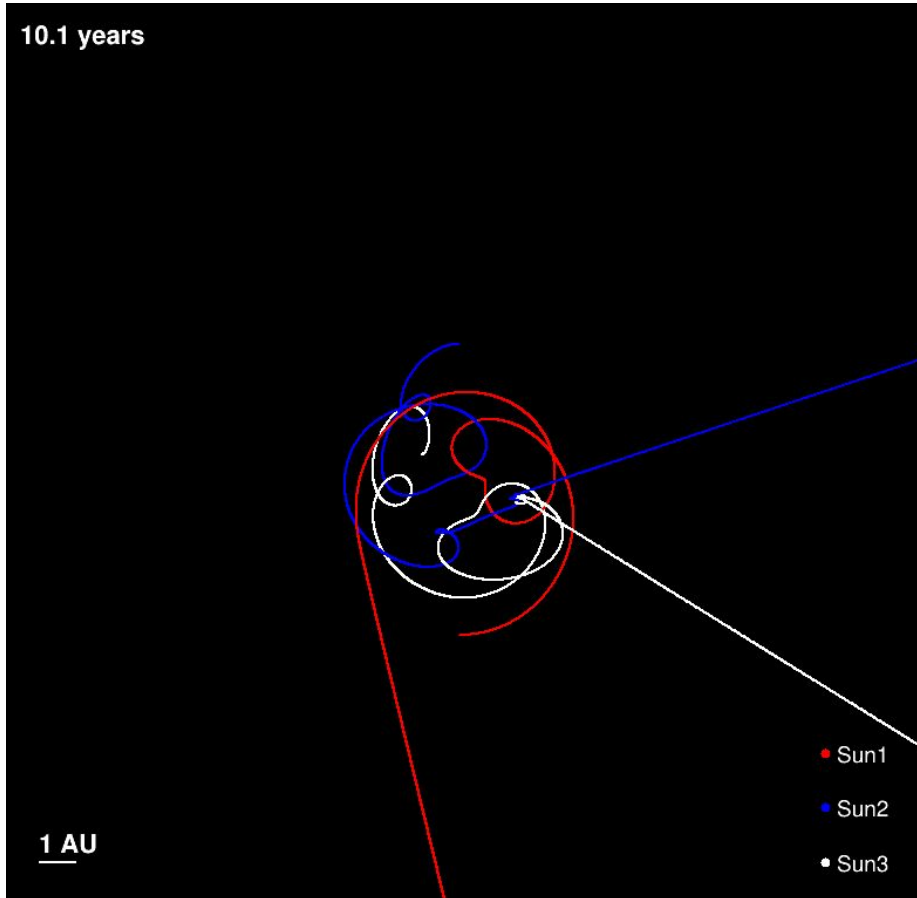
The Three Body Problem :: Problém tří těles



More:

https://youtu.be/mm3I4m8YsnM?si=Ajz0tH_7xfPyW74o

The Three Body Problem :: Problém tří těles



1% change in initial position of Sun2

More:

https://youtu.be/mm3l4m8YsnM?si=Ajz0tH_7xfPyW74o

The Three Body Problem :: “Oxford”

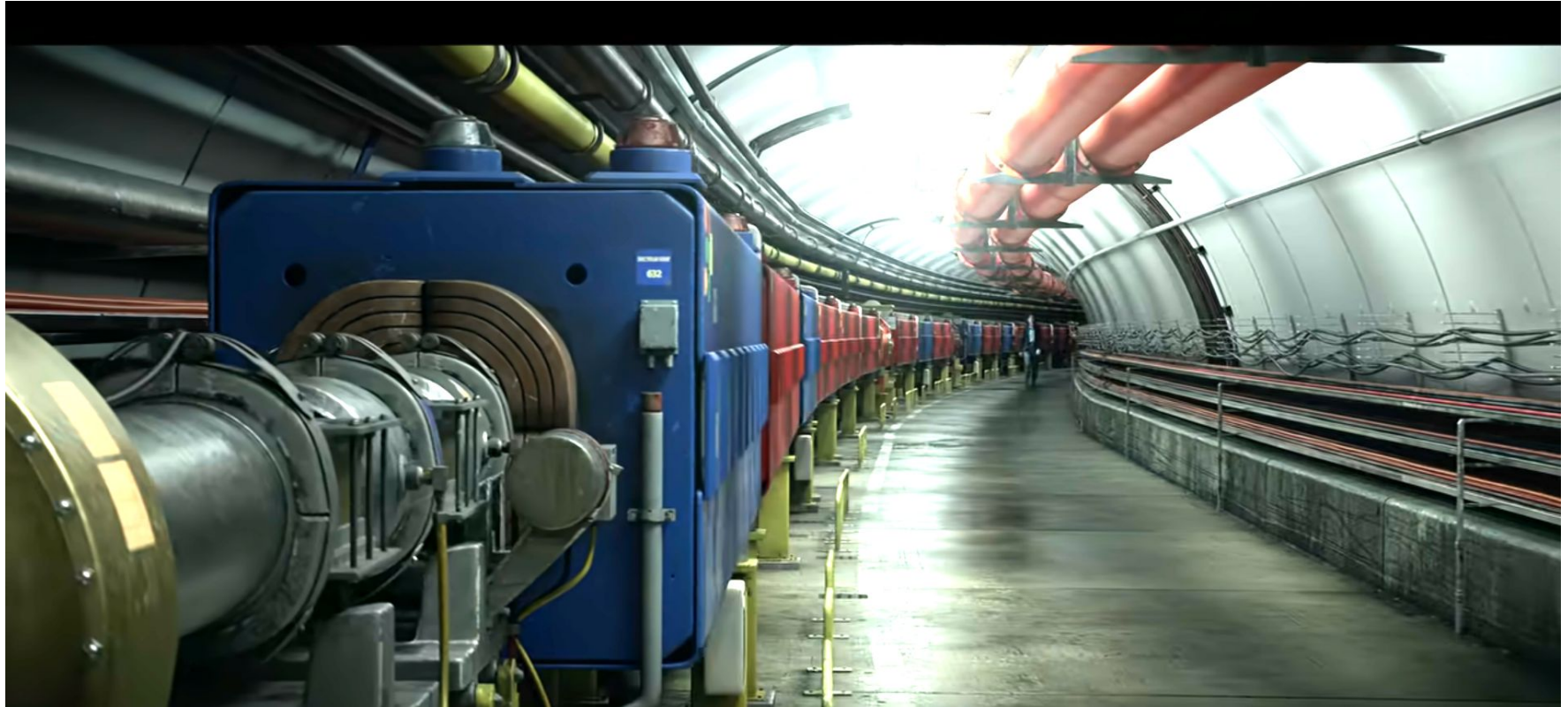
URYCHLOVAČ ČÁSTIC NA OXFORDSKÉ UNIVERZITĚ

OXFORD UNIVERSITY PARTICLE ACCELERATOR

The Three Body Problem :: Brookhaven National Light Source, USA



The Three Body Problem :: “Oxford”

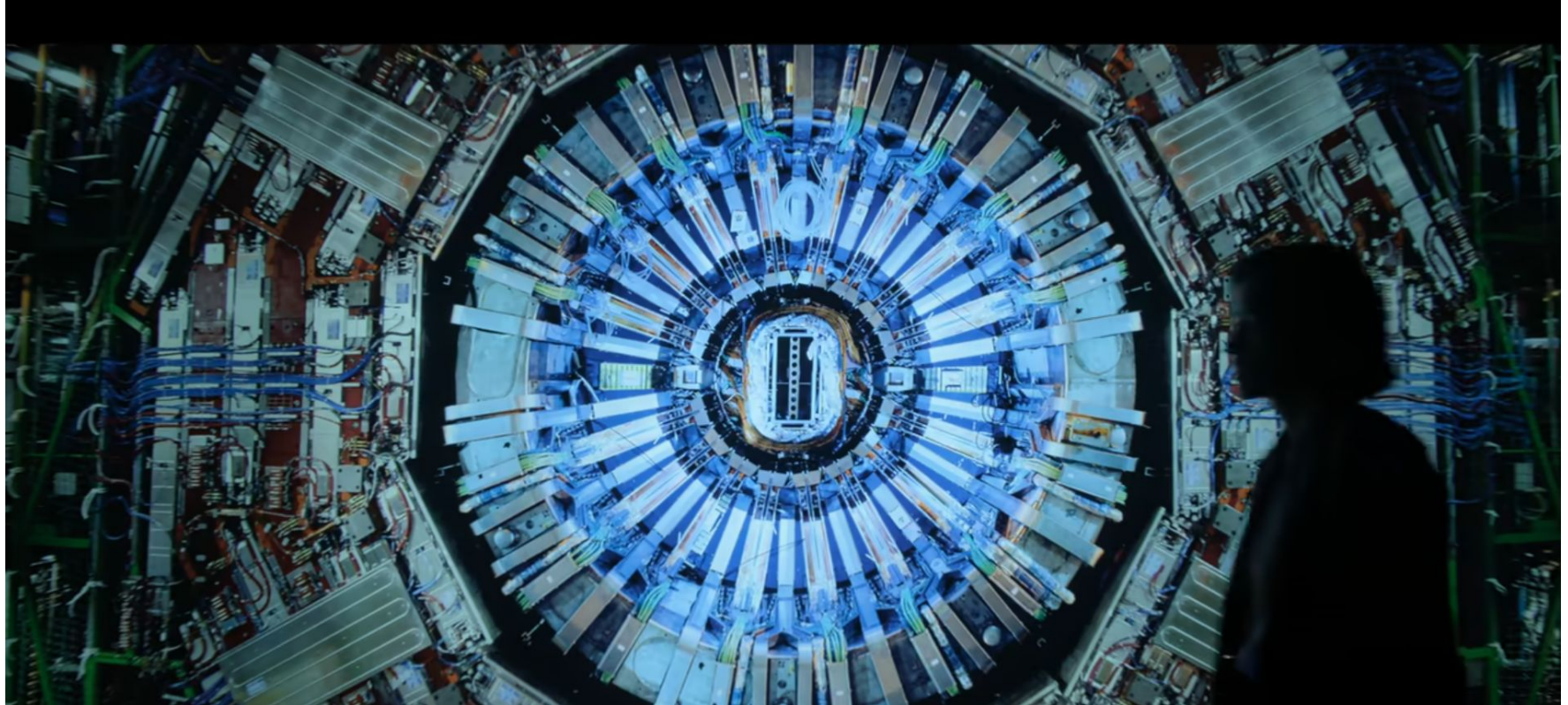


The Three Body Problem :: CERN's Super Proton Synchrotron

<https://cerncourier.com/a/the-sps-gets-ready-to-restart/>



The Three Body Problem :: “Oxford”



The Three Body Problem :: CMS Experiment At CERN

https://cms.cern/sites/default/files/2022-06/detector_stripe01-min_0_0.jpeg

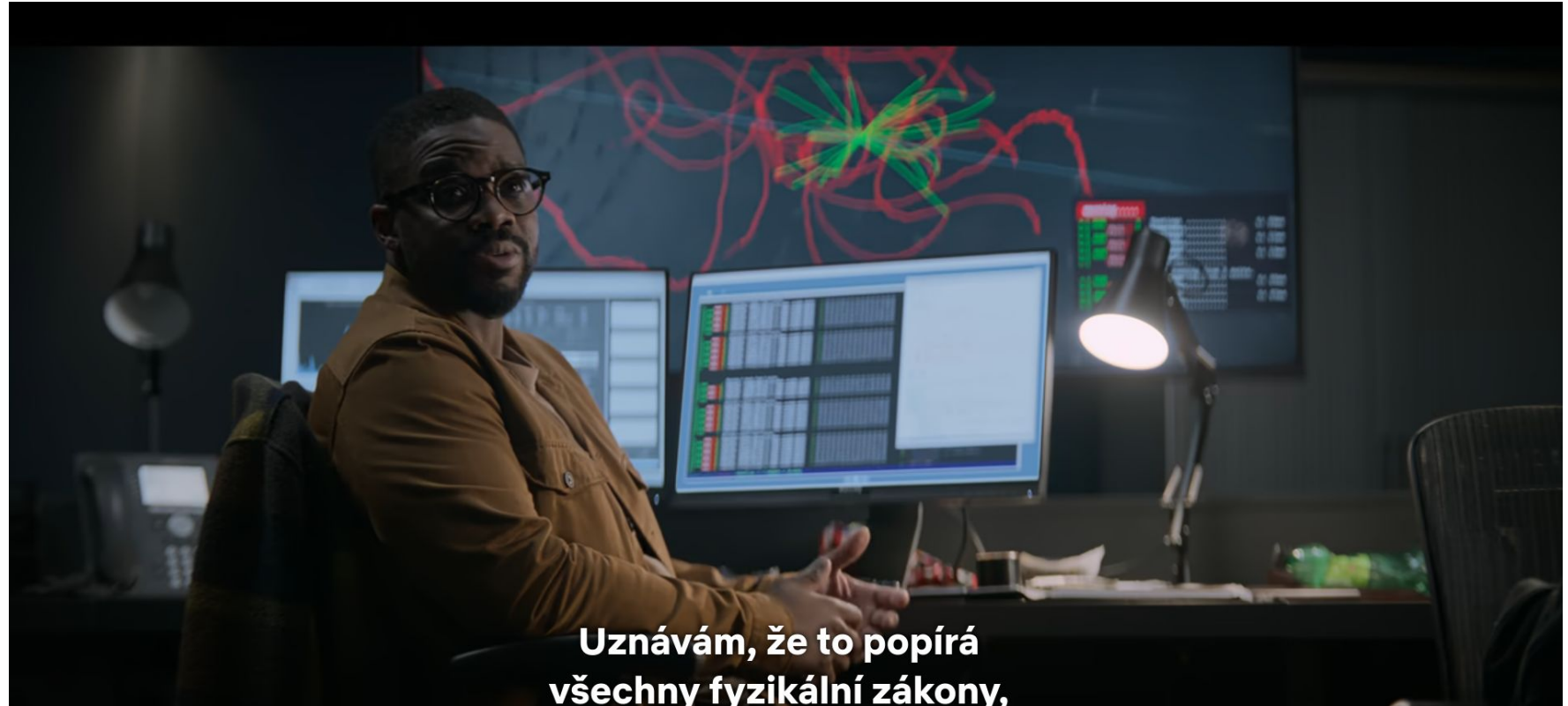
<https://cms.cern/detector>



The Three Body Problem :: “Oxford”



The Three Body Problem :: “Oxford”



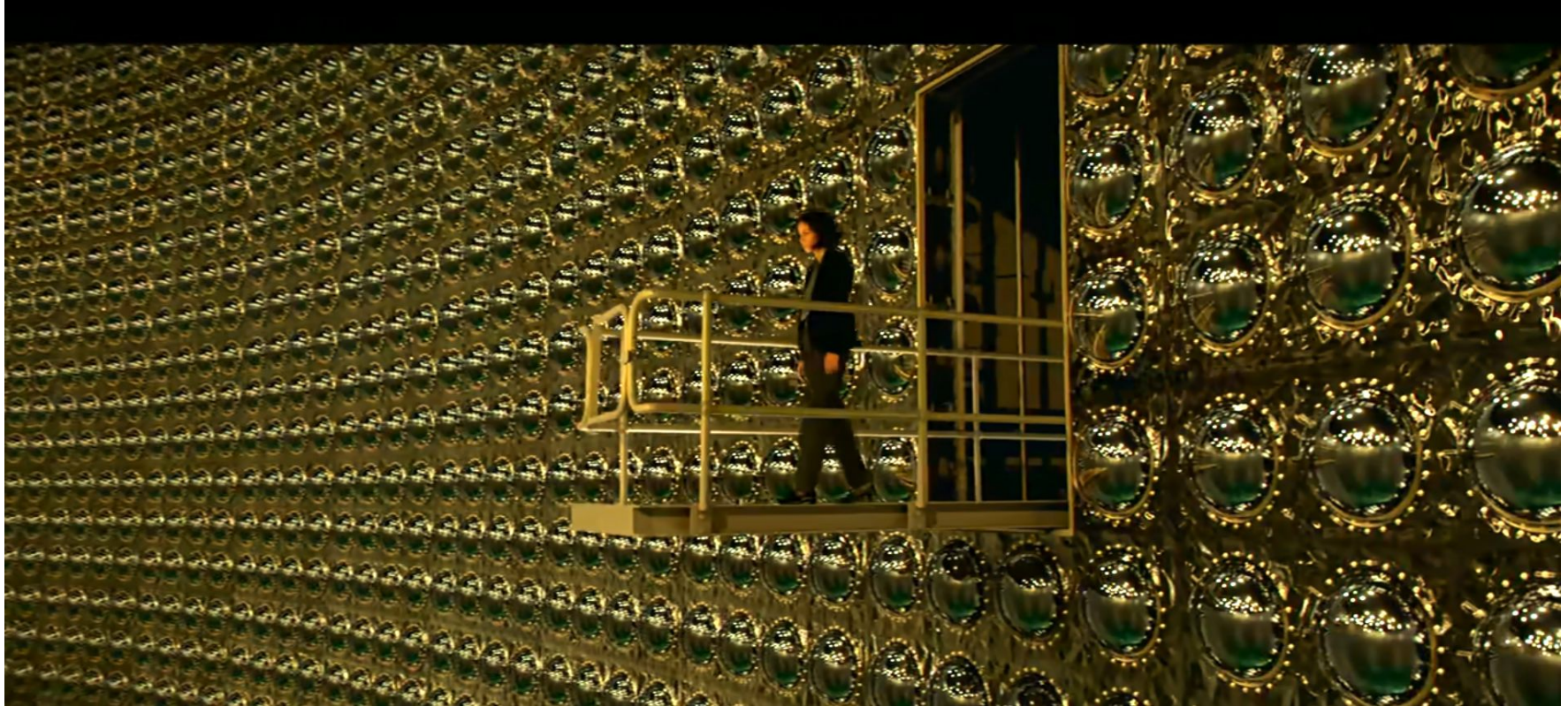
**Uznávám, že to popírá
všechny fyzikální zákony,**

The Three Body Problem :: ATLAS Exp. Control Room

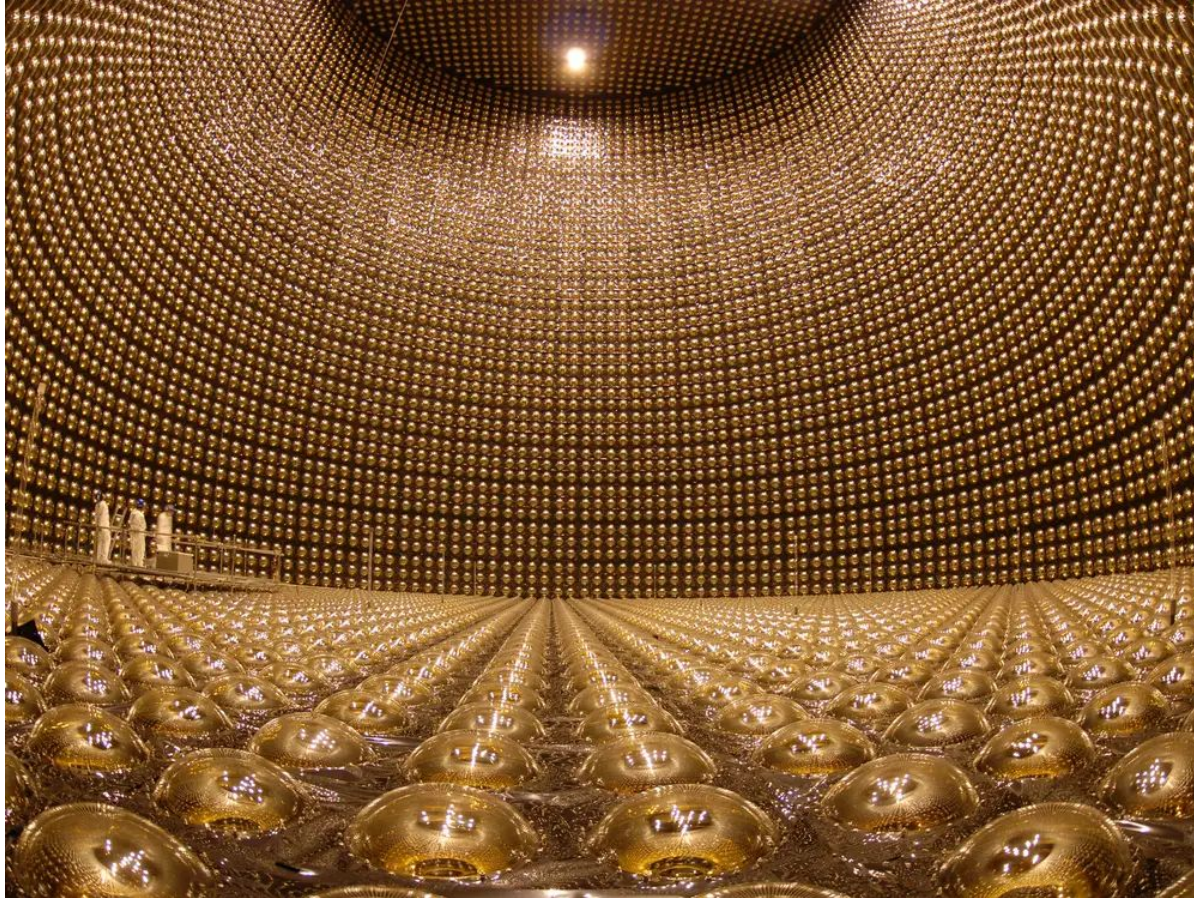


https://commons.wikimedia.org/wiki/File:ATLAS_experiment_control_room.jpg

The Three Body Problem :: “Oxford”



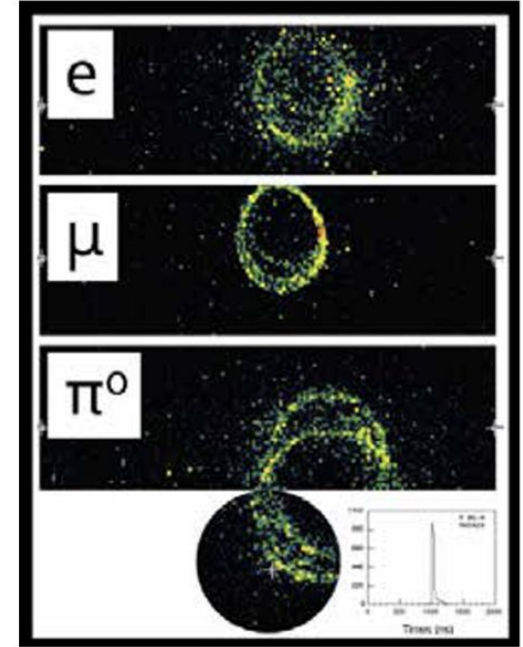
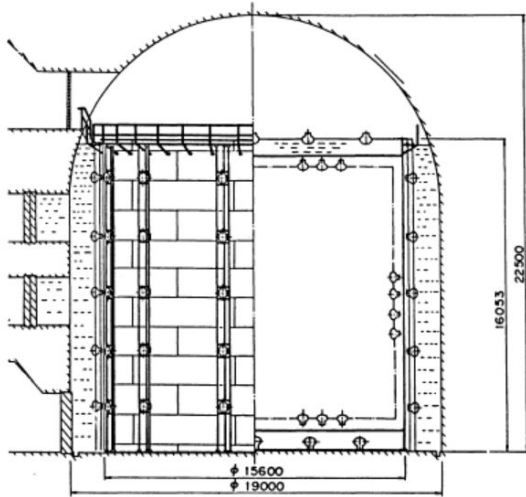
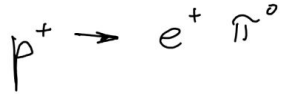
The Three Body Problem :: Super-Kamiokande Exp.



<https://www.businessinsider.com/super-kamiokande-neutrino-detector-is-unbelievably-beautiful-2018-6>

<https://www-sk.icrr.u-tokyo.ac.jp/en/sk/>

Experiment Kamiokande(-II) :: 3kt H₂O



Neutrino identification

Figure 2: The detector of KAMIOKANDE -II. The dimensions are given in millimeters.

<https://articles.adsabs.harvard.edu//full/1987ESOC...26..219K/0000220.000.html>

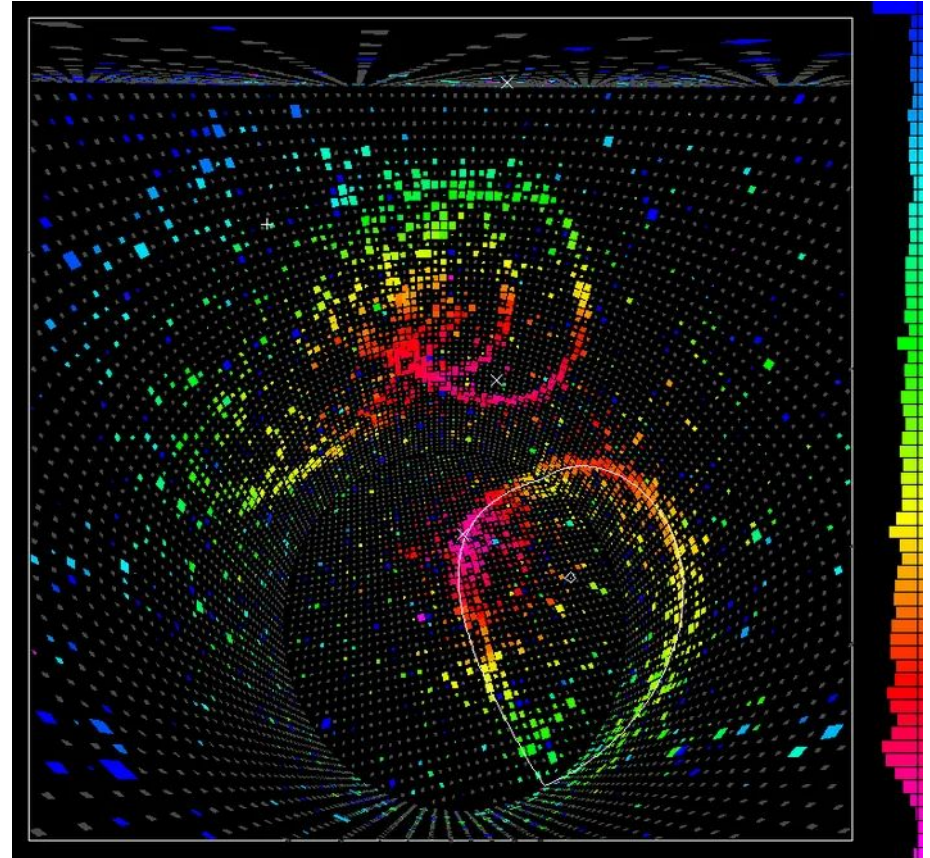
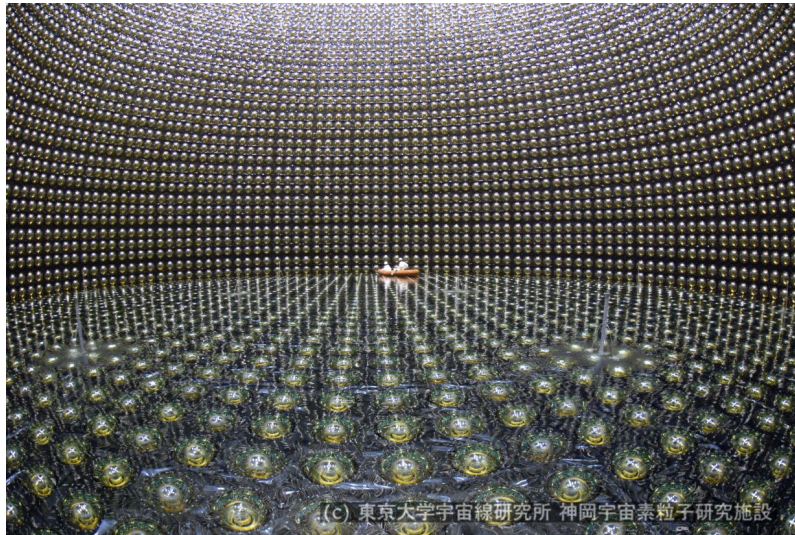
Super Kamiokande :: 50kt H₂O

Kamioka Nucleon Decay Experiment

<https://www-sk.icrr.u-tokyo.ac.jp/en/sk/about/history/>

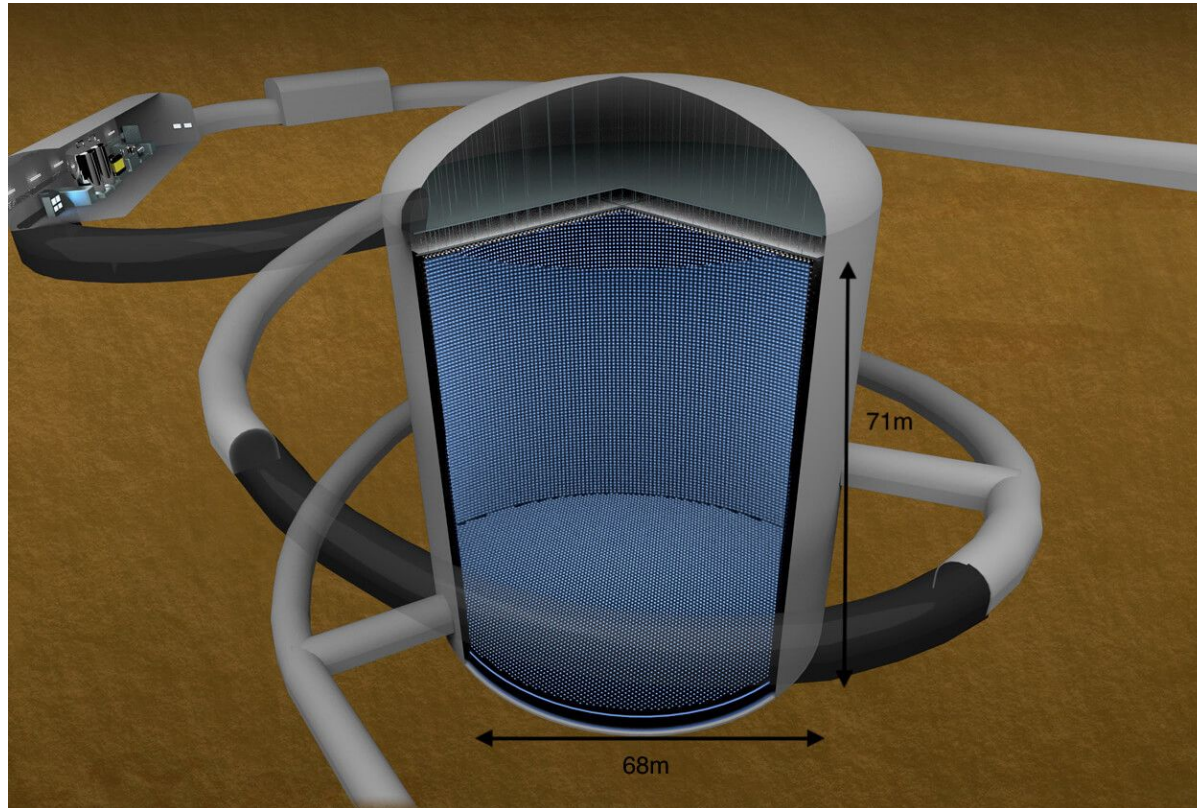
<https://www-sk.icrr.u-tokyo.ac.jp/en/sk/experience/gallery/>

<https://www.forbes.com/sites/startswithabang/2018/12/05/is-there-really-a-fourth-neutrino-out-there-in-the-universe/>

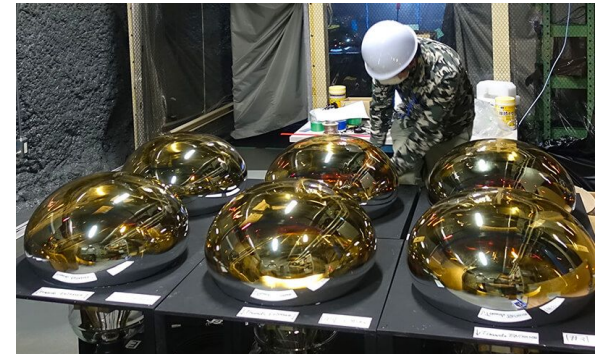


Hyper Kamiokande :: 250kt H₂O

<https://www-sk.icrr.u-tokyo.ac.jp/en/hk/about/detector/>
<https://www-sk.icrr.u-tokyo.ac.jp/en/hk/>



YT :: Hyper-Kamiokande
https://www.youtube.com/watch?v=JFOE3D2z7LM&t=12s&ab_channel=Hyper-Kamiokande



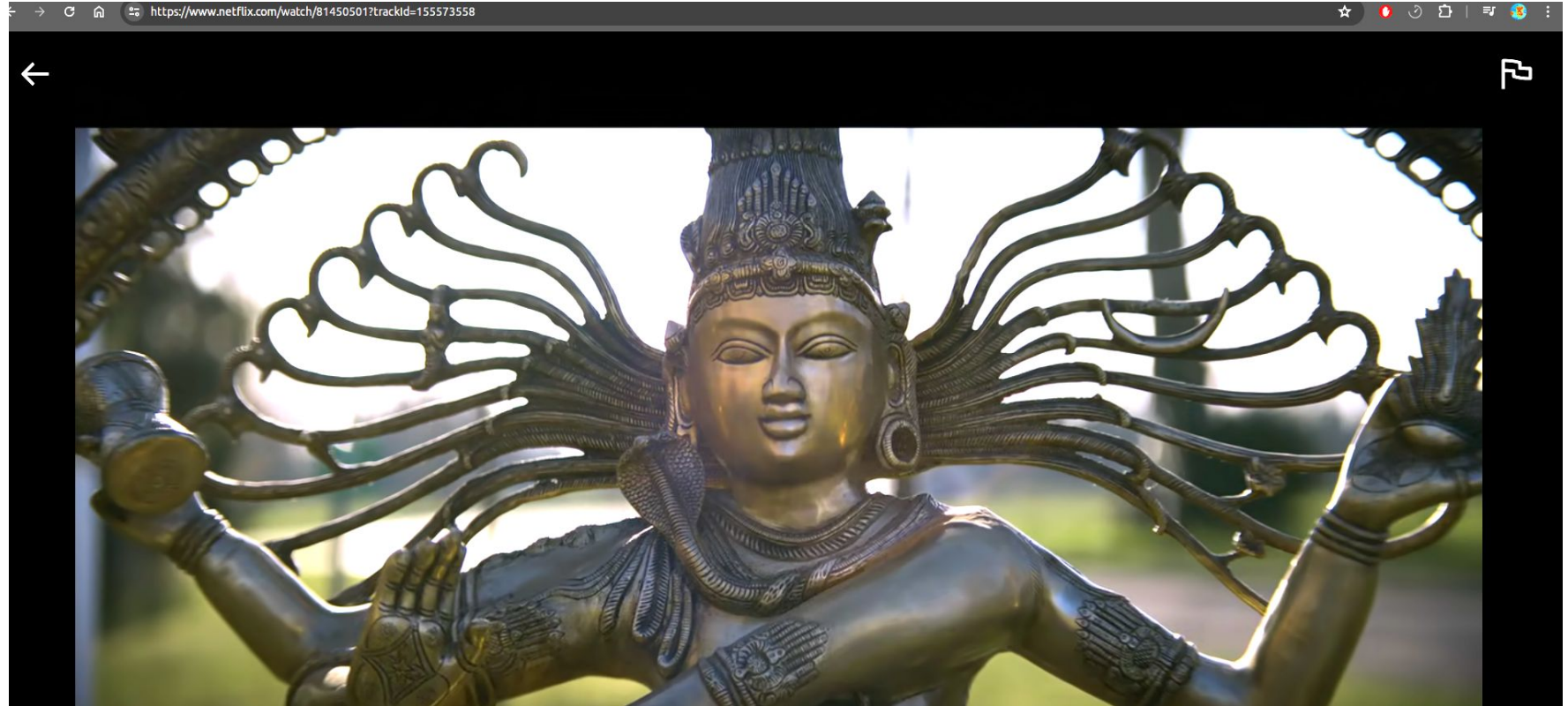
Hyper Kamiokande :: 250kt H₂O :: Dec. 2023



The Three Body Problem :: CERN



The Three Body Problem :: CERN



The Three Body Problem :: CERN – realita



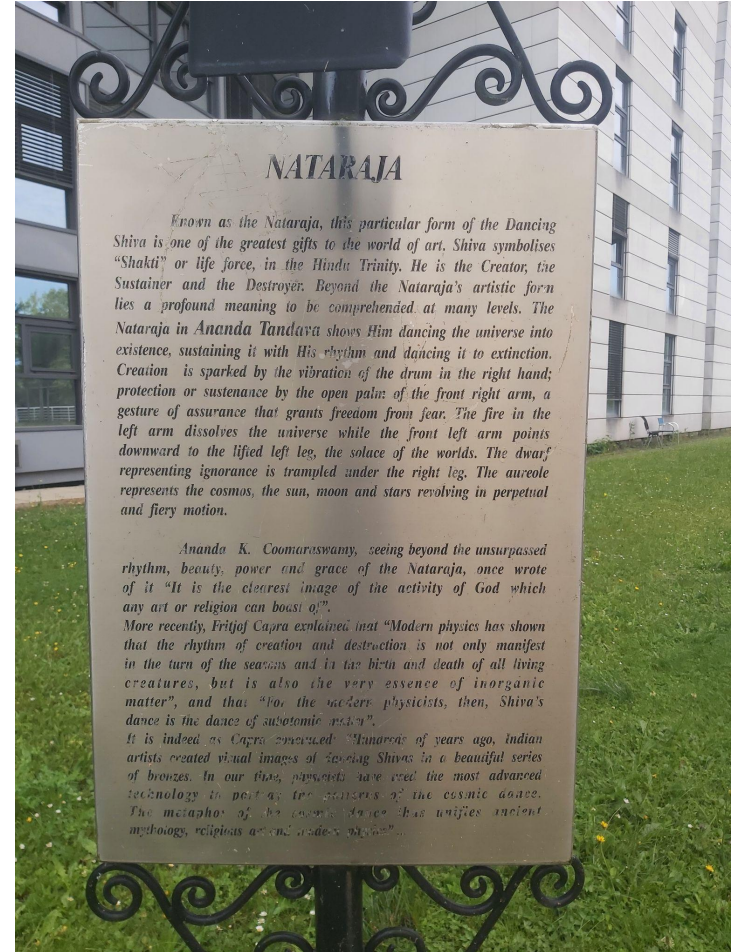
The Three Body Problem :: CERN – realita

नित्याय त्रिगुणात्मने पुरजिते कात्यायनी-श्रेयसे
सत्यायादिकुटुंबिने मुनिमनः प्रत्यक्ष-चिन्मूर्तये ।
मायासृष्ट-जगत्रयाय सकलाम्नायान्त-संचारिणे
सायं ताण्डव-संभ्रमाय जटिने सेयं नतिशंभवे ॥ ५६ ॥

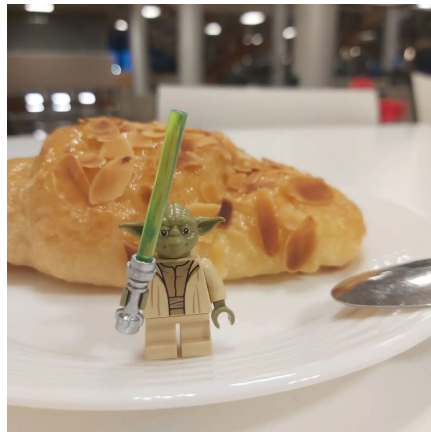
“O Omnipresent, the embodiment of all virtues, the creator of this cosmic universe, the king of dancers, who dances the *Ananda Tandava* in the twilight, I salute thee.”

(Source: Verse No. 56, Sivanandalahari by Sri Adi Sankara)

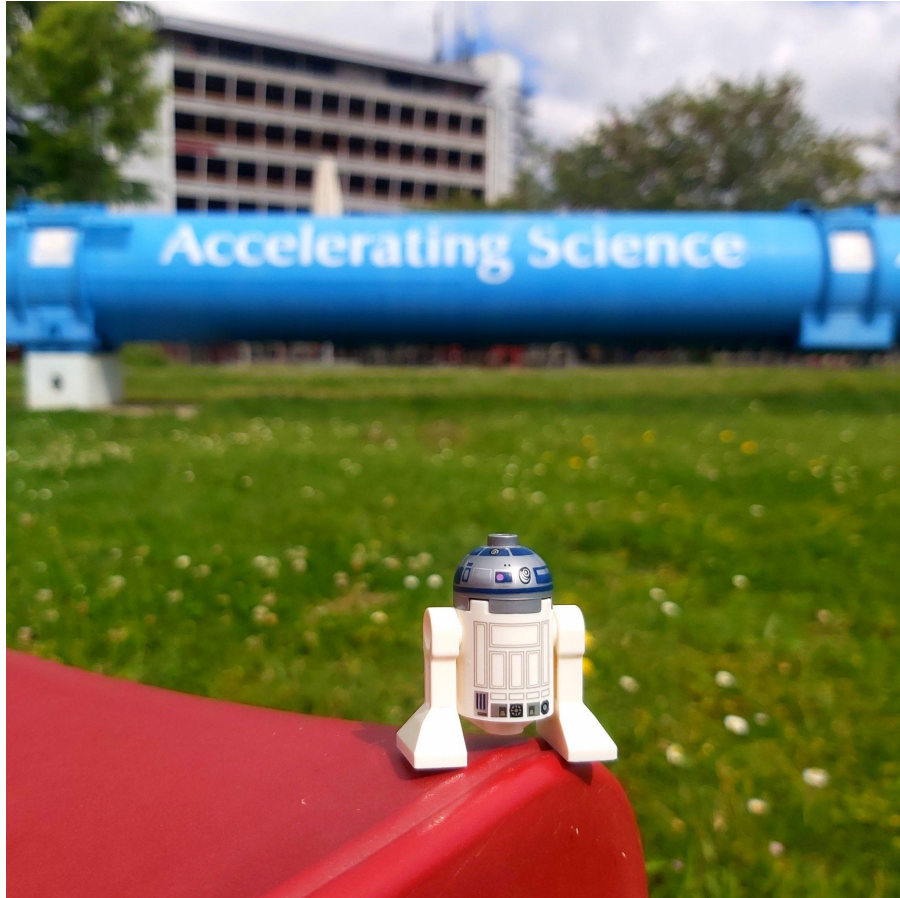
Presented by: The Department of Atomic Energy, Government of India.

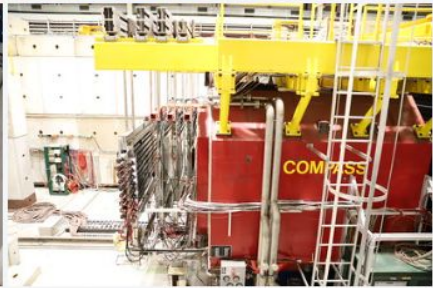
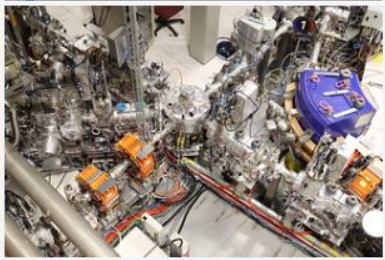
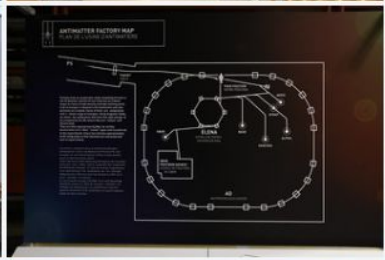


CERN

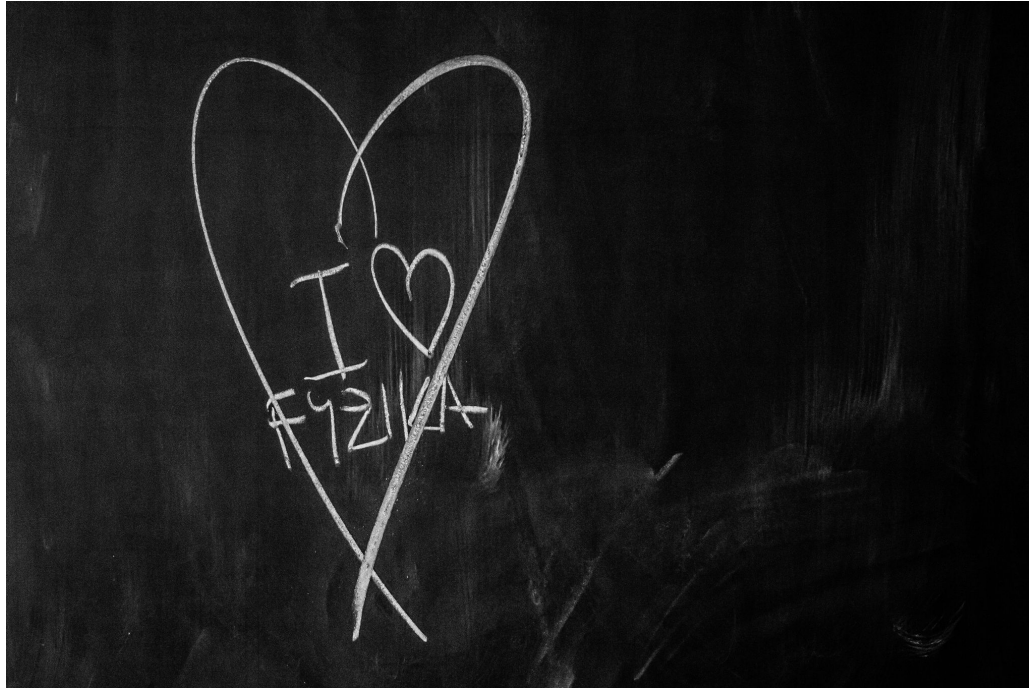








Backup



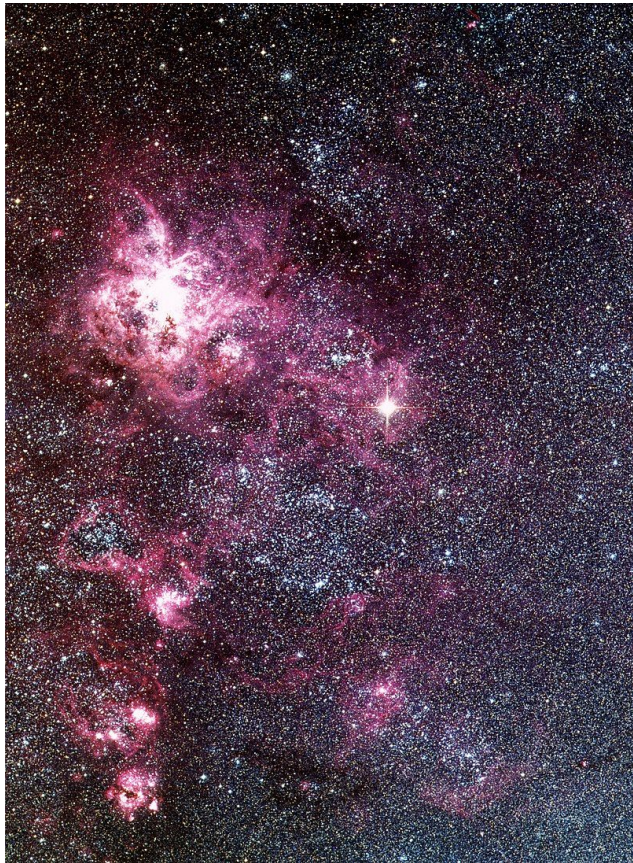
SN1987A 24.2.1987 :: 170kly



SN1987A 24.2.1987 :: 170kly



SN1987A 24.2.1987 :: 170kly



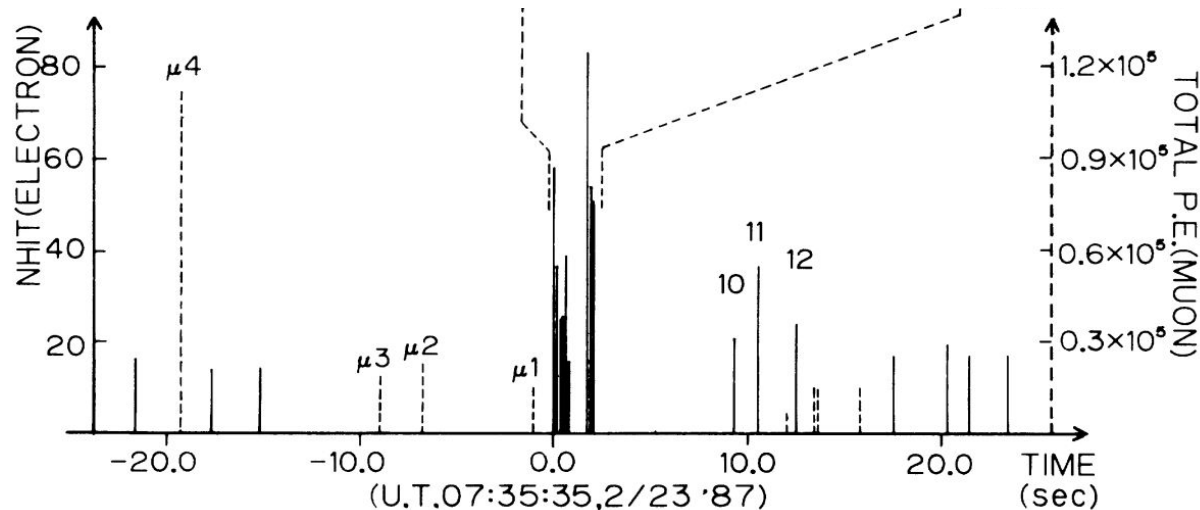
https://en.wikipedia.org/wiki/SN_1987A



<https://www-sk.icrr.u-tokyo.ac.jp/en/news/detail/324>

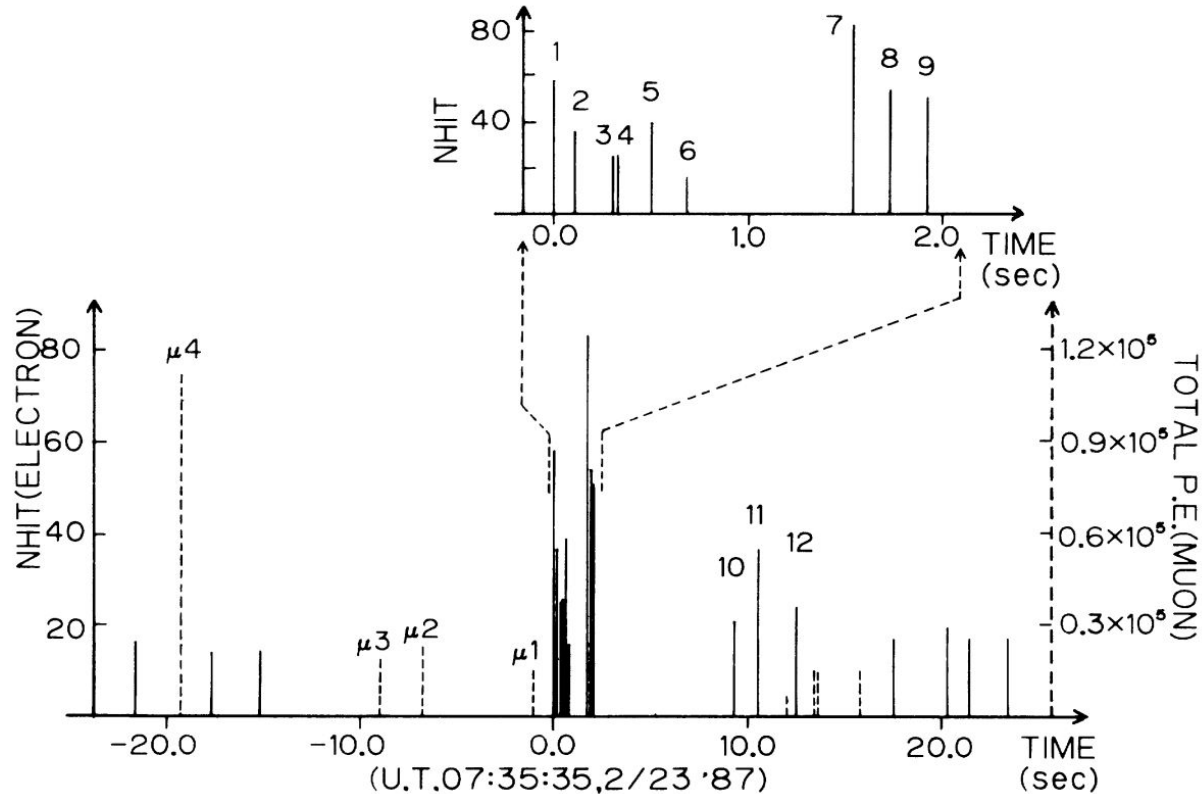
Observation of a Neutrino Burst from the Supernova SN1987A

K. Hirata,^(a) T. Kajita,^(a) **M. Koshiba,**^(a,b) M. Nakahata,^(b) Y. Oyama,^(b)
N. Sato,^(c) A. Suzuki,^(b) M. Takita,^(b) and Y. Totsuka^(a,c)
University of Tokyo, Tokyo 113, Japan



Experiment Kamiokande(-II) :: 3kt H₂O

- Počátky neutrinové astronomie! :-)



Experiment Kamiokande(-II) :: 3kt H₂O

- Počátky neutrinové astronomie! :-)
- 1 Erg is a unit of energy equal to 10^{-7} Joules (100 nJ)

form volume distribution. Additional support is provided by the correlation in angle of the first two observed events with the direction to SN1987A. The event burst occurred roughly 18 h prior to the first optical sighting.¹

Correcting for energy-dependent detection efficiency, and assuming that nine of the twelve events are due to $\bar{\nu}_e p \rightarrow e^+ n$, we obtain an integral flux of $1.0 \times 10^{10} \bar{\nu}_e \text{ cm}^{-2}$ for the burst, where the $\bar{\nu}_e$ energy (the observed electron energy plus 1.3 MeV) is above 8.8 MeV. This, in turn, leads to the $\bar{\nu}_e$ output of SN1987A of 8×10^{52} ergs for an assumed average energy of 15 MeV.

This observation is the first direct observation in neutrino astronomy, and coincides remarkably well with the current model of supernova collapse and neutron-star formation.⁶ In that model an aged, massive star, having exhausted its nuclear fuel, undergoes a supernova explosion. In supernovae of Type II almost all of the gravitational binding energy of the resultant neutron star, $\sim 3 \times 10^{53}$ ergs, is radiated within a few seconds in the form of 10^{58} neutrinos of all flavors with average energy in the vicinity of 10–15 MeV.

Experiment Kamiokande(-II) :: 3kt H₂O

- Počátky neutrinové astronomie! :-)
- 1 Erg is a unit of energy equal to 10^{-7} Joules (100 nJ)

form volume distribution. Additional support is provided by the correlation in angle of the first two observed events with the direction to SN1987A. The event burst occurred roughly 18 h prior to the first optical sighting.¹

Correcting for energy-dependent detection efficiency, and assuming that nine of the twelve events are due to $\bar{\nu}_e p \rightarrow e^+ n$, we obtain an integral flux of $1.0 \times 10^{10} \bar{\nu}_e \text{ cm}^{-2}$ for the burst, where the $\bar{\nu}_e$ energy (the observed electron energy plus 1.3 MeV) is above 8.8 MeV. This, in turn, leads to the $\bar{\nu}_e$ output of SN1987A of 8×10^{52} ergs for an assumed average energy of 15 MeV.

This observation is the first direct observation in neutrino astronomy, and coincides remarkably well with the current model of supernova collapse and neutron-star formation.⁶ In that model an aged, massive star, having exhausted its nuclear fuel, undergoes a supernova explosion. In supernovae of Type II almost all of the gravitational binding energy of the resultant neutron star, $\sim 3 \times 10^{53}$ ergs, is radiated within a few seconds in the form of 10^{58} neutrinos of all flavors with average energy in the vicinity of 10–15 MeV.

Experiment Kamiokande(-II) :: 3kt H₂O

- Počátky neutrinové astronomie! :-)
- 1 Erg is a unit of energy equal to 10^{-7} Joules (100 nJ)

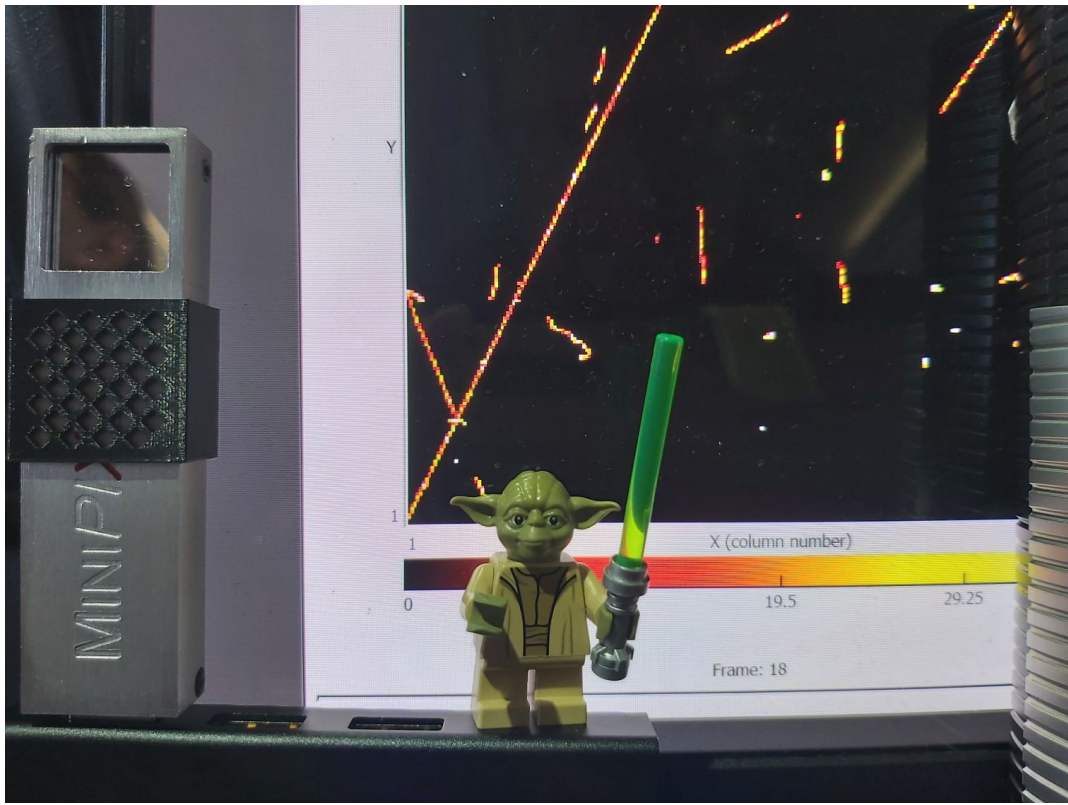
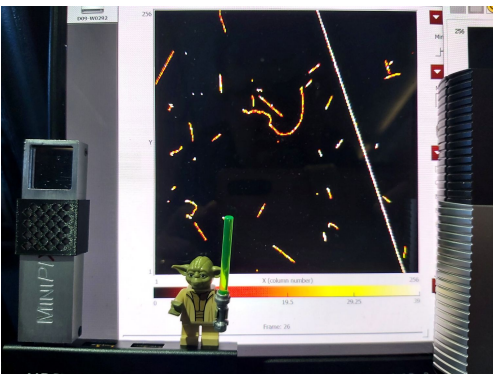
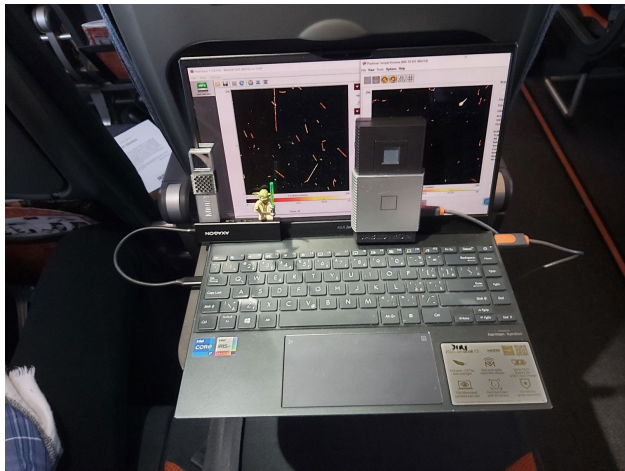
form volume distribution. Additional support is provided by the correlation in angle of the first two observed events with the direction to SN1987A. The event burst occurred roughly 18 h prior to the first optical sighting.¹

Correcting for energy-dependent detection efficiency, and assuming that nine of the twelve events are due to $\bar{\nu}_e p \rightarrow e^+ n$, we obtain an integral flux of $1.0 \times 10^{10} \bar{\nu}_e \text{ cm}^{-2}$ for the burst, where the $\bar{\nu}_e$ energy (the observed electron energy plus 1.3 MeV) is above 8.8 MeV. This, in turn, leads to the $\bar{\nu}_e$ output of SN1987A of 8×10^{52} ergs for an assumed average energy of 15 MeV.

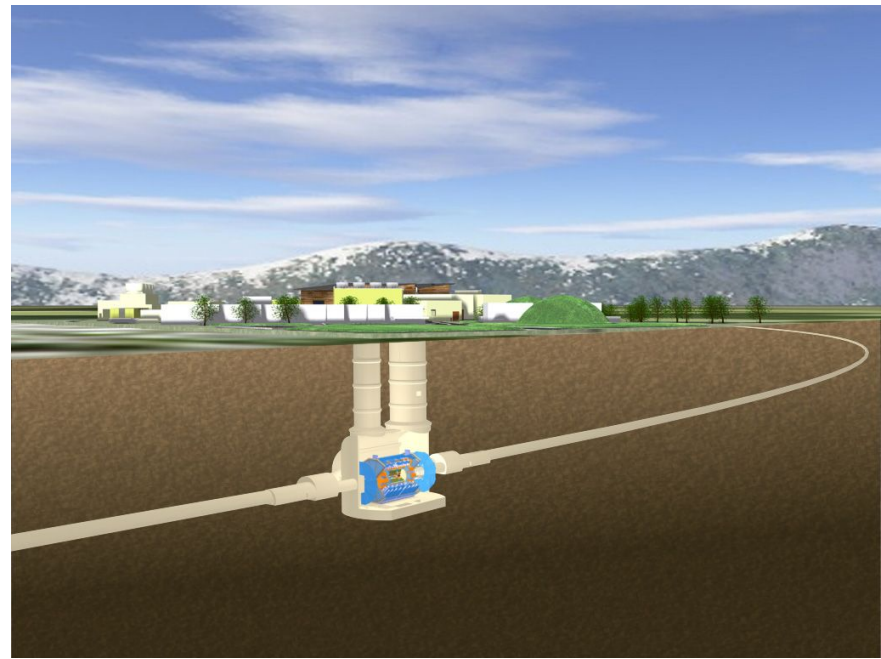
This observation is the first direct observation in neutrino astronomy, and coincides remarkably well with the current model of supernova collapse and neutron-star formation.⁶ In that model an aged, massive star, having exhausted its nuclear fuel, undergoes a supernova explosion.

In supernovae of Type II almost all of the gravitational binding energy of the resultant neutron star, $\sim 3 \times 10^{53}$ ergs, is radiated within a few seconds in the form of 10^{58} neutrinos of all flavors with average energy in the vicinity of 10–15 MeV.

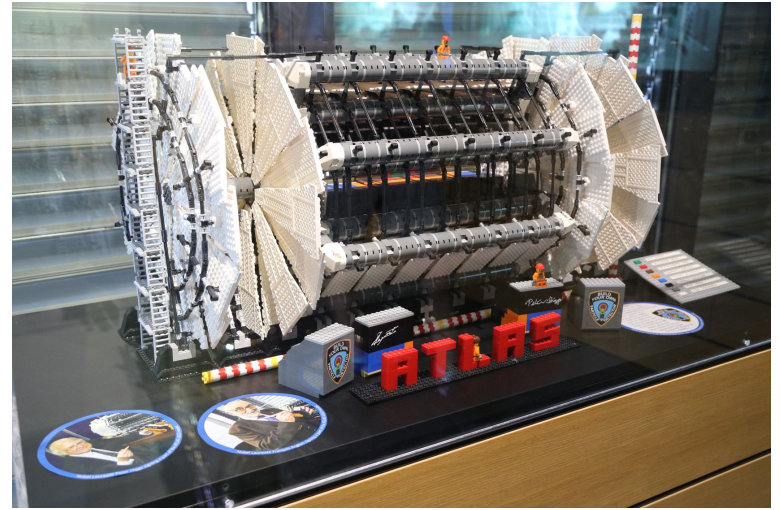
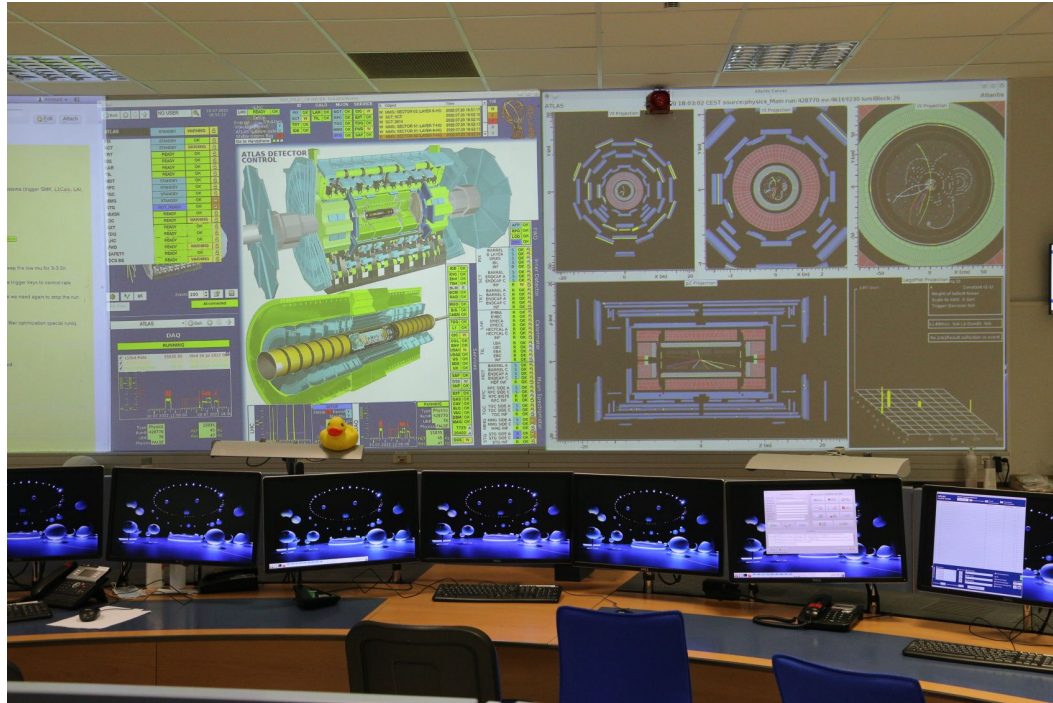
Flying to CERN – with particle cameras!



The ATLAS Experiment



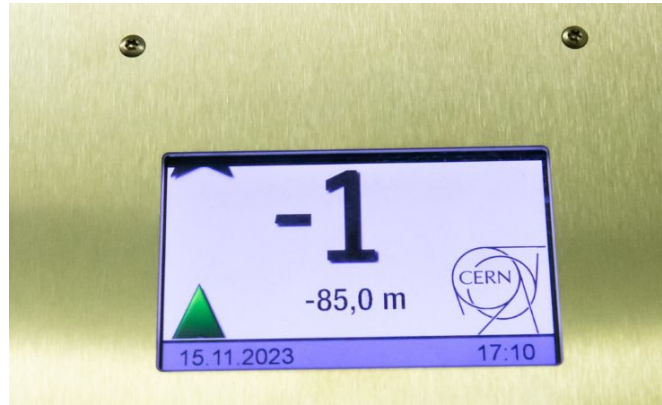
The ATLAS Experiment



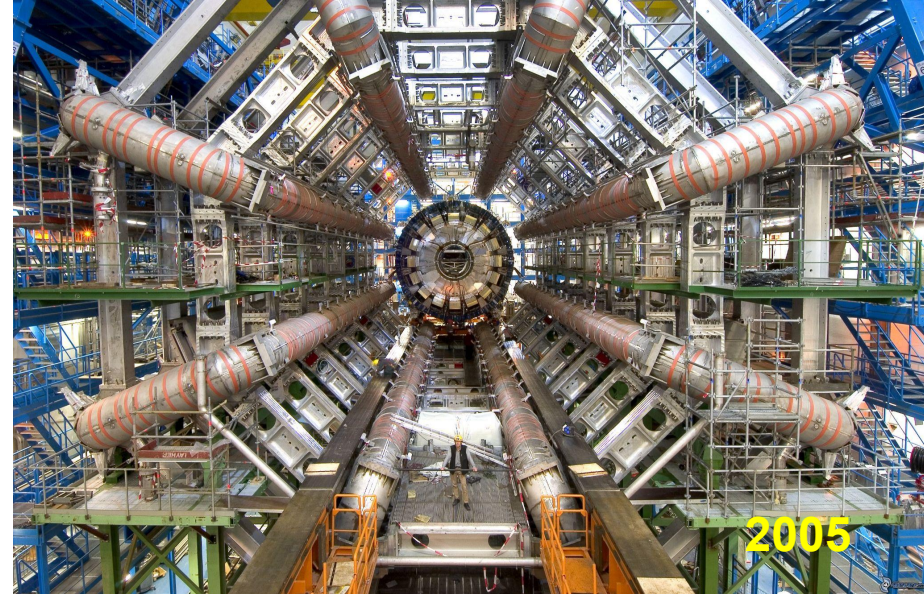
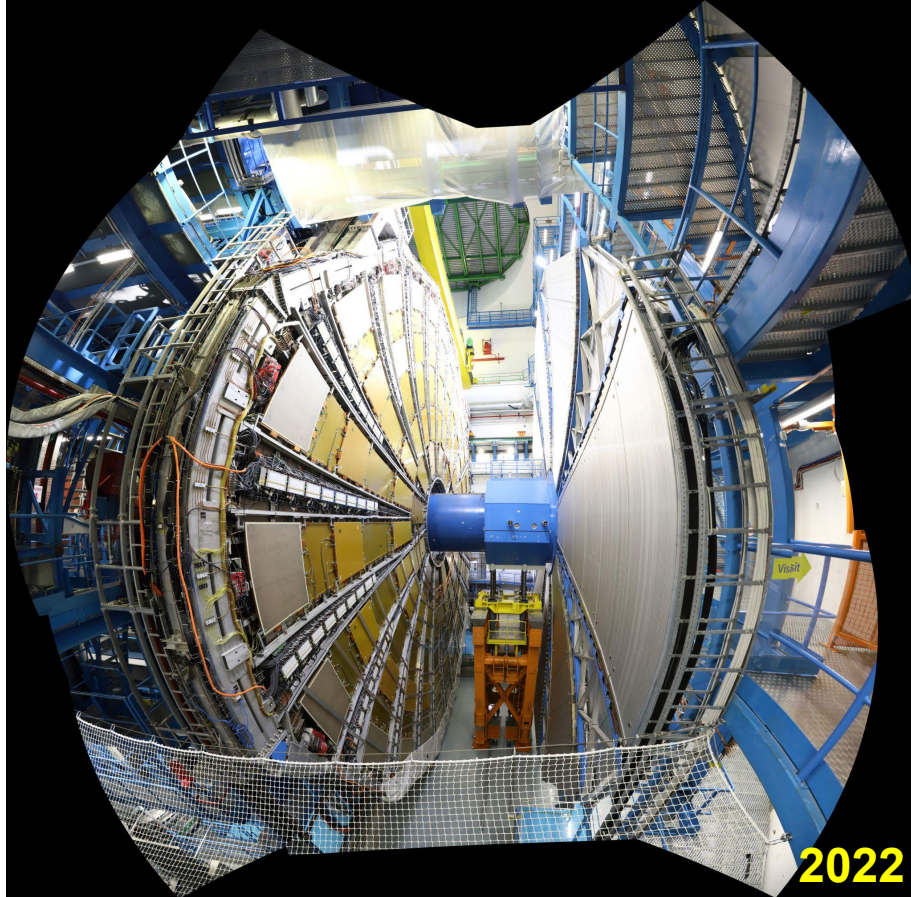
The ATLAS Experiment



The ATLAS Experiment

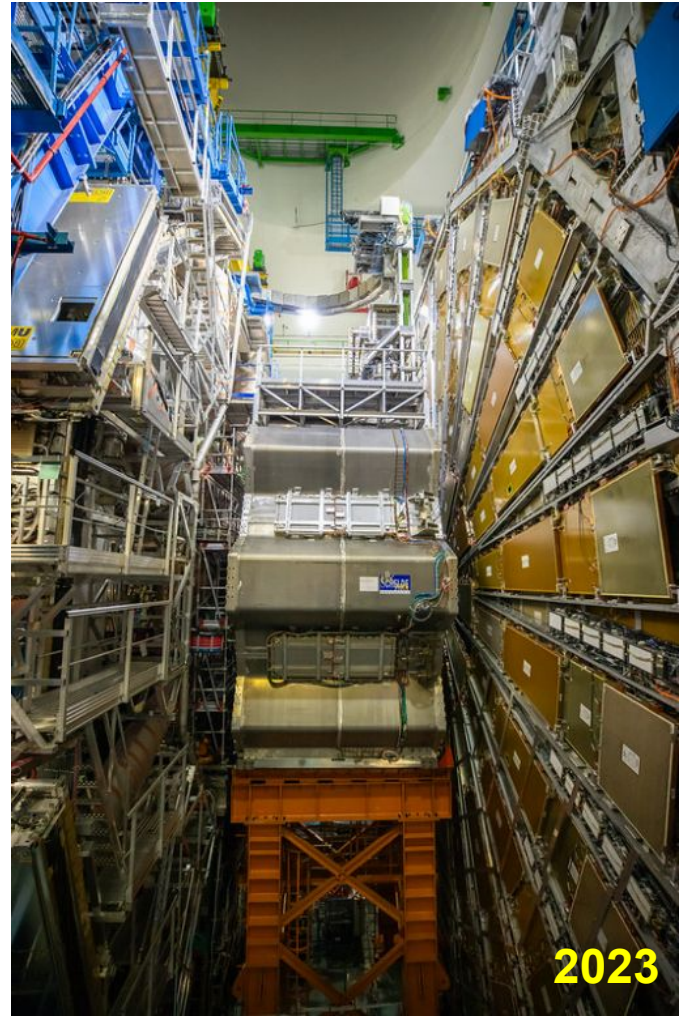


The ATLAS Experiment

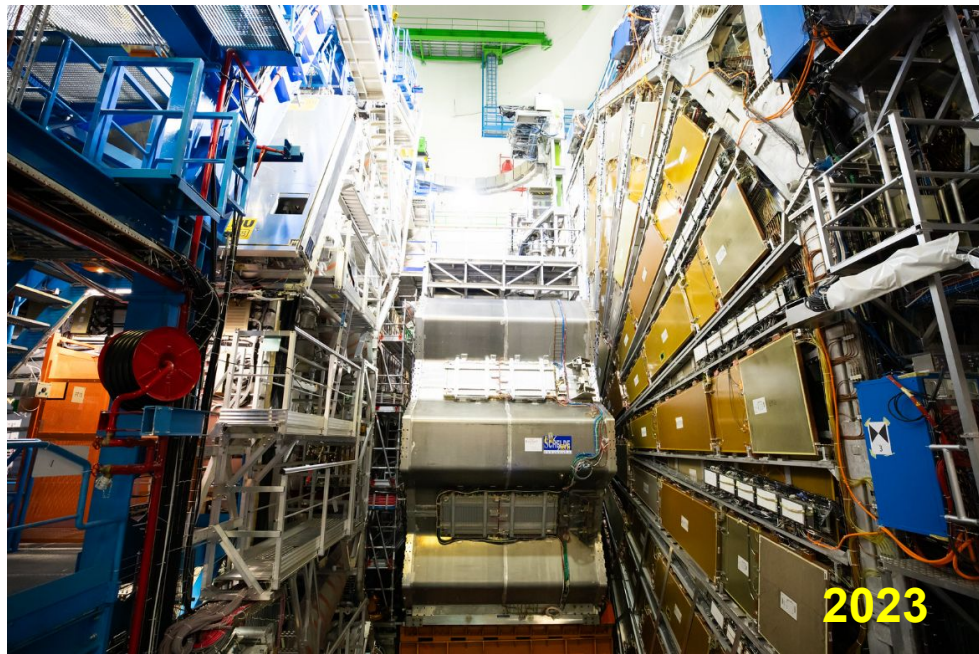


<https://timeline.web.cern.ch/timeline-header/143>

The ATLAS Experiment



The ATLAS Experiment



SWI swissinfo.ch

Swiss perspectives in 10 languages



▲ French President Emmanuel Macron, left, with Fabiola Gianotti, Director General of CERN, and Swiss President Alain Berset in the ATLAS experiment on Thursday © Keystone / Martial Trezzini



read aloud

French President Emmanuel Macron spoke on Thursday of France's ambitions to remain in "first place" during a visit to the world's most powerful particle accelerator on the Franco-Swiss border, where a successor, even more powerful, is being studied.

November 17, 2023 - 10:33

🕒 2 minutes

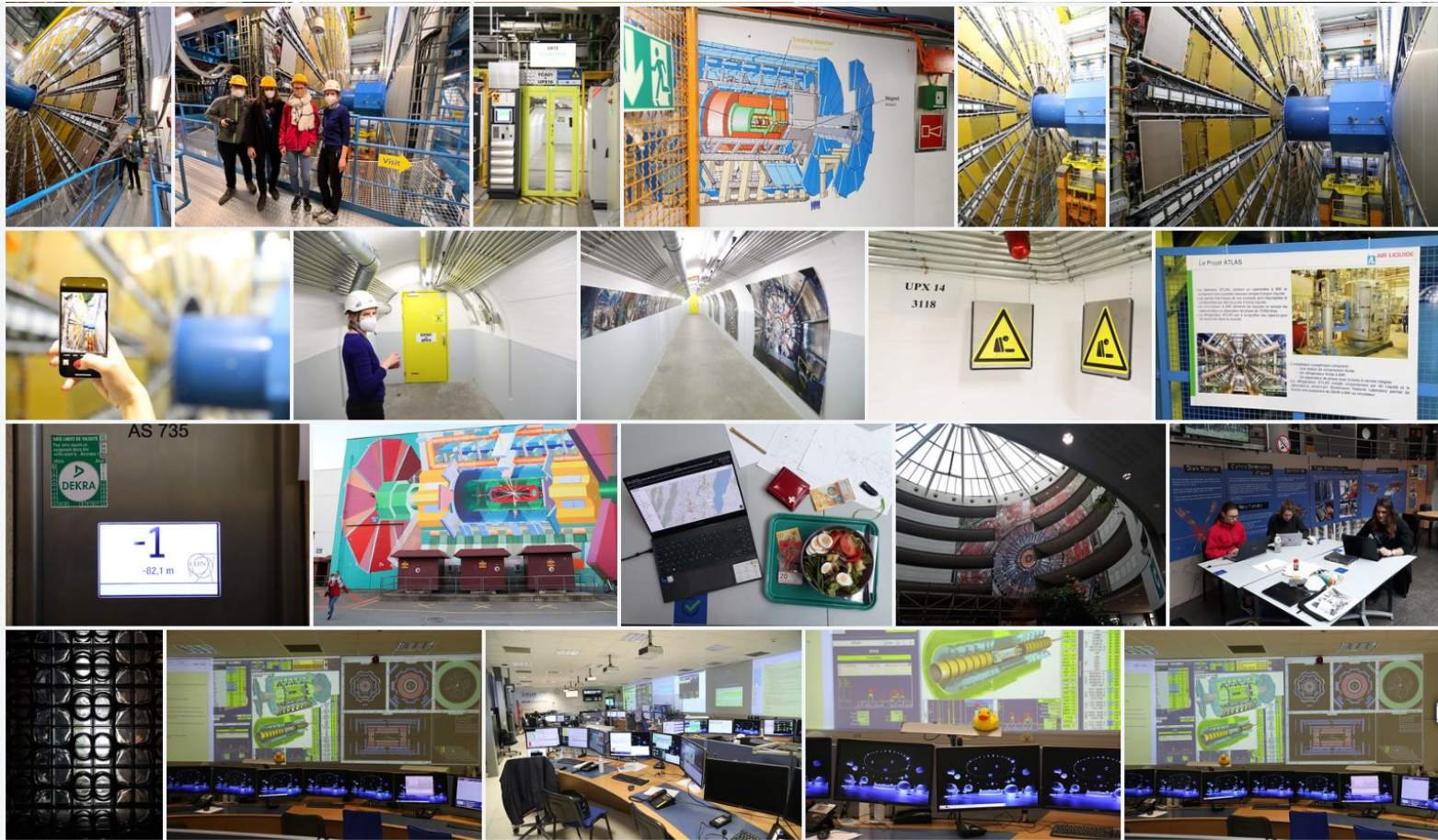
The ATLAS Experiment



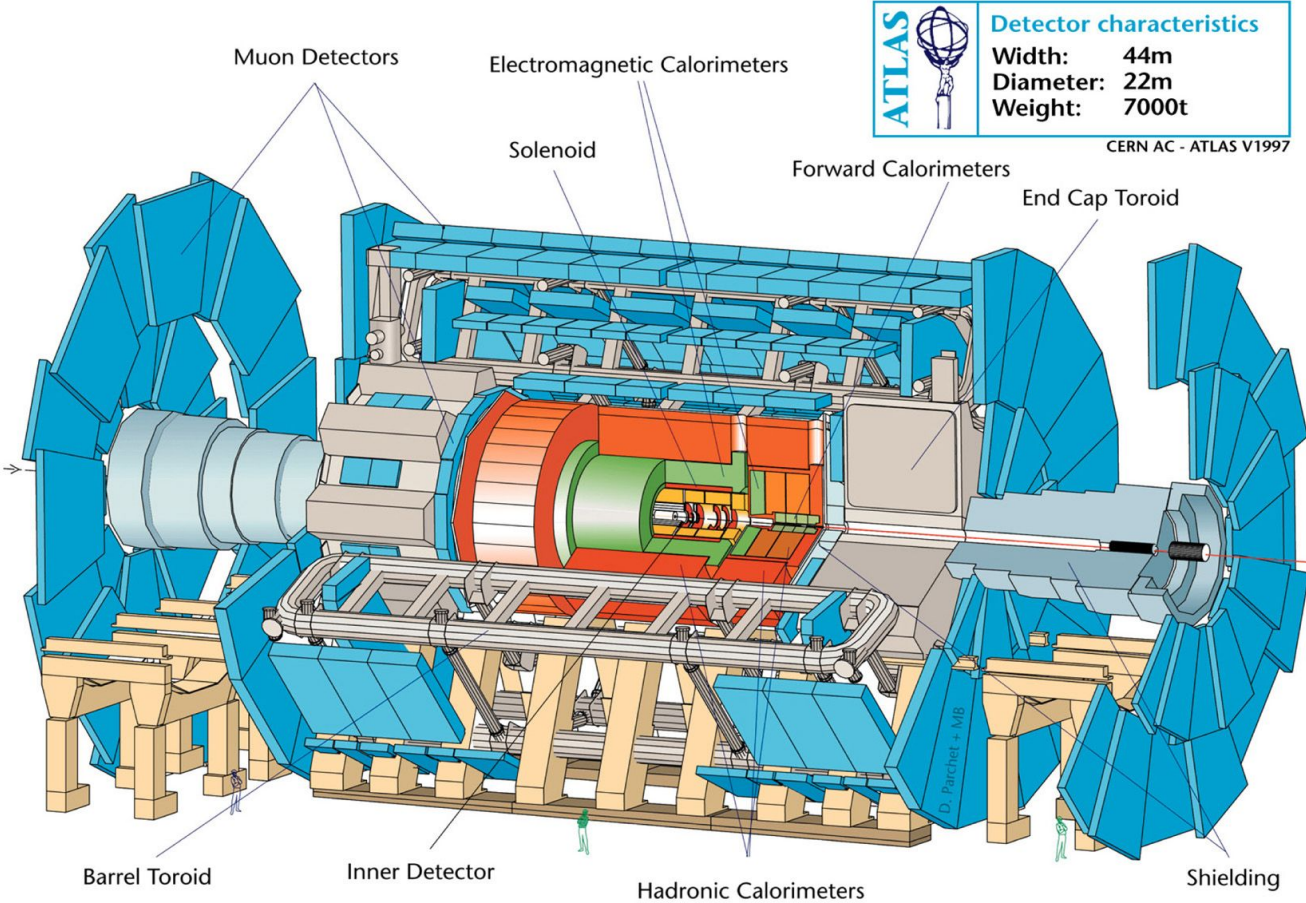
The ATLAS Experiment



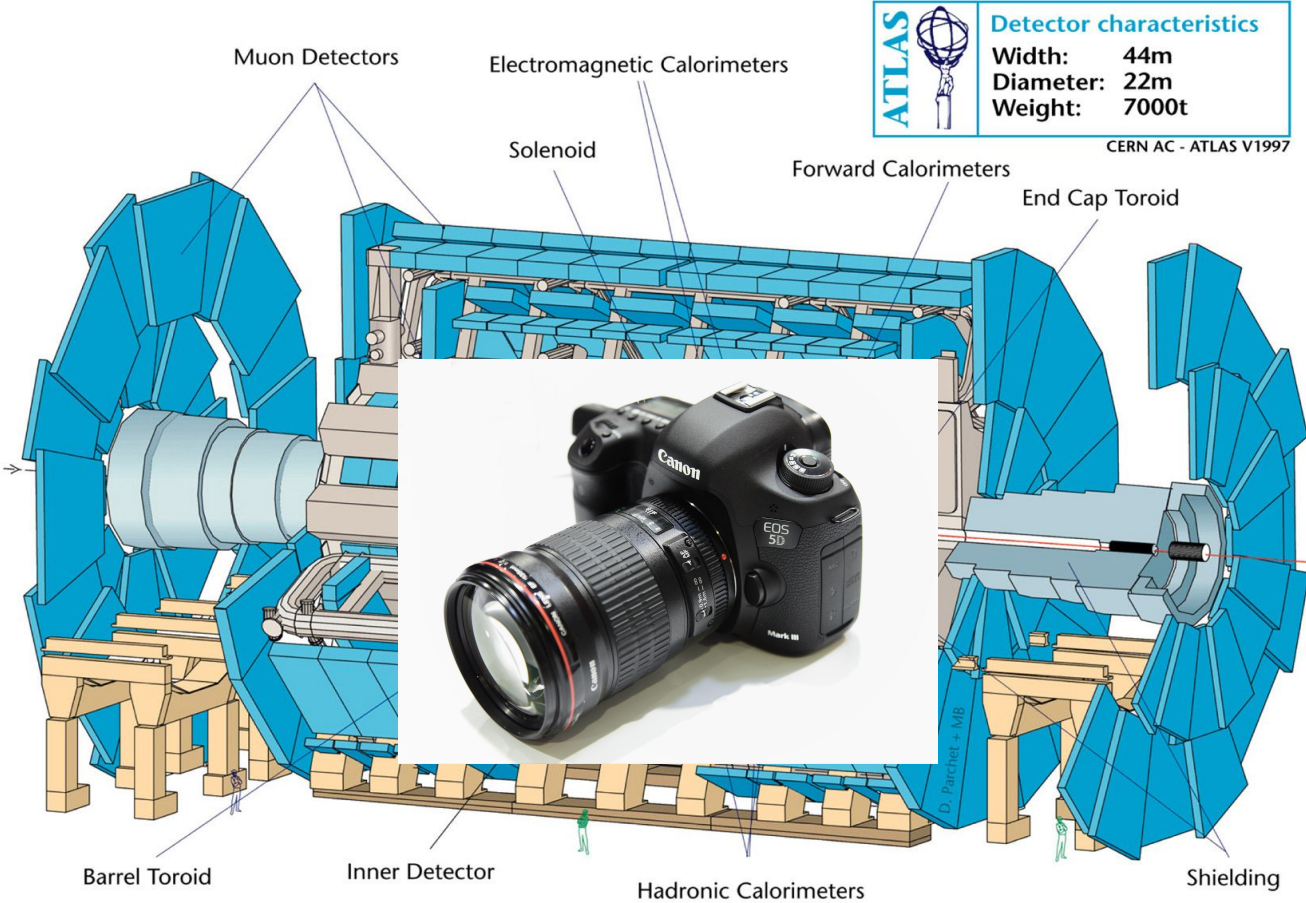
The ATLAS Experiment



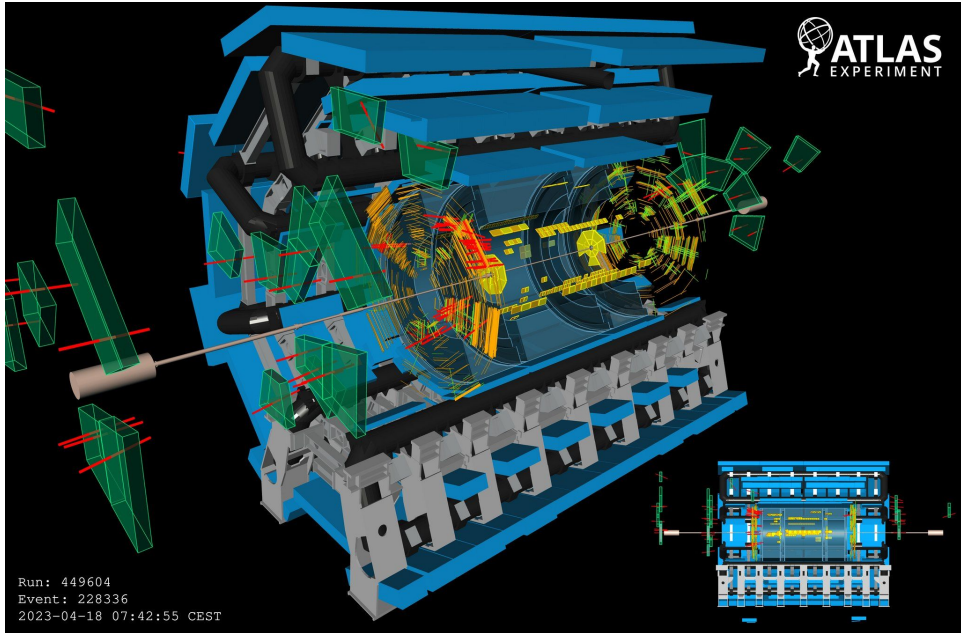
The ATLAS Experiment



The ATLAS Experiment

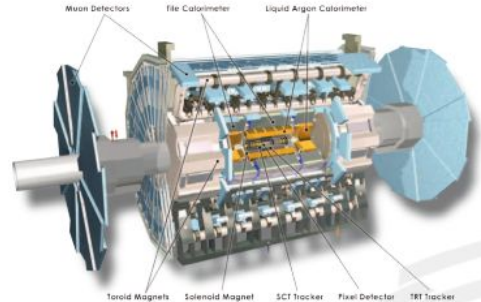


The ATLAS Experiment



The ATLAS Detector

- Diameter 25 m -- Length : 46 m
- Barrel toroid length 26 m
- Overall weight 7 000 tonnes
- ~ 100 million electronic channels
- ~ 3 000 km of cables



Calorimeters

Measure the energies carried by the particles

- Liquid Argon (LAr) Calorimeter**
- Barrel 6.4 m long, 53 cm thick, 110 000 channels.
 - Works with Liquid Argon at -183° C
 - LAr endcap consists of the forward calorimeter, electromagnetic (EM) and hadronic endcaps.
 - EM endcaps each have thickness 0.632 m and radius 2.077 m.
 - Hadronic endcaps consist of two wheels of thickness 0.8 m and 1.0 m with radius 2.09 m.
 - Forward calorimeter has three modules of radius 0.455 m and thickness 0.450 m each.

- Tile Calorimeter (TileCal)**
- Barrel made of 64 wedges, each 5.6 m long and 20 tonnes.
 - Each Endcap has 64 wedges, each 2.6 m long.
 - 500 000 plastic scintillator tiles.

Muon System

Identifies and measures the momenta of muons

- Thin Gap Chambers**
- For triggering and 2nd coordinate measurement (non-bending direction) at ends of detector.
- 440 000 channels
- Resistive Plate Chambers**
- For triggering and 2nd coordinate measurement in central region.
- 380 000 channels
 - Electric Field 5 000 V/mm
- Monitored Drift Tubes**
- Measure curves of tracks.
- 1 171 chambers with total 354 240 tubes (3 cm diameter, 0.85-6.5 m long).
 - Tube resolution 80 μ m
- Cathode Strip Chambers**
- Measure precision coordinates at ends of detector.
- 70 000 channels
 - Resolution 60 μ m

The ATLAS Experiment – particle identification

