

# Fyzikální laboratoř CERN

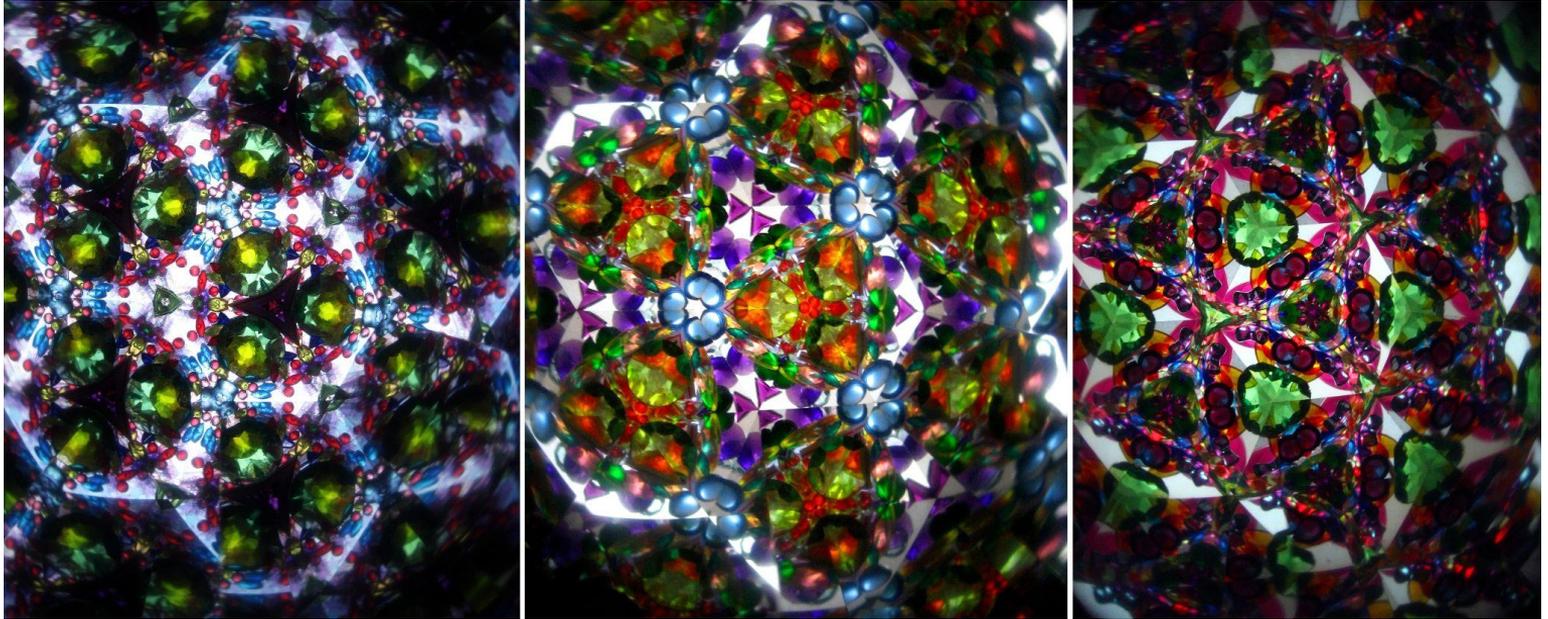


**Fyzikální Καλεϊδοσκοπ 2024**  
**Mgr. Jiří Kvita, Ph.D.**  
**Společná laboratoř optiky**  
**UP a FZÚ AV ČR**



# Etymology [ [edit](#) ]

Coined by its Scottish inventor [David Brewster](#),<sup>[1]</sup> "kaleidoscope" is derived from the [Ancient Greek](#) word [καλός](#) (*kalos*), "beautiful, beauty",<sup>[2]</sup> [εἶδος](#) (*eidos*), "that which is seen: form, shape"<sup>[3]</sup> and [σκοπέω](#) (*skopeō*), "to look to, to examine",<sup>[4]</sup> hence "observation of beautiful forms."<sup>[5]</sup> It was first published in the patent that was granted on July 10, 1817.<sup>[6]</sup>



<https://upload.wikimedia.org/wikipedia/commons/f/f4/Kaleidoscopes.jpg>

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# Science for peace

CERN's convention states: "The Organization shall have no concern with work for military requirements and the results of its experimental and theoretical work shall be published or otherwise made generally available."

Conseil européen pour la Recherche Nucléaire  
European Council for Nuclear Research



YEARS / ANS CERN

1954-2024

# CERN

United Nations / Founded



October 24, 1945, San Francisco, California, United States

NASA / Founded



July 29, 1958, United States

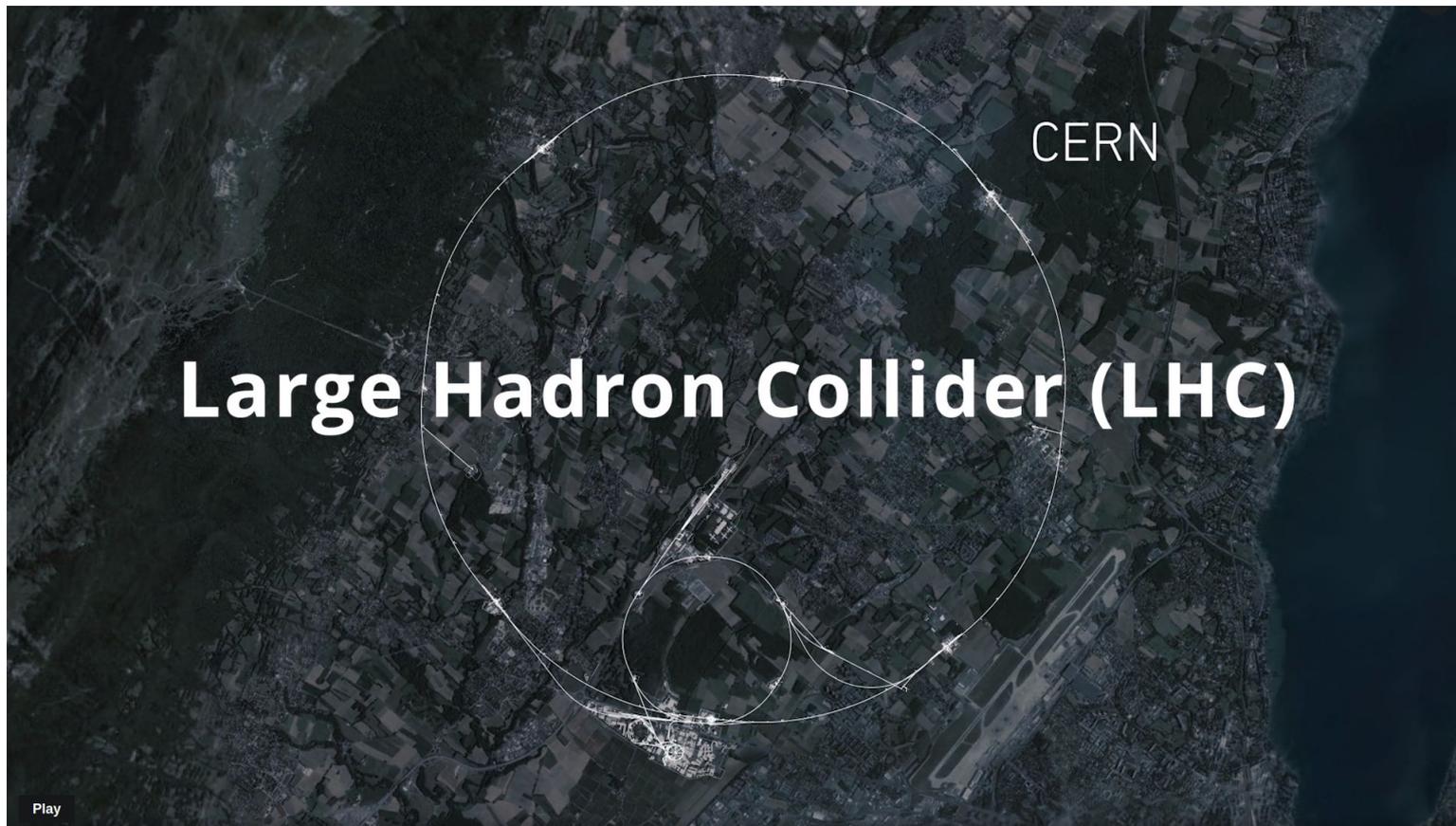


CERN

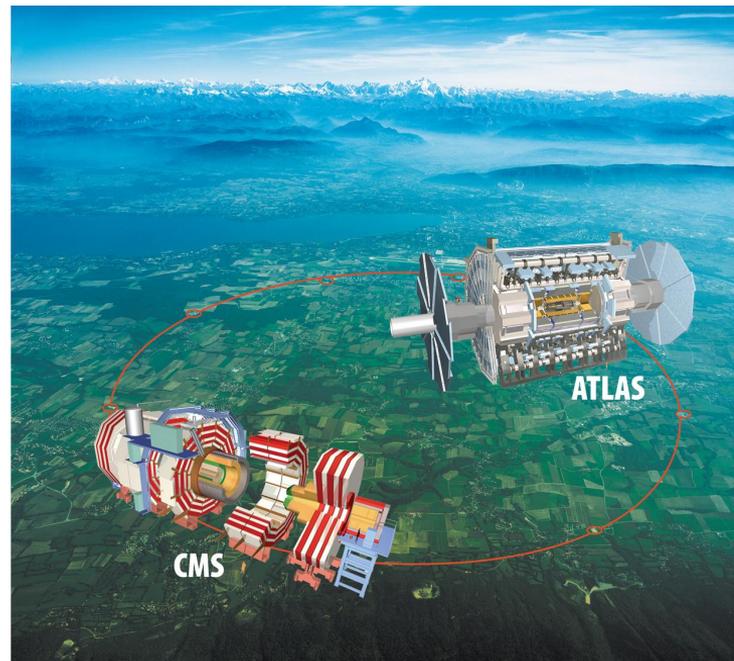
23 členských zemí + pozorovatelské státy



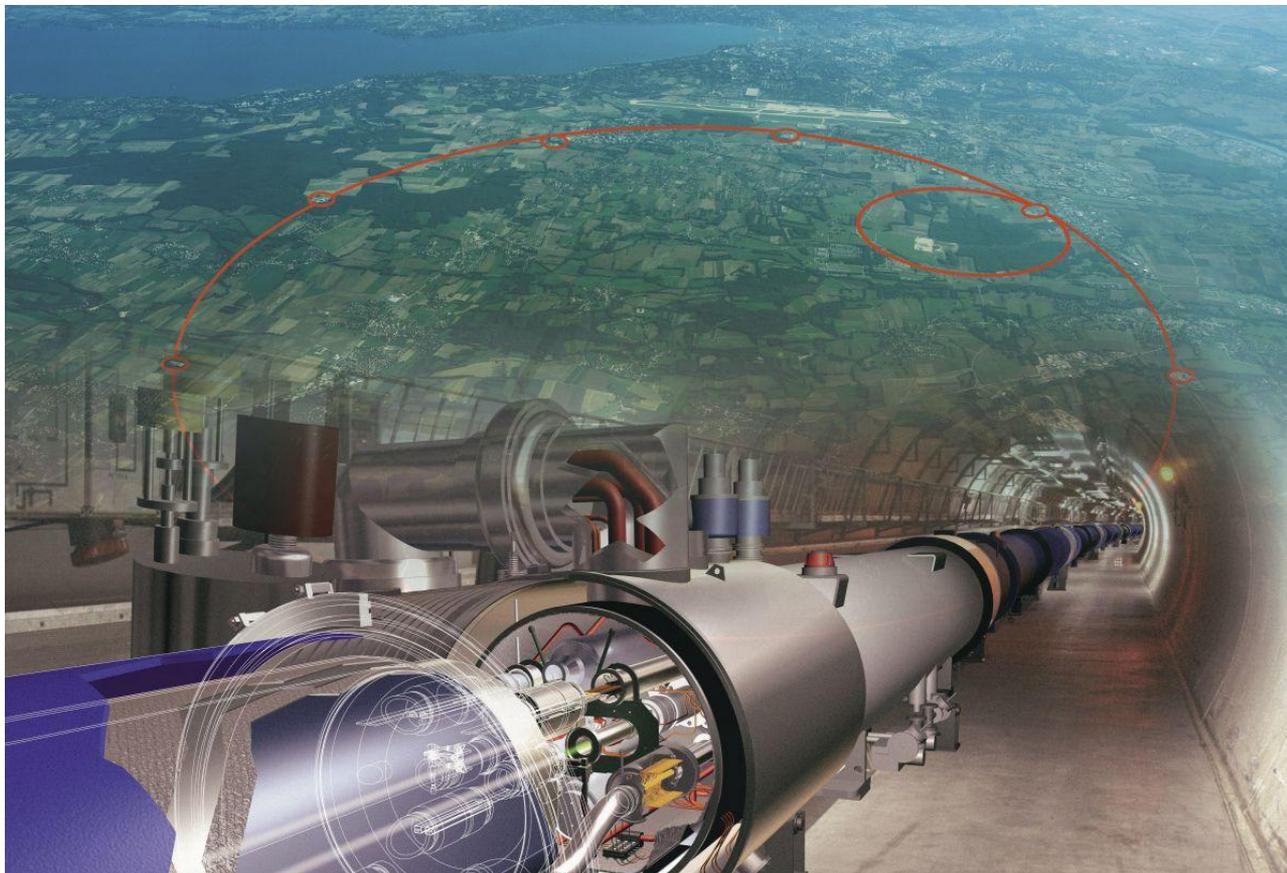
Domov největšího a nejenergetičtějšího urychlovače na světě  
– Large Hadron Collider, LHC



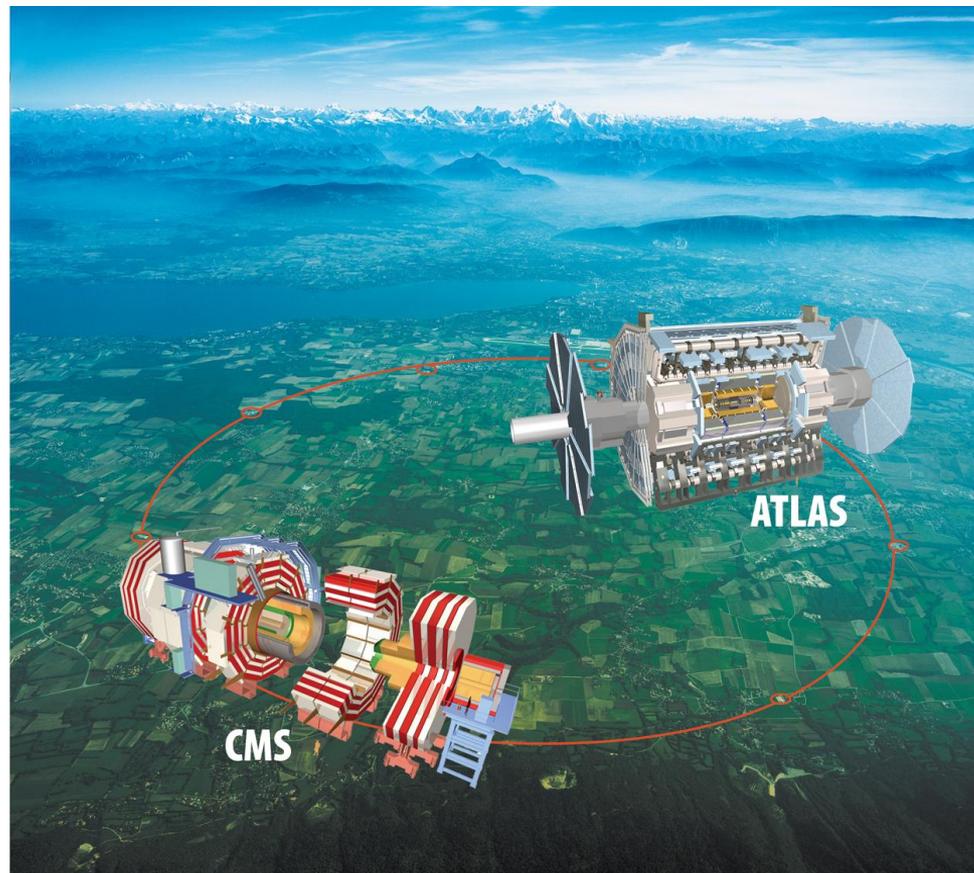
## Domov největšího a nejenergetičtějšího urychlovače na světě – Large Hadron Collider, LHC



# Domov největšího a nejenergetičtějšího urychlovače na světě – Large Hadron Collider, LHC



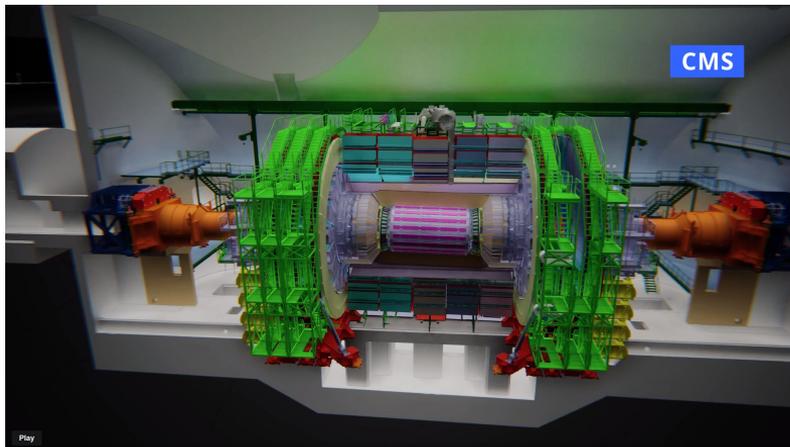
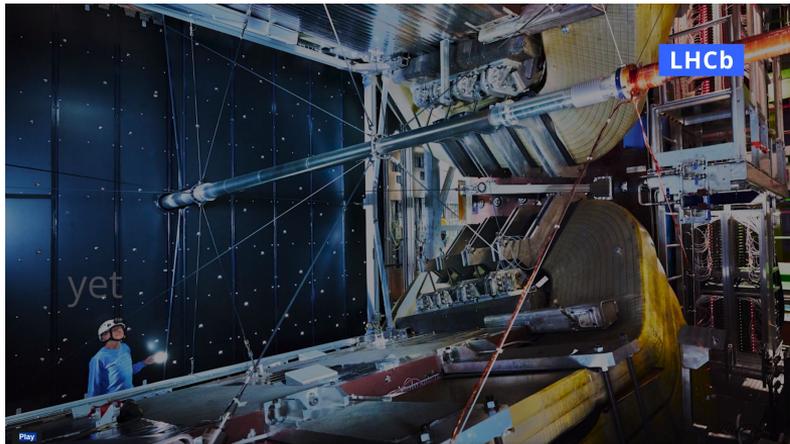
# CERN – Hlavní LHC Experimenty

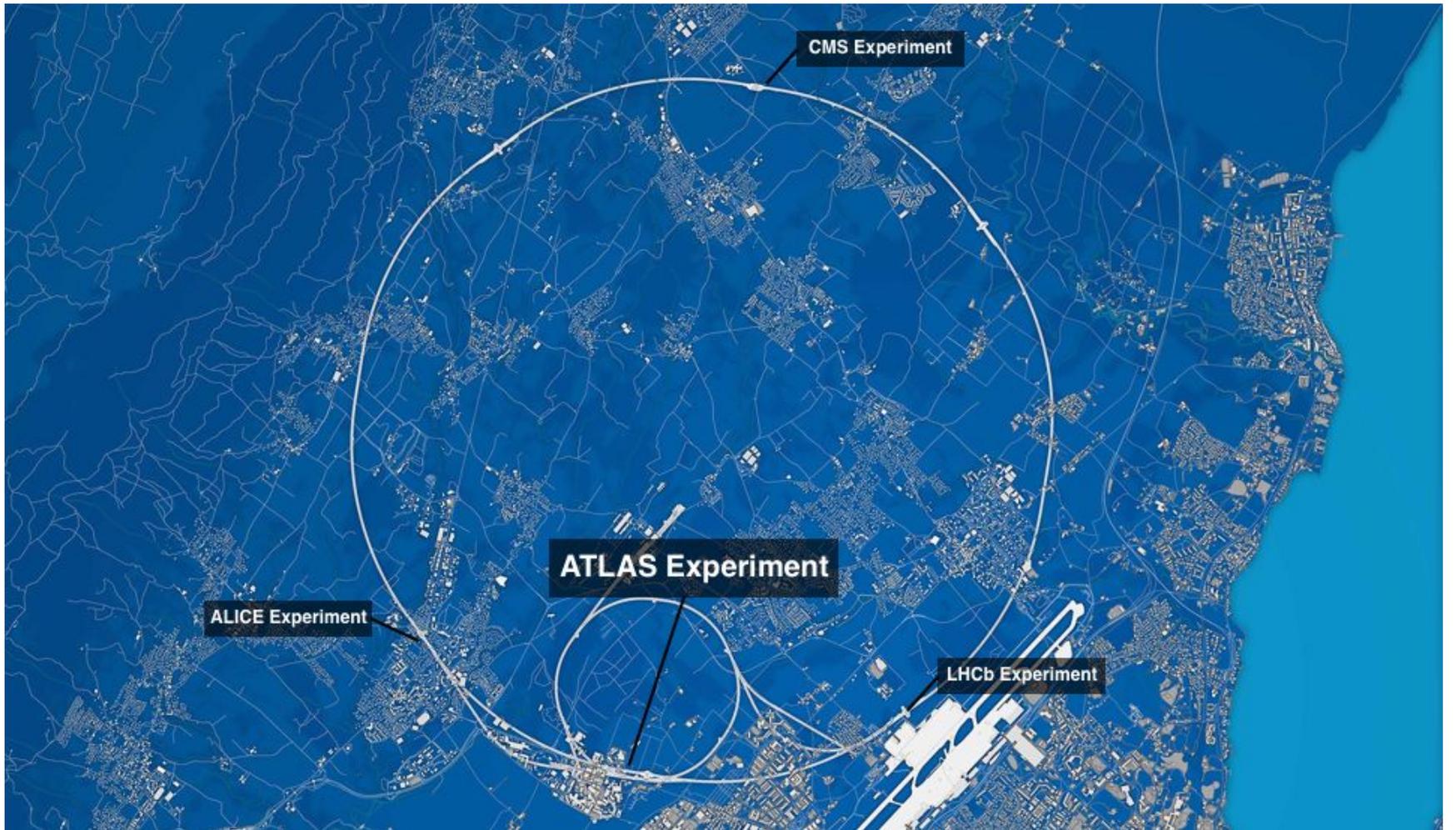


# CERN – Hlavní LHC Experimenty



# CERN – Hlavní LHC Experimenty





# \$CERN\$



WHERE THE  
**WEB**  
WAS BORN

In the offices of this corridor, all the fundamental technologies of the World Wide Web were developed.

Started in 1990 from a proposal made by Tim Berners-Lee in 1989, the effort was first divided between an office in building 31 of the Computing and Networking Division (CN) and one in building 2 of the Electronics and Computing for Physics Division (ECP).

In 1991 the team came together in these offices, then belonging to ECP. It was composed of two CERN staff members, Tim Berners-Lee (GB) and Robert Cailliau (BE), aided by a number of Fellows, Technical Students, a Coopérant and Summer Students.

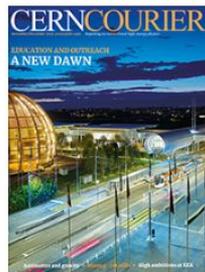
At the end of 1994 Tim Berners-Lee left CERN to direct the WWW Consortium (W3C), a world-wide organization devoted to leading the Web to its full potential. The W3C was founded with the help of CERN, the European Commission, the Massachusetts Institute of Technology (MIT), the Institut National pour la Recherche en Informatique et en Automatique (INRIA), and the Advanced Research Projects Agency (ARPA).

In 1995 Tim Berners-Lee and Robert Cailliau received the ACM Software System Award for the World Wide Web. In 2004, Tim Berners-Lee was awarded the first Millennium Technology Prize by the Finnish Technology Award Foundation.

*The CERN Library  
June 2004*



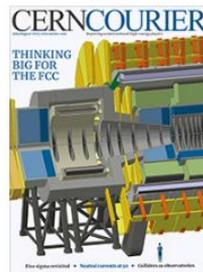
Jan/Feb 2024



Nov/Dec 2023



Sep/Oct 2023



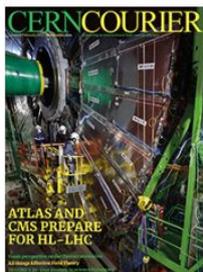
Jul/Aug 2023



May/June 2023



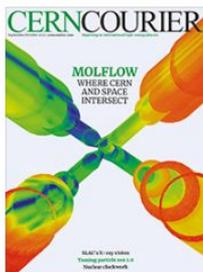
Mar/Apr 2023



Jan/Feb 2023



Nov/Dec 2022



Sep/Oct 2022



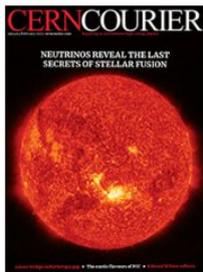
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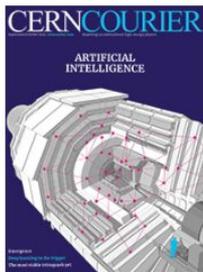
Mar/Apr 2022



Jan/Feb 2022



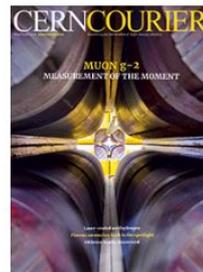
Nov/Dec 2021



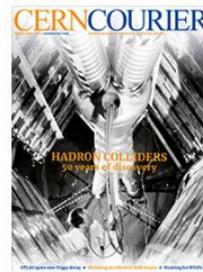
Sep/Oct 2021



Jul/Aug 2021



May/June 2021



Mar/Apr 2021



## Sobota 7.10.2023

 **CERN** •  
5 hod · 🌐

Today, CERN inaugurates CERN Science Gateway, its new emblematic centre for science education and outreach. It will be open to the public from tomorrow, 8 October 2023.

Join us in celebration and watch the inauguration ceremony live today at 11.00 CEST: <https://webcast.cern.ch/event/i1332909>

[#CERNScienceGateway](#)

 **CERN** •  
2 h · 🌐

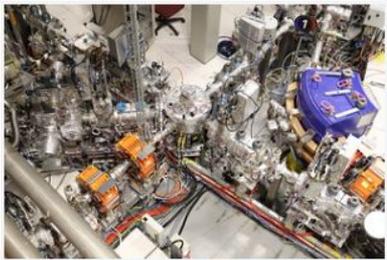
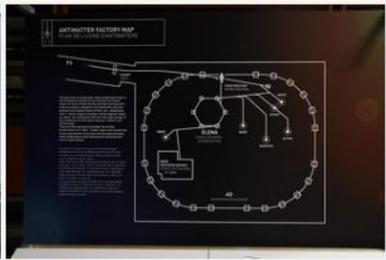
CERN launches Science Gateway

Today, CERN inaugurates its new emblematic centre for science education and outreach, CERN Science Gateway. The building was designed by architect Renzo Piano and funded through external donations, with the leading contribution coming from Stellantis.

CERN Science Gateway will be open to the public as of tomorrow, 8 October 2023.

Find out more: <https://home.cern/.../cern-inaugurates-science-gateway...>

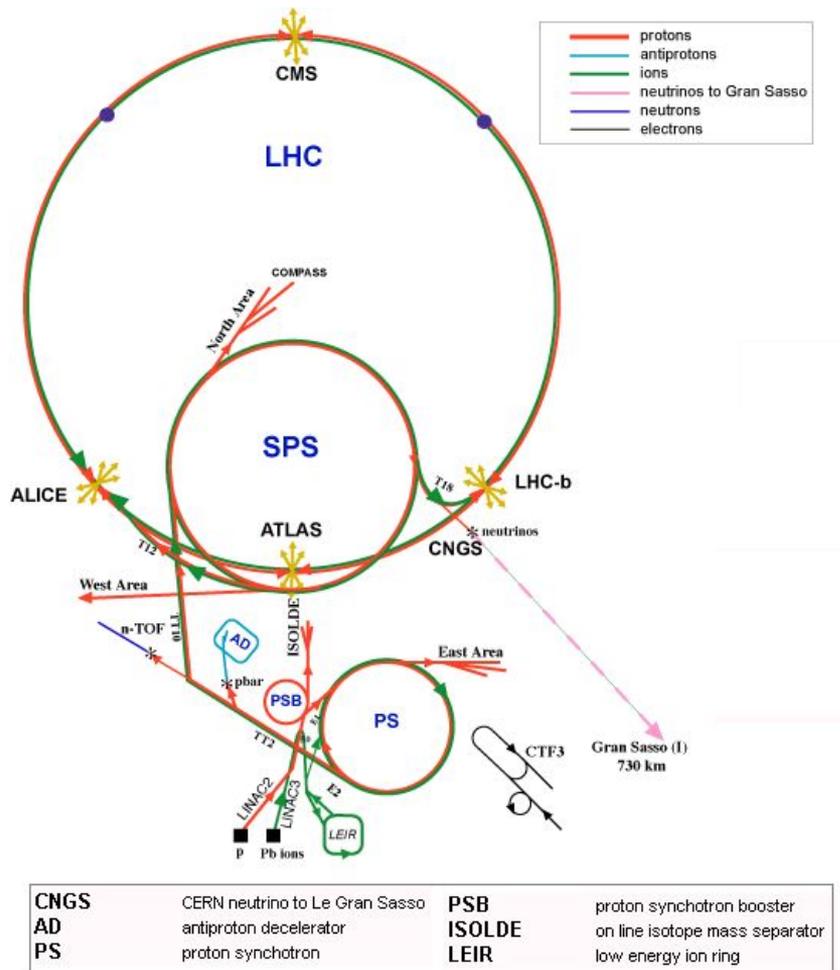
[#CERNScienceGateway](#) [Stellantis](#)



# LHC

- Obvod 26.66km
- Urychluje protony na energii 6.8 TeV (design 7 TeV)
  - silně relativistické, rychlosti blízké rychlosti světla
  - celk. těžišťová energie srážek: 13.6 TeV (design 14 TeV)
- Dipólových magnetů: 1232
  - délka 14.3m, 35t, 2-in-1 design, 0.5M CHF.
  - $B=8.33T$ ,  $R=2804m$
  - Teplota: 1.9 K, proud: 12 kA
- Všech magnetů: 9.6k
- Cena: 4332 MCHF

<https://cds.cern.ch/record/2262862/files/fermilab-conf-15-635-td.pdf>  
<https://home.cern/resources/faqs/facts-and-figures-about-lhc>

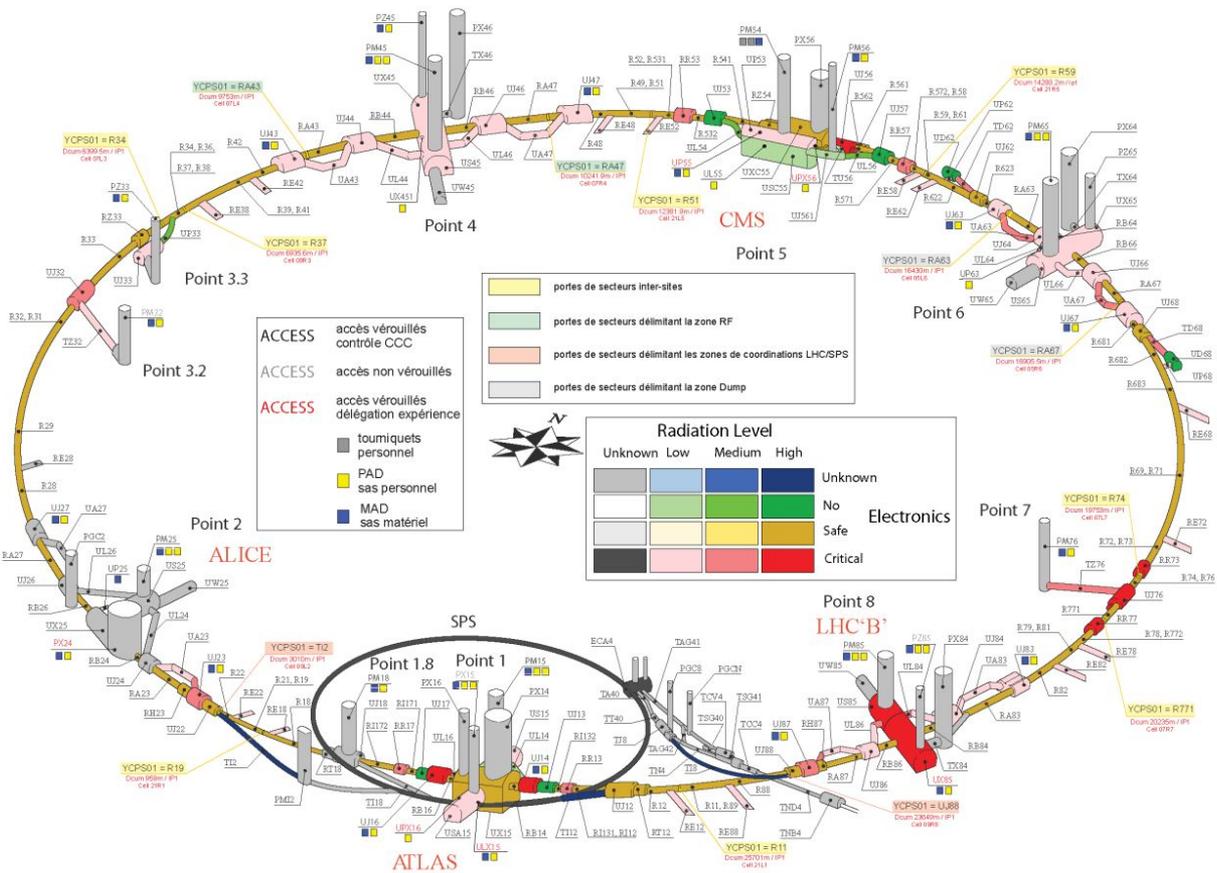


- Fixed target areas

- North Area (French sector, high energy, 450 GeV primary, 120 GeV secondary beams)
- East Area (Swiss, low energy PS Booster 24 GeV primary, ~1–8 GeV secondary beams)

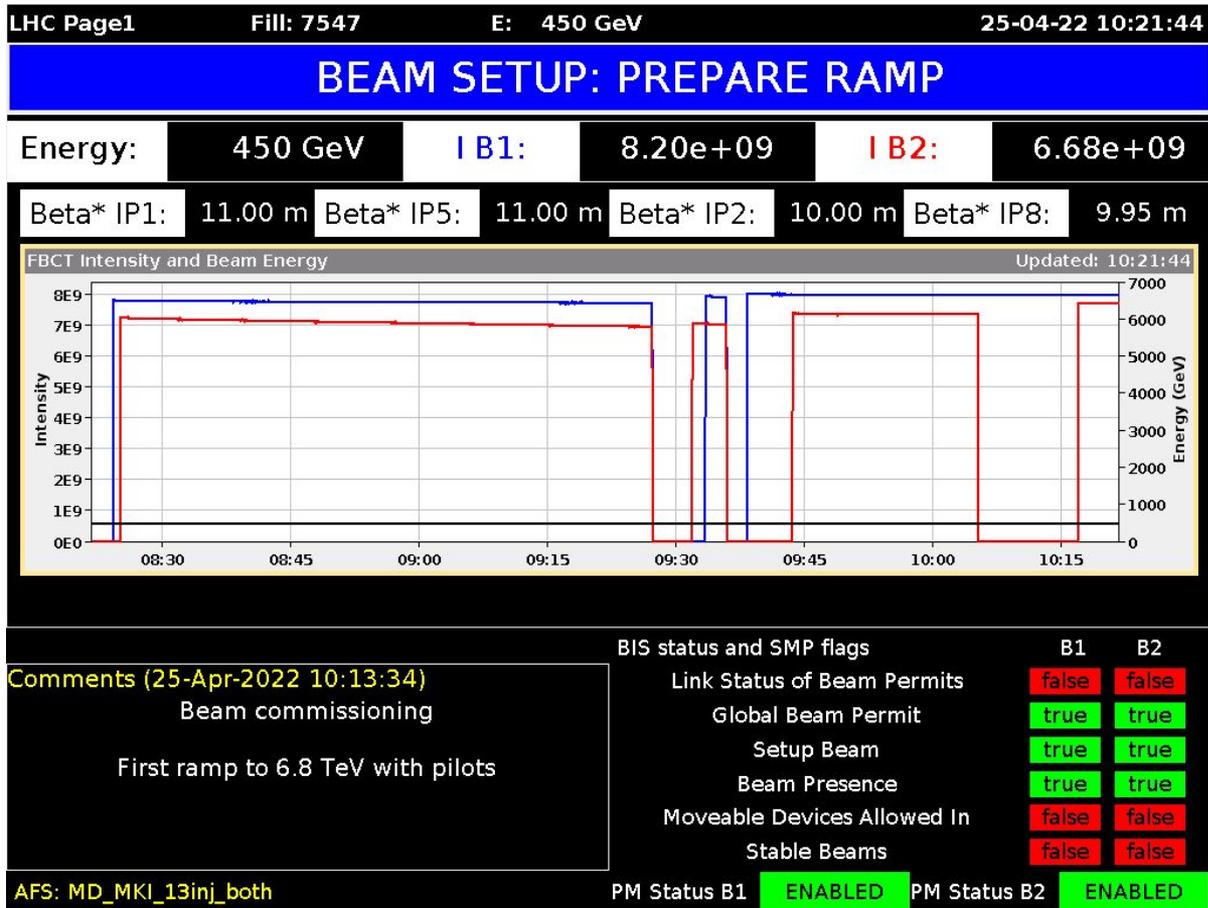
- Neutrino beams to Gran Sasso (IT)
- (not just) pre-accelerators for the highest-energy LHC.

# The LHC



# The LHC

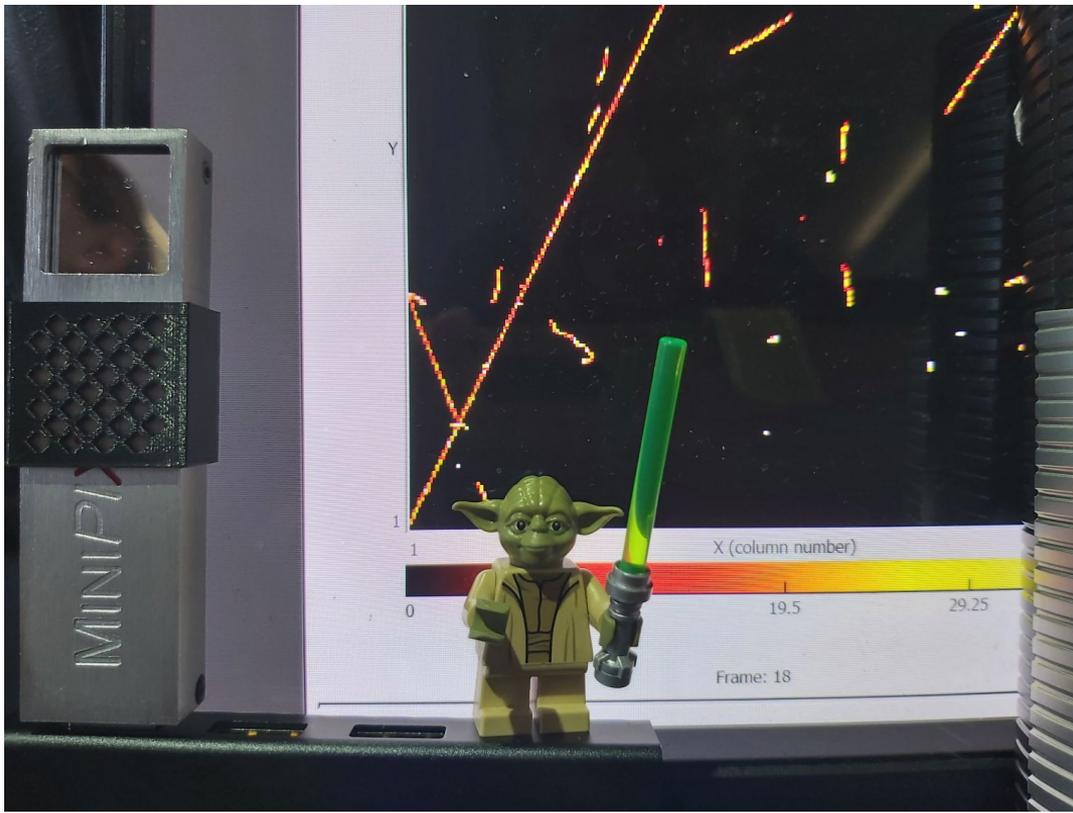
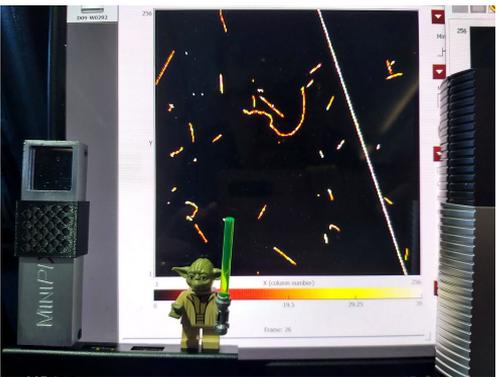
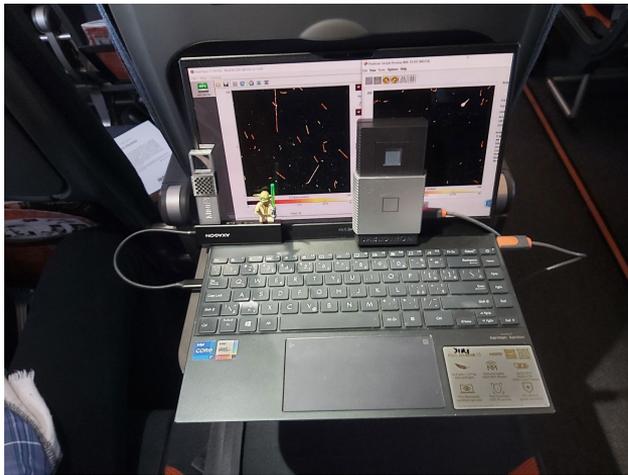




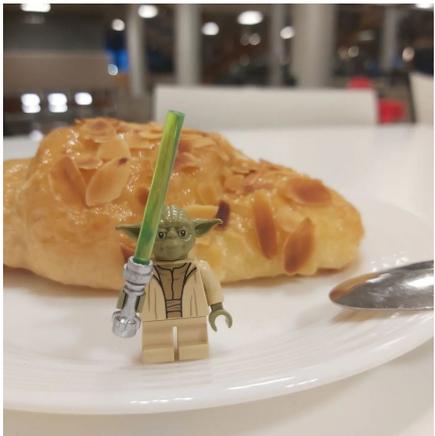
# Going to CERN



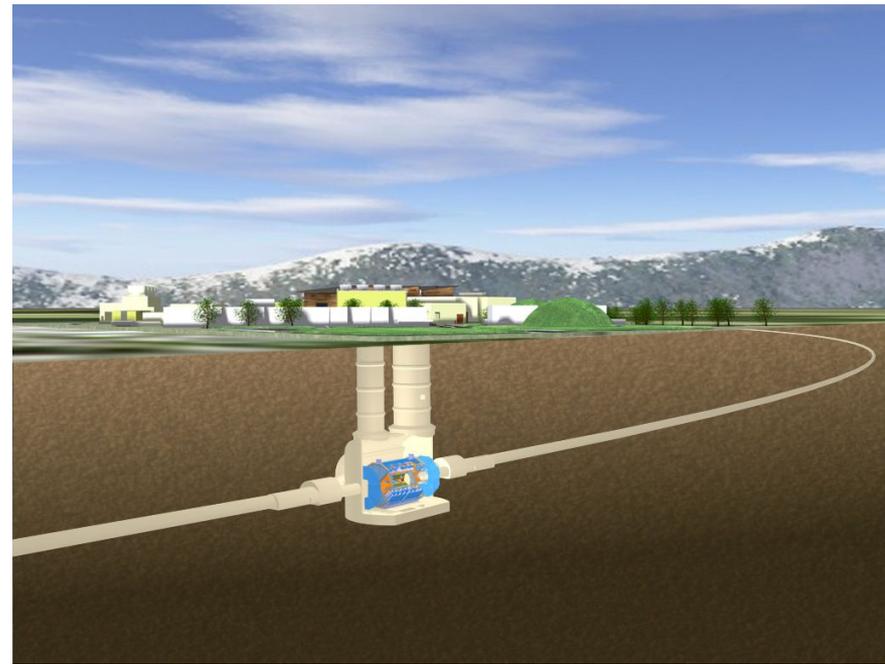
# Flying to CERN – with particle cameras!



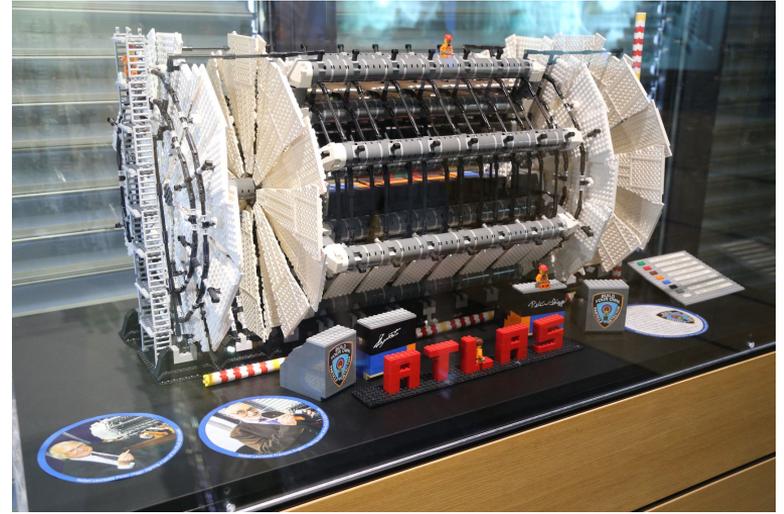
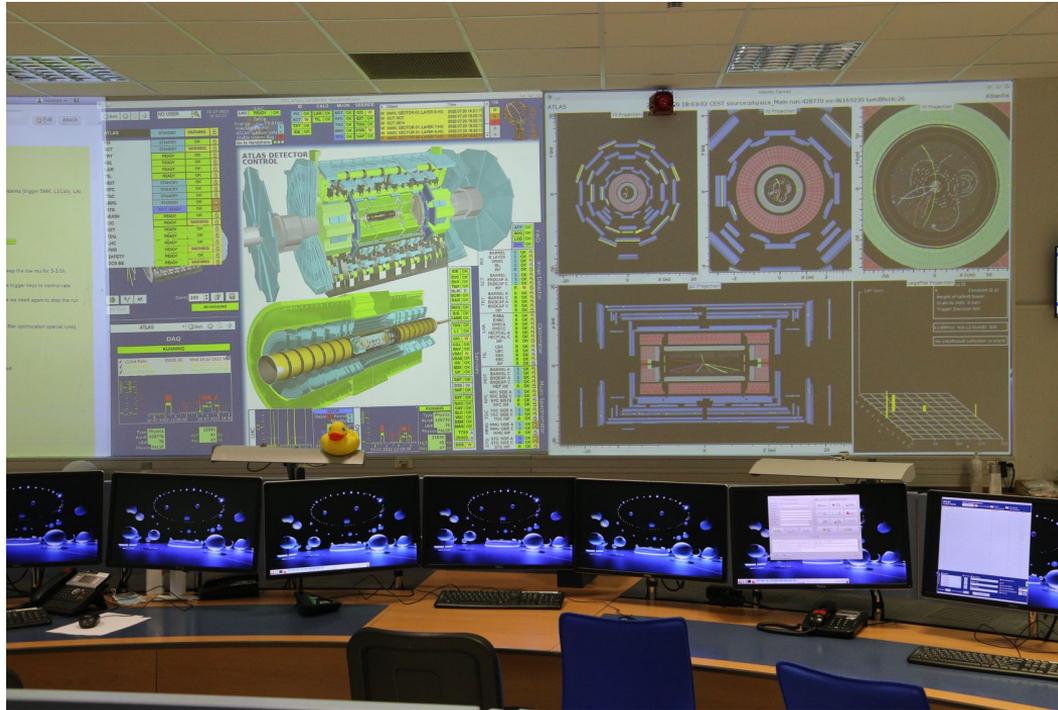
# Excursion Day



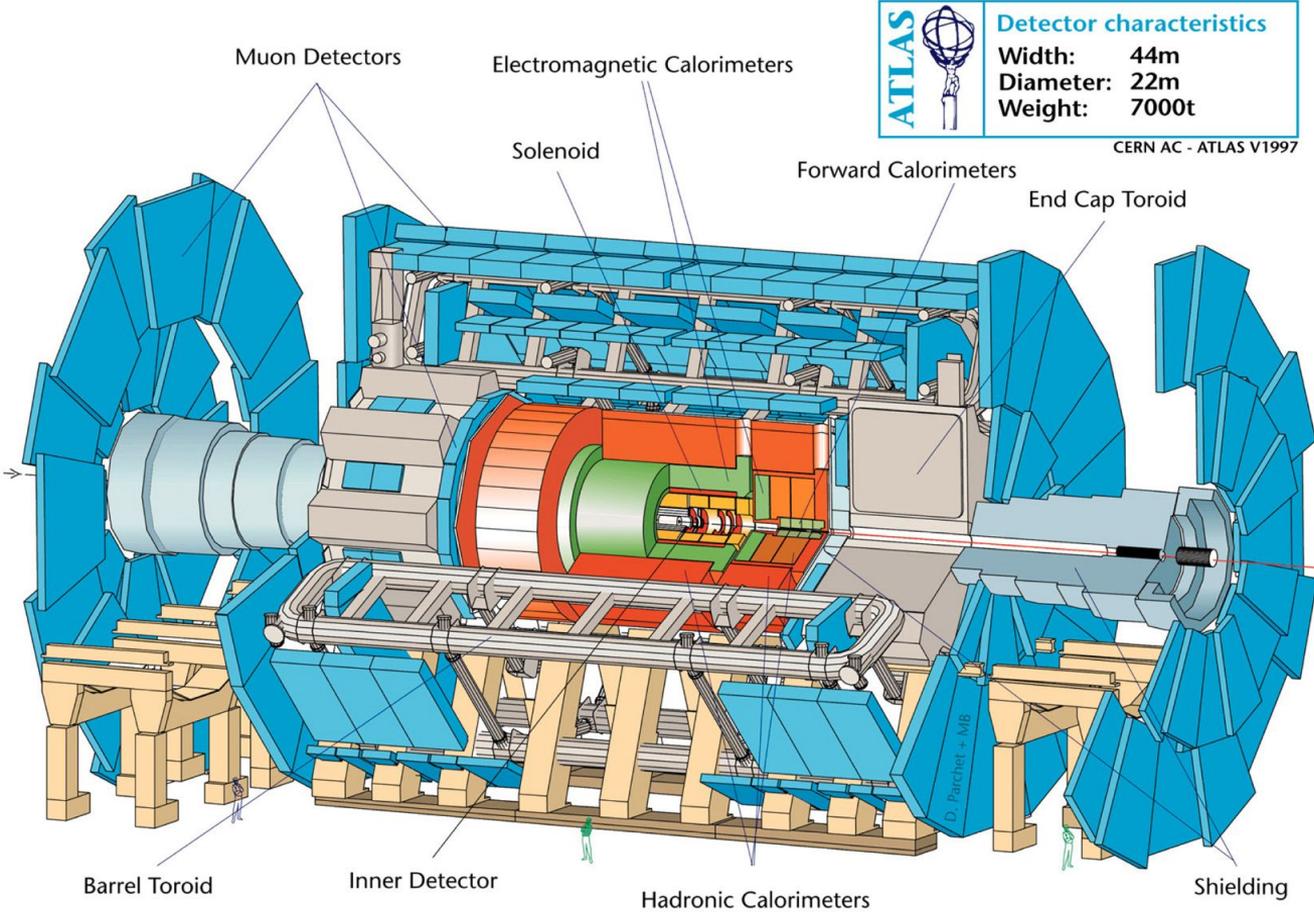
# The ATLAS Experiment



# The ATLAS Experiment



# The ATLAS Experiment



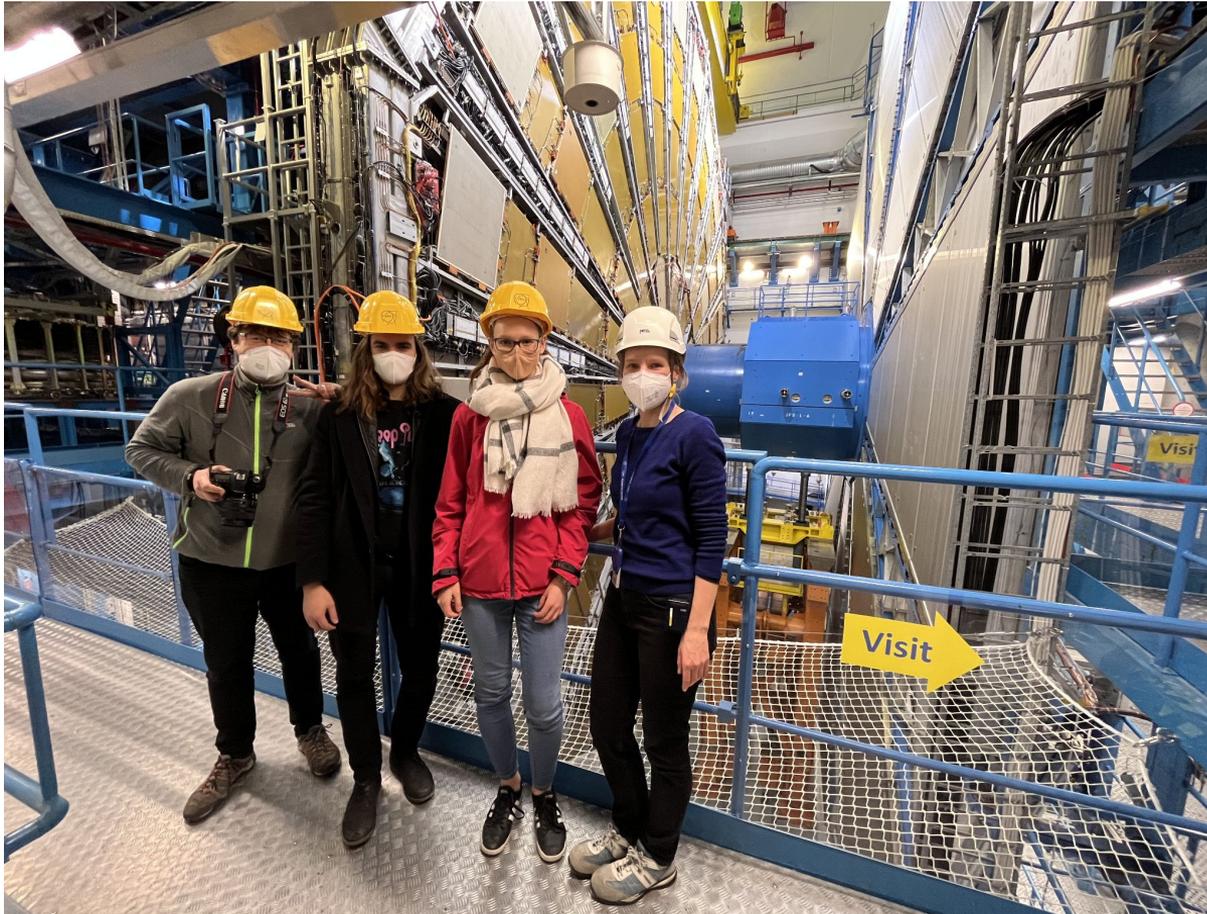
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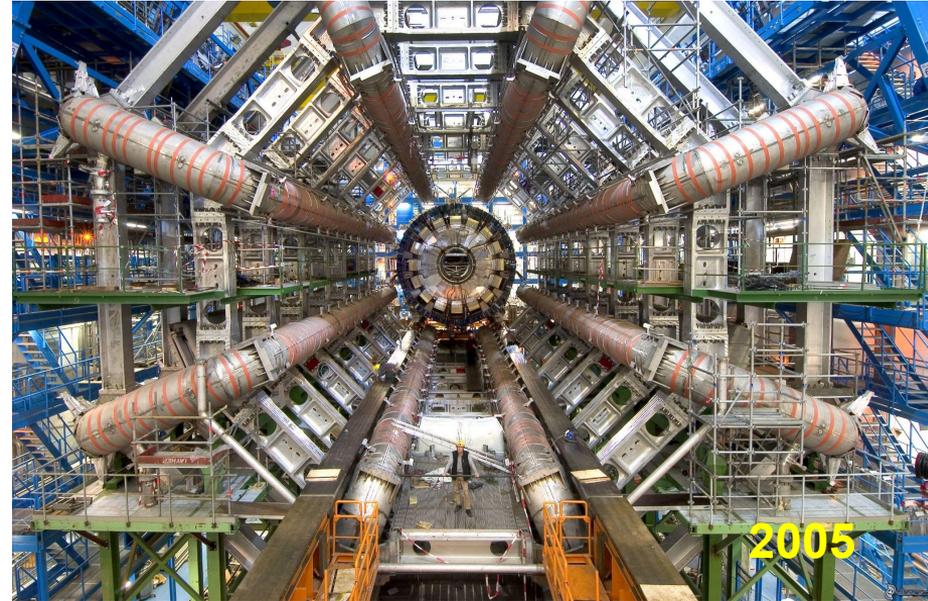
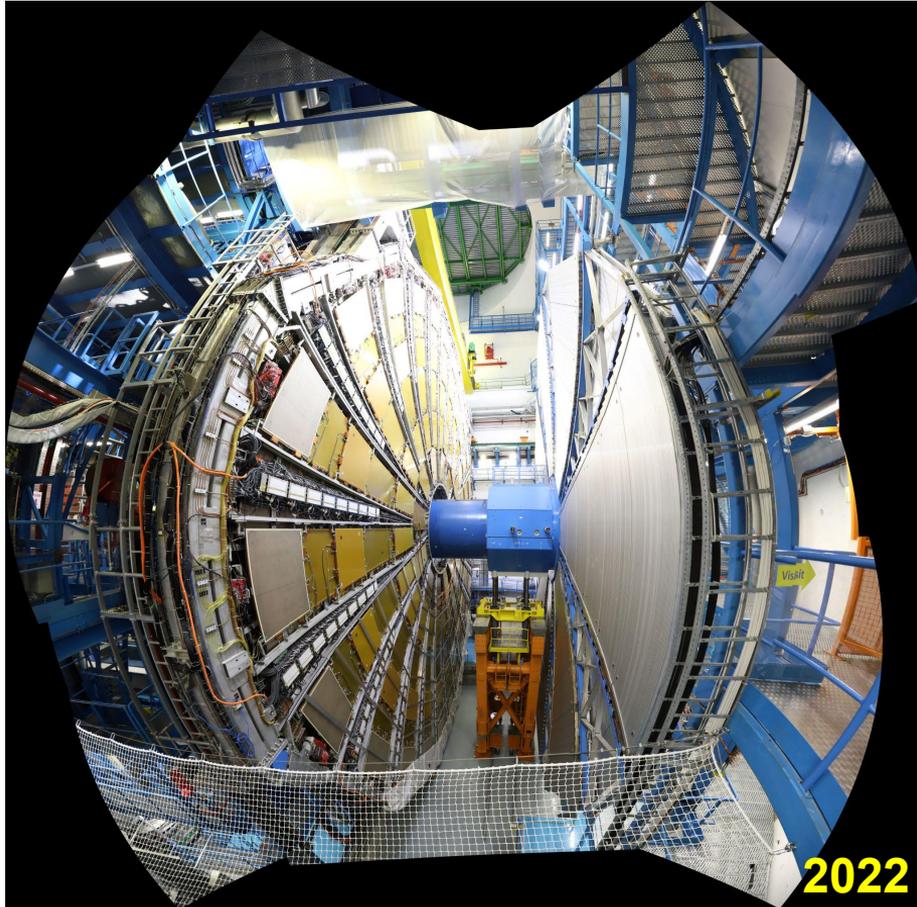
# The ATLAS Experiment



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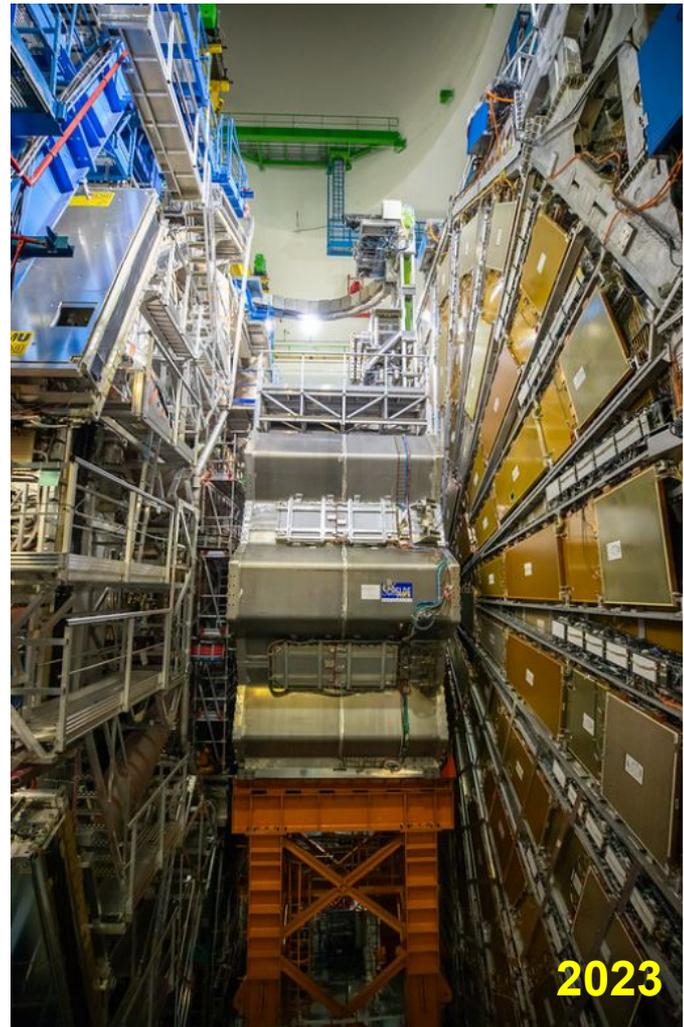


# The ATLAS Experiment

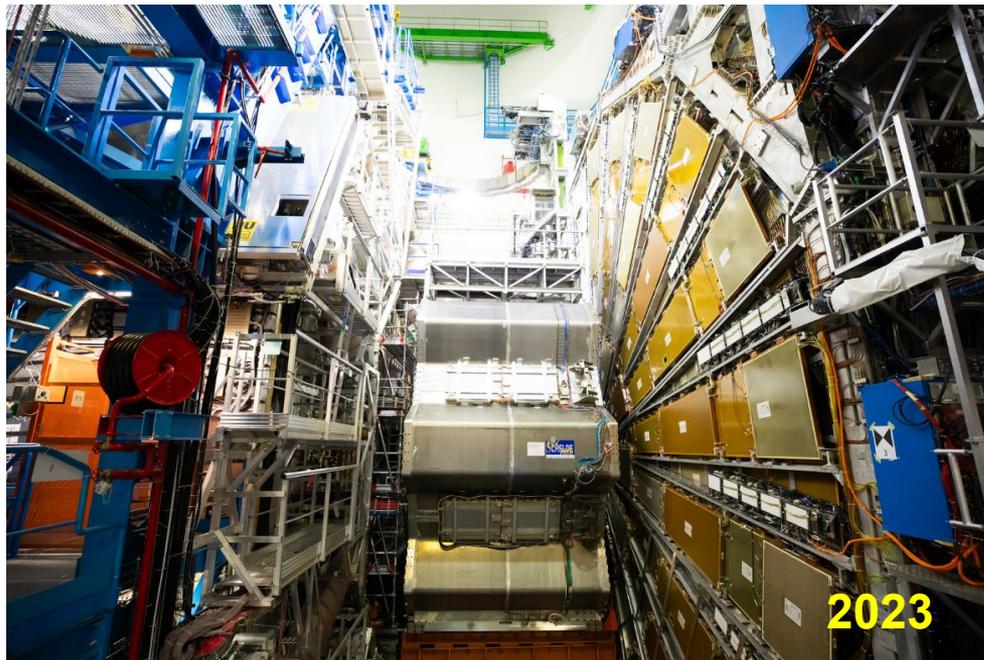


<https://timeline.web.cern.ch/timeline-header/143>

# The ATLAS Experiment



# The ATLAS Experiment



SWI swissinfo.ch

Swiss perspectives in 10 languages



▲ French President Emmanuel Macron, left, with Fabiola Gianotti, Director General of CERN, and Swiss President Alain Berset in the ATLAS experiment on Thursday © Keystone / Martial Trezzini



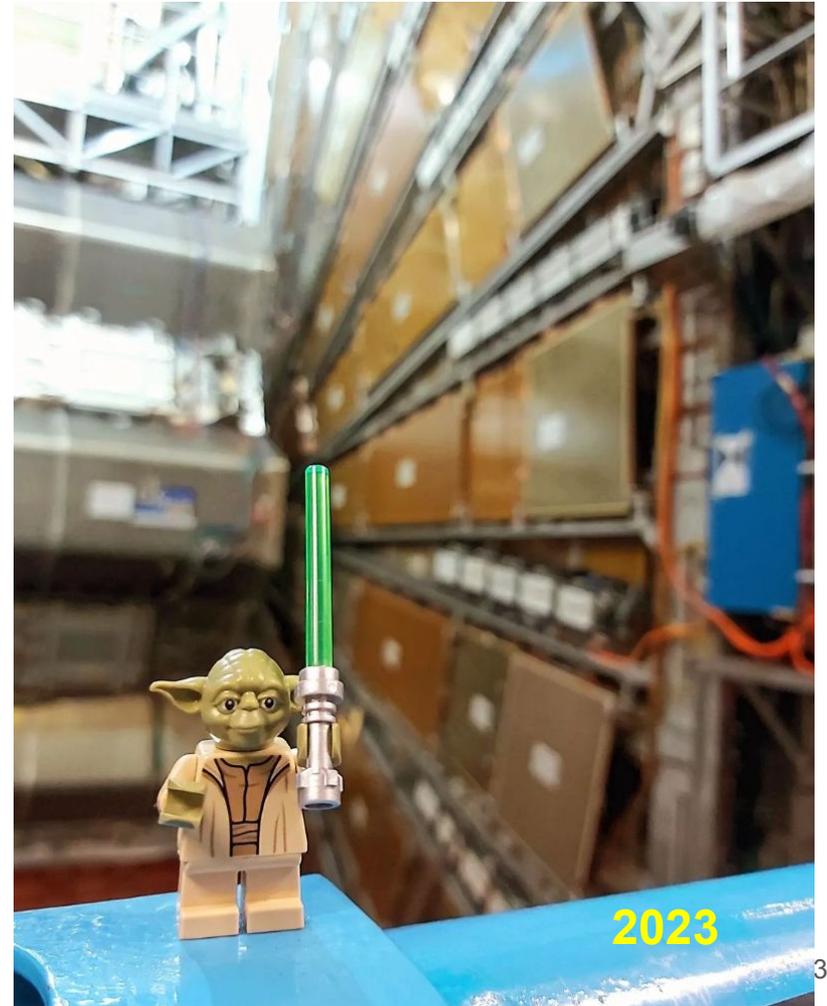
read aloud

French President Emmanuel Macron spoke on Thursday of France's ambitions to remain in "first place" during a visit to the world's most powerful particle accelerator on the Franco-Swiss border, where a successor, even more powerful, is being studied.

November 17, 2023 - 10:33

🕒 2 minutes

# The ATLAS Experiment



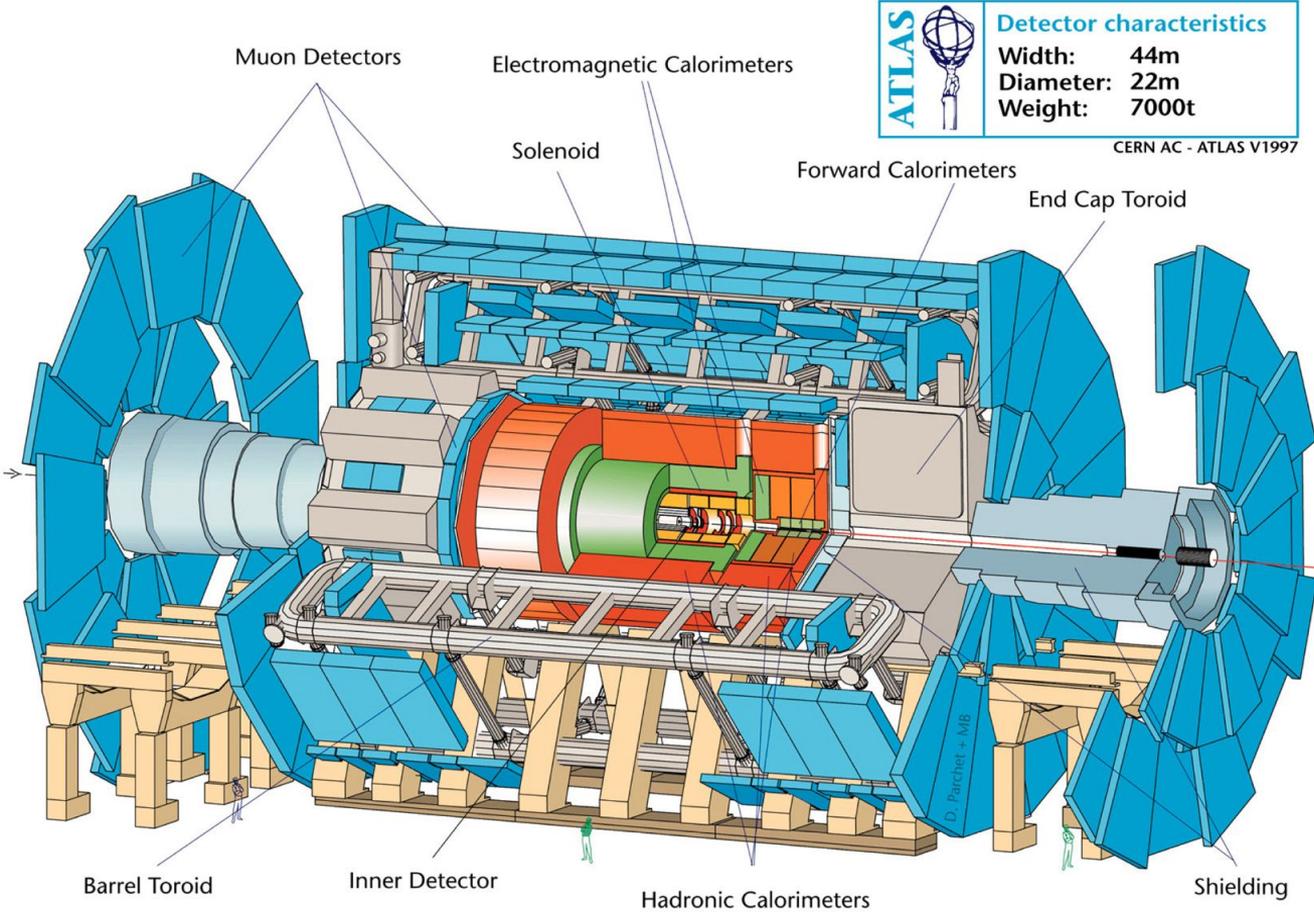
# The ATLAS Experiment



# The ATLAS Experiment



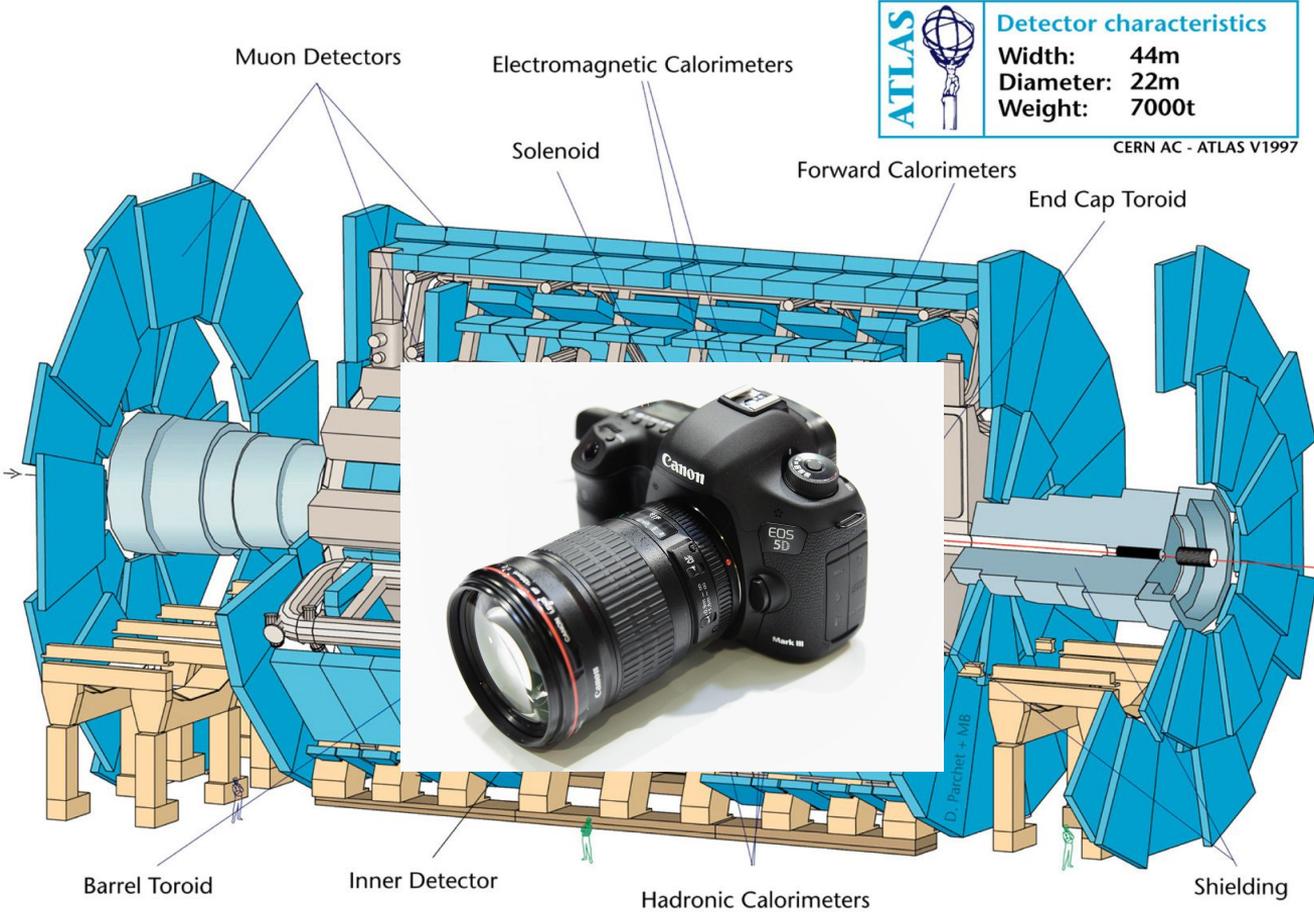
# The ATLAS Experiment



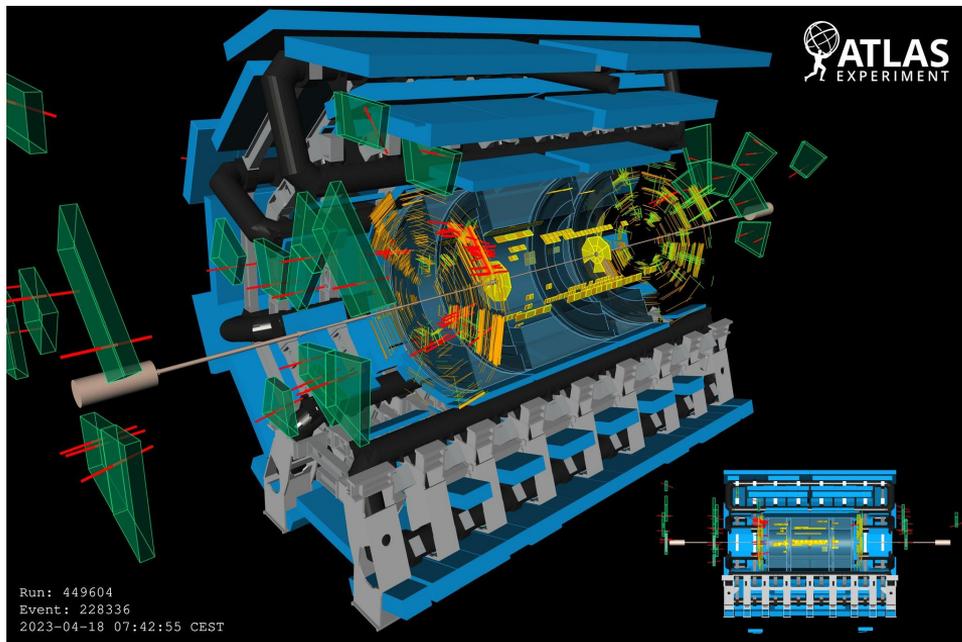
	<b>Detector characteristics</b>	
	<b>Width:</b>	<b>44m</b>
	<b>Diameter:</b>	<b>22m</b>
	<b>Weight:</b>	<b>7000t</b>

CERN AC - ATLAS V1997

# The ATLAS Experiment



# The ATLAS Experiment

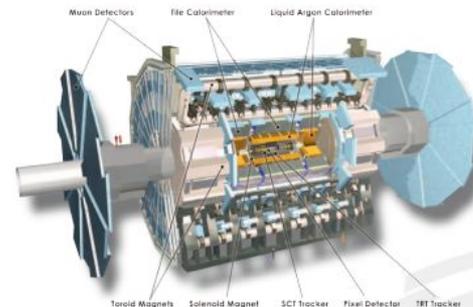


## ATLAS Fact Sheet

1

### The ATLAS Detector

- Diameter 25 m -- Length : 46 m
- Barrel toroid length 26 m
- Overall weight 7 000 tonnes
- ~ 100 million electronic channels
- ~ 3 000 km of cables



### Calorimeters

Measure the energies carried by the particles

#### Liquid Argon (LAr) Calorimeter

- Barrel 6.4 m long, 53 cm thick, 110 000 channels.
- Works with Liquid Argon at  $-183^{\circ}\text{C}$
- LAr endcap consists of the forward calorimeter, electromagnetic (EM) and hadronic endcaps.
- EM endcaps each have thickness 0.632 m and radius 2.077 m.
- Hadronic endcaps consist of two wheels of thickness 0.8 m and 1.0 m with radius 2.09 m.
- Forward calorimeter has three modules of radius 0.455 m and thickness 0.450 m each.

#### Tile Calorimeter (TileCal)

- Barrel made of 64 wedges, each 5.6 m long and 20 tonnes.
- Each Endcap has 64 wedges, each 2.6 m long.
- 500 000 plastic scintillator tiles.

### Muon System

Identifies and measures the momenta of muons

#### Thin Gap Chambers

- For triggering and 2nd coordinate measurement (non-bending direction) at ends of detector.
- 440 000 channels

#### Resistive Plate Chambers

- For triggering and 2nd coordinate measurement in central region.
- 380 000 channels
  - Electric Field 5 000 V/mm

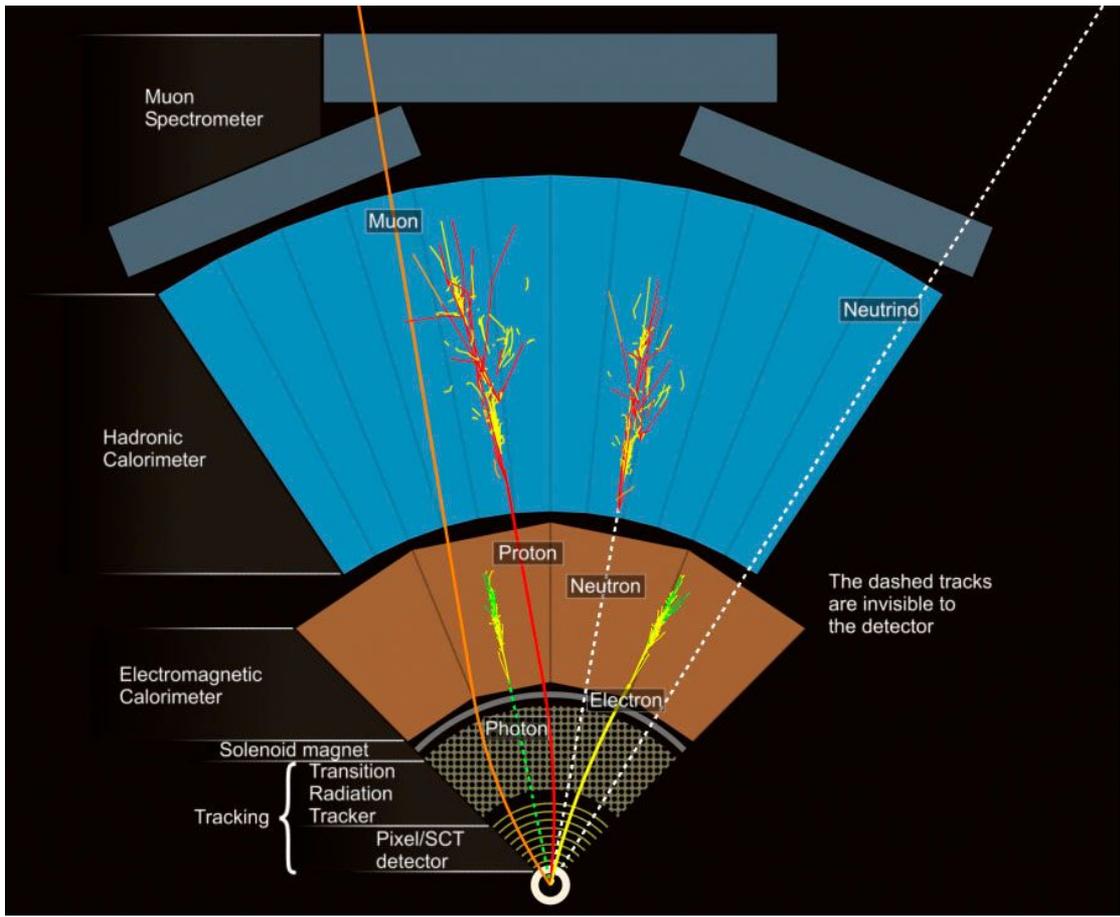
#### Monitored Drift Tubes

- Measure curves of tracks.
- 1 171 chambers with total 354 240 tubes (3 cm diameter, 0.85-6.5 m long).
  - Tube resolution 80  $\mu\text{m}$

#### Cathode Strip Chambers

- Measure precision coordinates at ends of detector.
- 70 000 channels
  - Resolution 60  $\mu\text{m}$

# The ATLAS Experiment – particle identification



# The ATLAS Experiment

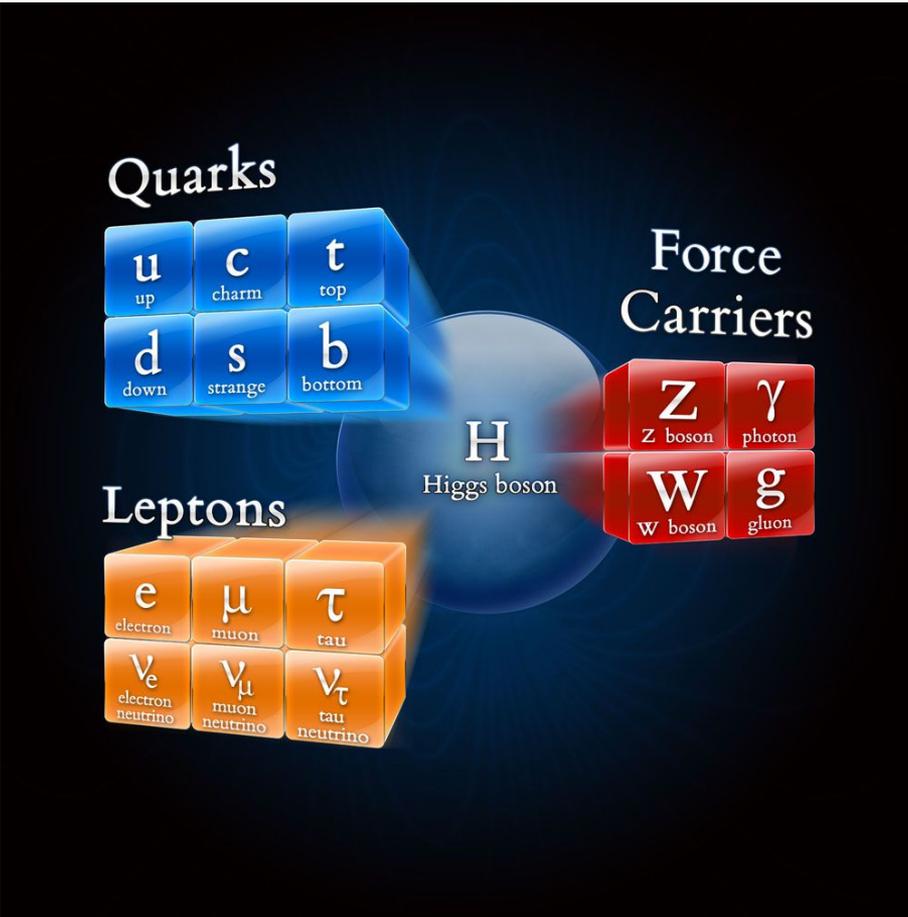
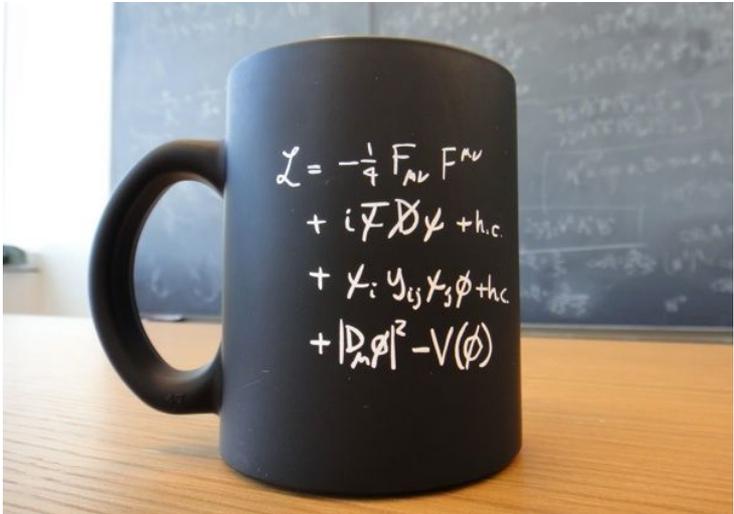


<https://atlas.cern/Discover/Collaboration>

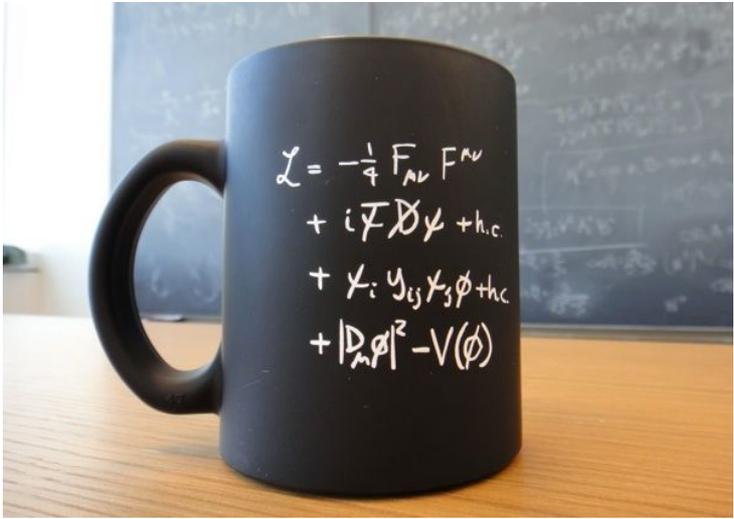


- 6000 členů
- 3000 vědeckých autorů
- 182 institucí z 42 zemí
- 1200 doktorských studentů

# Quarks, Leptons, Gauge Bosons and the BEH boson



# Quarks, Leptons, Gauge Bosons and the BEH boson



## WHAT PART OF

$$-\frac{1}{2} \partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_a f^{abc} \partial_\mu g_\nu^b g_\nu^c g_\mu^a - \frac{1}{2} g_a^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \frac{1}{2} i g_a^2 (\bar{\psi} \gamma^\mu \psi) g_\mu^a$$

$$\bar{C}^a \partial^\mu C^a + g_a f^{abc} \partial_\mu C^a g_\mu^b g_\mu^c - \partial_\mu W_\mu^+ \partial_\nu W_\nu^- - M^2 W_\mu^+ W_\mu^- - \frac{1}{2} \partial_\nu Z_\nu^0 Z_\nu^0 - \frac{1}{2} M^2 Z_\nu^0 Z_\nu^0 -$$

$$\frac{1}{2} \partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2} \partial_\mu H \partial_\mu H - \frac{1}{2} m_H^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2} \partial_\mu \phi^0 \partial_\mu \phi^0$$

$$+ \frac{1}{2} M^2 \phi^0 \phi^0 - \beta_h [2\lambda \phi^4 + 2\lambda H^4 + \frac{1}{2} (H^2 + \phi^0)^2 + 2\phi^+ \phi^-] + 2\lambda \phi^0 - \alpha_h - i g_{cW} [\partial_\mu Z_\nu^0 (W_\mu^+ W_\nu^- -$$

$$W_\mu^- W_\nu^+) - Z_\nu^0 (\partial_\mu \partial_\nu W_\mu^+ - W_\mu^- \partial_\nu W_\mu^+) + Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\mu W_\mu^+)] - i g_{sW} \partial_\nu A_\mu (W_\mu^+ W_\nu^- -$$

$$W_\mu^- W_\nu^+) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\mu W_\mu^+) + A_\nu (W_\mu^- \partial_\nu W_\mu^+ - W_\nu^+ \partial_\mu W_\mu^-)] - \frac{1}{2} g^2 W_\mu^+ W_\nu^- W_\nu^+ W_\mu^- +$$

$$\frac{1}{2} g^2 W_\mu^+ W_\nu^- W_\nu^+ W_\mu^- + g^2 Z_\nu^0 (Z_\nu^0 W_\mu^+ W_\mu^- - Z_\nu^0 W_\mu^+ W_\nu^-) + g^2 Z_\nu^0 (A_\mu W_\mu^+ A_\nu W_\nu^- -$$

$$A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_W c_W A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - W_\nu^- W_\mu^+) - 2 A_\mu Z_\nu^0 W_\mu^+ W_\nu^- - g \alpha [H^2 +$$

$$H \phi^0 \phi^0 + 2 H \phi^+ \phi^-] - \frac{1}{2} g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4 H^2 \phi^+ \phi^- +$$

$$2(\phi^0)^2 H^2] - g M W_\mu^+ W_\nu^- H - \frac{1}{2} g \frac{M}{\Lambda} Z_\nu^0 Z_\nu^0 H - \frac{1}{2} i g [W_\mu^+ (\phi^0 \partial_\nu \phi^- - \phi^- \partial_\nu \phi^0) - W_\nu^- (\phi^0 \partial_\mu \phi^+ -$$

$$\phi^+ \partial_\mu \phi^0)] + \frac{1}{2} i g [W_\mu^+ (H \partial_\nu \phi^- - \phi^- \partial_\nu H) - W_\nu^- (H \partial_\mu \phi^+ - \phi^+ \partial_\mu H)] + \frac{1}{2} g^2 [Z_\nu^0 (H \partial_\mu \phi^0 -$$

$$\phi^0 \partial_\mu H) - i g \frac{M}{\Lambda} Z_\nu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + i g s_W M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - i g \frac{1-2c_W^2}{2c_W} Z_\nu^0 \phi^+ \partial_\nu \phi^- -$$

$$\frac{1}{2} g^2 \frac{1}{2c_W} Z_\nu^0 Z_\nu^0 [H^2 + (\phi^0)^2 + 2(2c_W^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2} g^2 \frac{1}{2c_W} Z_\nu^0 Z_\nu^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) -$$

$$\frac{1}{2} i g^2 \frac{2c_W}{c_W} Z_\nu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2} g^2 s_W A_\mu \partial^\mu (W_\mu^+ \phi^- + W_\mu^- \phi^+) + \frac{1}{2} i g^2 s_W A_\mu H (W_\mu^+ \phi^-$$

$$- W_\mu^- \phi^+) - g^2 \frac{2c_W}{c_W} (2c_W^2 - 1) Z_\nu^0 A_\mu \phi^+ \phi^- - g^2 s_W^2 A_\mu \phi^+ \phi^- - \tilde{e}^3 (\gamma^0 + m_e) e^3 -$$

$$\tilde{\nu} \gamma^0 \nu - \tilde{u}^3 (\gamma^0 + m_u) u^3 - d^3 (\gamma^0 + m_d) d^3 + i g s_W A_\mu [-(e^3 \gamma^\mu e^3) + \frac{2}{3} (\tilde{u}^3 \gamma^\mu u^3) -$$

$$\frac{1}{3} (\tilde{d}^3 \gamma^\mu d^3)] + \frac{2}{3} g \frac{M}{\Lambda} Z_\nu^0 [(\tilde{\nu} \gamma^\mu \nu) + (e^3 \gamma^\mu e^3) + 4(e^3 \gamma^\mu e^3) - (\tilde{u}^3 \gamma^\mu u^3) -$$

$$(1 - \gamma^5) \tilde{u}^3] + (\tilde{d}^3 \gamma^\mu (1 - \frac{1}{3} \gamma^5) d^3)] + \frac{1}{\sqrt{2}} W_\mu^+ [(\nu^3 \gamma^\mu (1 + \gamma^5) e^3) - (\tilde{u}^3 \gamma^\mu (1 +$$

$$\gamma^5) C_{\nu d} d^3)] + \frac{1}{\sqrt{2}} W_\mu^- [(e^3 \gamma^\mu (1 + \gamma^5) \nu) + (\tilde{d}^3 \gamma^\mu (1 + \gamma^5) u^3)] + \frac{1}{\sqrt{2}} M [ - \phi^+ (\tilde{\nu} (1 -$$

$$\gamma^5) e^3) + \phi^- (e^3 (1 + \gamma^5) \nu) ] - \frac{1}{\sqrt{2}} M [H (e^3 e^3) + i \phi^0 (e^3 \gamma^5 e^3)] + \frac{1}{\sqrt{2}} M \phi^+ [ - m_d (\tilde{u}^3 C_{\nu d} (1 -$$

$$\gamma^5) d^3) + m_u (\tilde{d}^3 C_{\nu u} (1 + \gamma^5) u^3) ] + \frac{1}{\sqrt{2}} M \phi^- [ m_u (\tilde{d}^3 C_{\nu u} (1 + \gamma^5) u^3) - m_d (\tilde{u}^3 C_{\nu d} (1 -$$

$$\gamma^5) d^3) ] - \frac{1}{\sqrt{2}} M H (u^3 u^3) - \frac{1}{\sqrt{2}} M H (d^3 d^3) + \frac{1}{\sqrt{2}} M \phi^0 (\tilde{u}^3 \gamma^5 u^3) - \frac{1}{\sqrt{2}} M \phi^0 (\tilde{d}^3 \gamma^5 d^3) +$$

$$X^+ (\partial^\mu - M^2) X^+ + X^- (\partial^\mu - M^2) X^- + X^0 (\partial^\mu - M^2) X^0 + Y \partial^\mu Y + i g_{cW} W_\mu^+ (\partial_\nu X^0 X^\nu -$$

$$\partial_\nu X^+ X^0) + i g_{sW} W_\mu^+ (\partial_\nu \bar{X} X^\nu - \partial_\nu X^+ X^0) + i g_{cW} W_\mu^- (\partial_\nu X^0 X^\nu - \partial_\nu X^- X^0) +$$

$$i g_{sW} W_\mu^- (\partial_\nu X^- X^0 - \partial_\nu \bar{X} X^+) + i g_{cW} Z_\nu^0 (\partial_\mu X^+ X^\mu - \partial_\mu X^- X^\mu) + i g_{sW} A_\mu (\partial_\nu X^+ X^\nu -$$

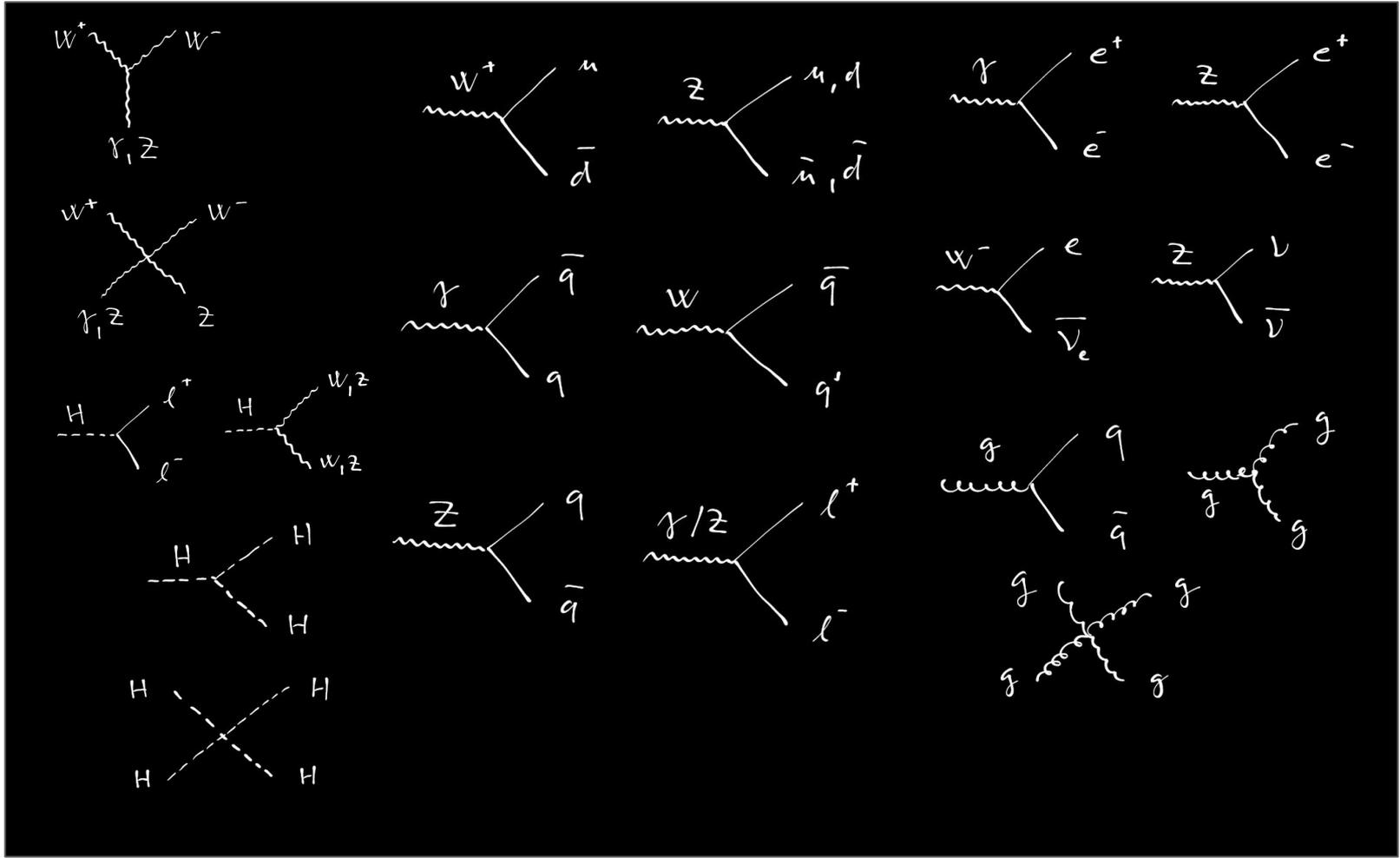
$$\partial_\nu X^- X^\nu) - \frac{1}{2} g M [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{2} \bar{X}^0 X^0 H] + \frac{1-2c_W^2}{2c_W} i g M [\bar{X}^+ X^0 \phi^+ -$$

$$X^- X^0 \phi^-] + \frac{1}{2c_W} i g M [X^0 X^- \phi^+ - X^0 X^+ \phi^-] + i g M s_W [X^0 X^- \phi^+ - X^0 X^+ \phi^-] +$$

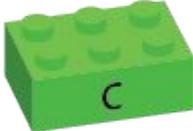
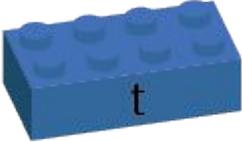
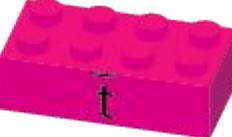
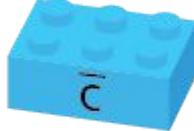
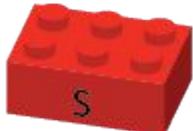
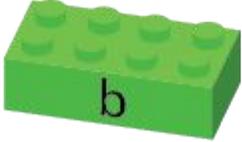
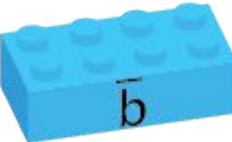
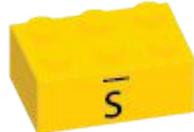
$$\frac{1}{2} i g M \bar{X}^+ X^+ \phi^0 - X^- X^- \phi^0]$$

## DO YOU NOT UNDERSTAND?

# Quarks, Leptons, Gauge Bosons and the BEH boson

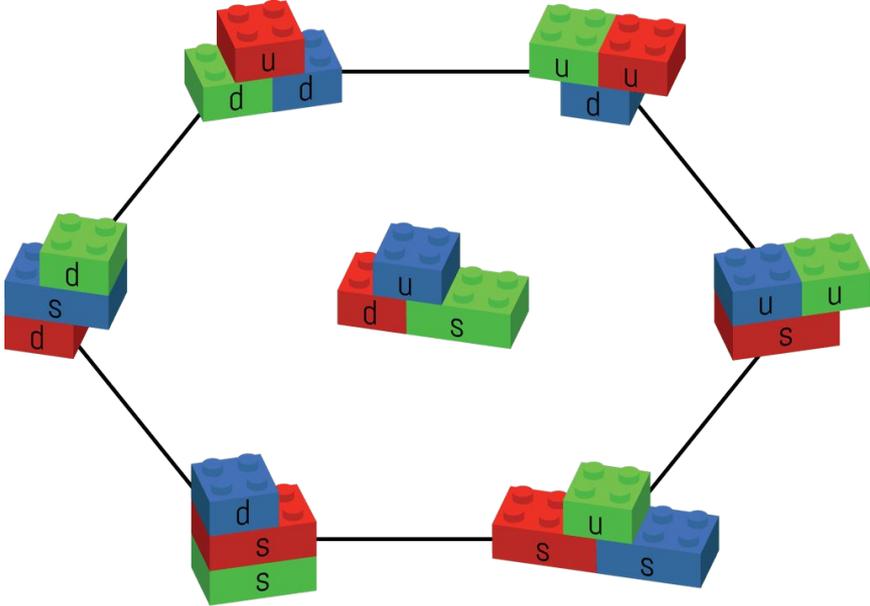


# Kvarky a antikvarky

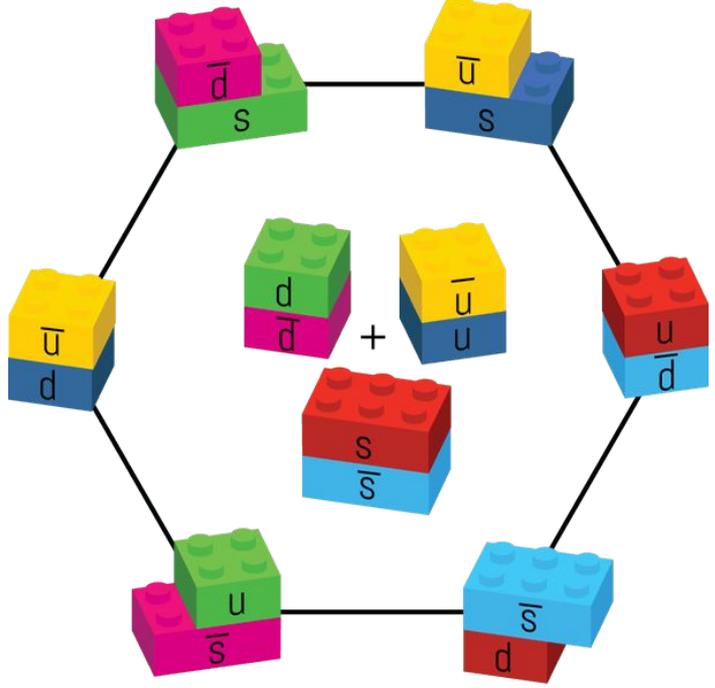
Quarks			Antiquarks		
1	2	3	3	2	1
Up 	Charm 	Top 	Antitop 	Anticharm 	Antiup 
 Down	 Strange	 Bottom	 Antibottom	 Antistrange	 Antidown

# Hadrony, skládání kvarků u, d, s

Baryony = qqq



Mesony = q + anti-q





# “Stabilní” částice: pozorovatelné v detektoru



SHORTCUTS ▾

## Particle Listings

R.L. Workman *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. 2022, 083C01 (2022) and 2023 update

Cut-off date for Listings/Summary Tables was Jan. 15, 2023. Files can be downloaded directly by clicking on the icon: .

For a key to the listings [click here](#).

Expand/Collapse All

Gauge & Higgs Bosons +

Leptons (e, mu, tau, neutrinos, heavy leptons ...)

Quarks (u,d,s,c,b,t...)

Mesons (pi, K, D, B, psi, Upsilon, ...)

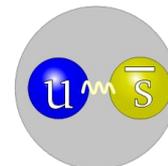
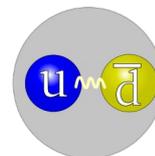
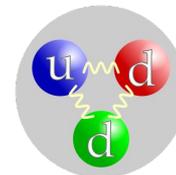
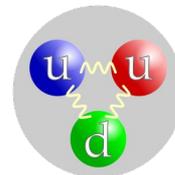
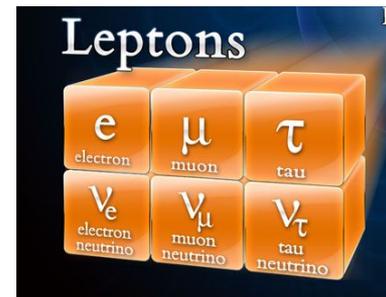
Baryons (p, n, Lambda\_b, Xi, ...)

Searches not in Other Sections +

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# “Stabilní” částice: pozorovatelné v detektoru

- Fotony,  $m=0$
- Elektrony  $e^-$ ,  $mc^2 = 0.511 \text{ MeV}$ 
  - Pozitrony  $e^+$
- Protony  $p^+$ ,  $938.3 \text{ MeV}$
- Neutrony  $n^0$ ,  $939.6 \text{ MeV}$
- Neutrína – málo interagující
- Miony  $\mu^\pm$  – doba života  $2\mu\text{s}$ ,  $c\tau = 600\text{m}$ .  $mc^2 = 105.6 \text{ MeV}$ 
  - 200x těžší než elektrony, pronikavější, kosmické záření.
- Piony,  $140 \text{ MeV}$ 
  - $\pi^\pm$  – doba života  $26\text{ns}$ ,  $c\tau = 7.8\text{m}$
  - $\pi^0$  – rychle se rozpadá
- Kaony
  - $K^+$ ,  $K^-$ ,  $c\tau = 3.7\text{m}$
  - $K^0$ , anti- $K^0$



# “Stabilní” částice: pozorovatelné v detektoru

$$m_e c^2 \doteq 0.511 \text{ MeV}$$

$$m_\mu c^2 \doteq 105.7 \text{ MeV}$$

$$m_p c^2 \doteq 938 \text{ MeV}$$

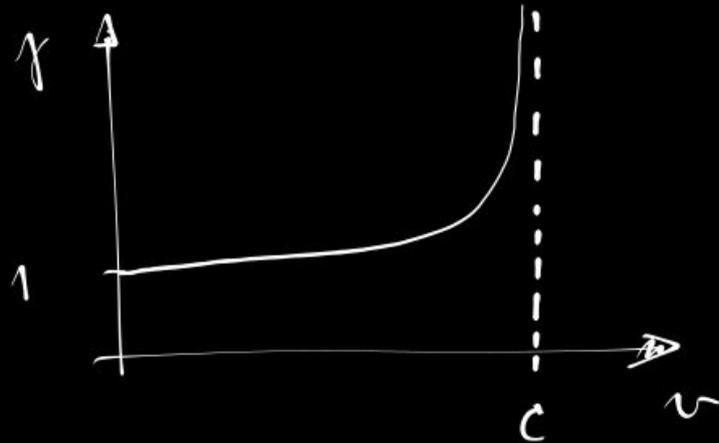
$$m_n c^2 \doteq 939 \text{ MeV}$$

$$m_{\pi^+} c^2 \doteq 140 \text{ MeV}$$

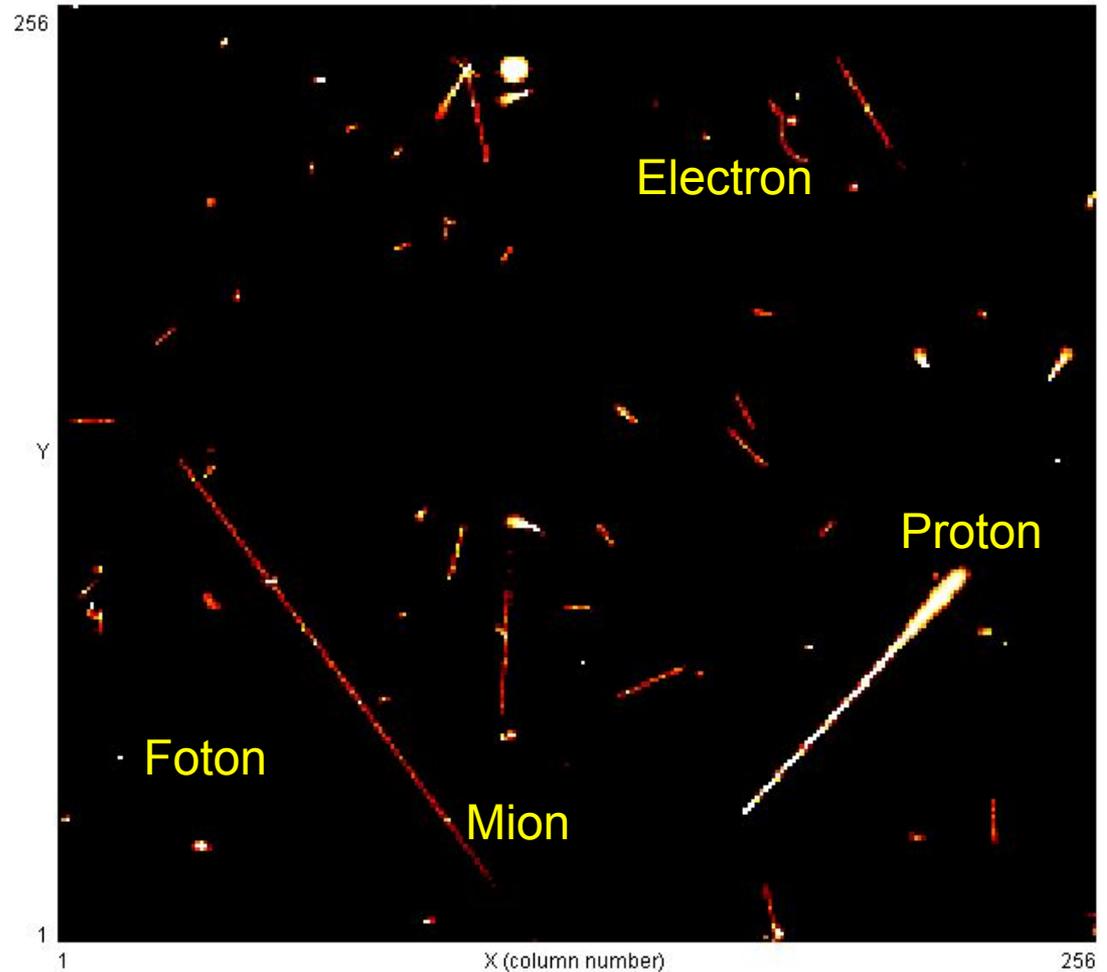
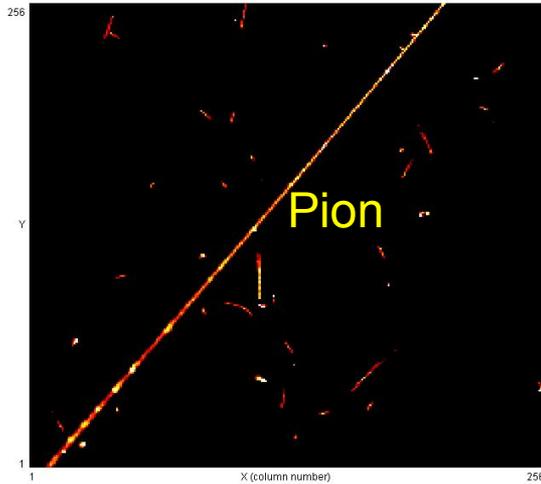
$$m_{\pi^0} c^2 \doteq 135 \text{ MeV}$$

$$E = \gamma m c^2$$

$$\gamma \equiv \frac{1}{\sqrt{1 - \beta^2}} \quad \beta \equiv \frac{v}{c}$$

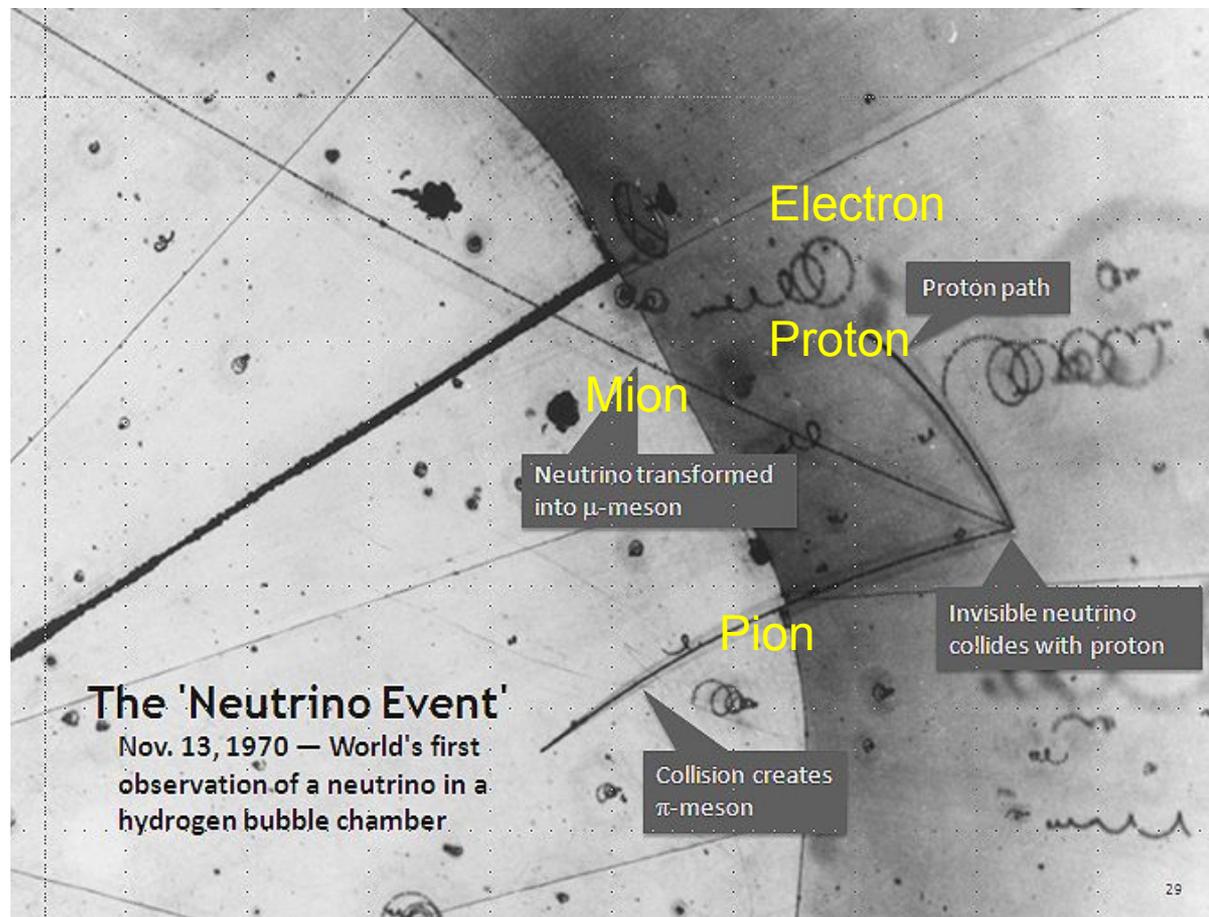


# Částicová kamera

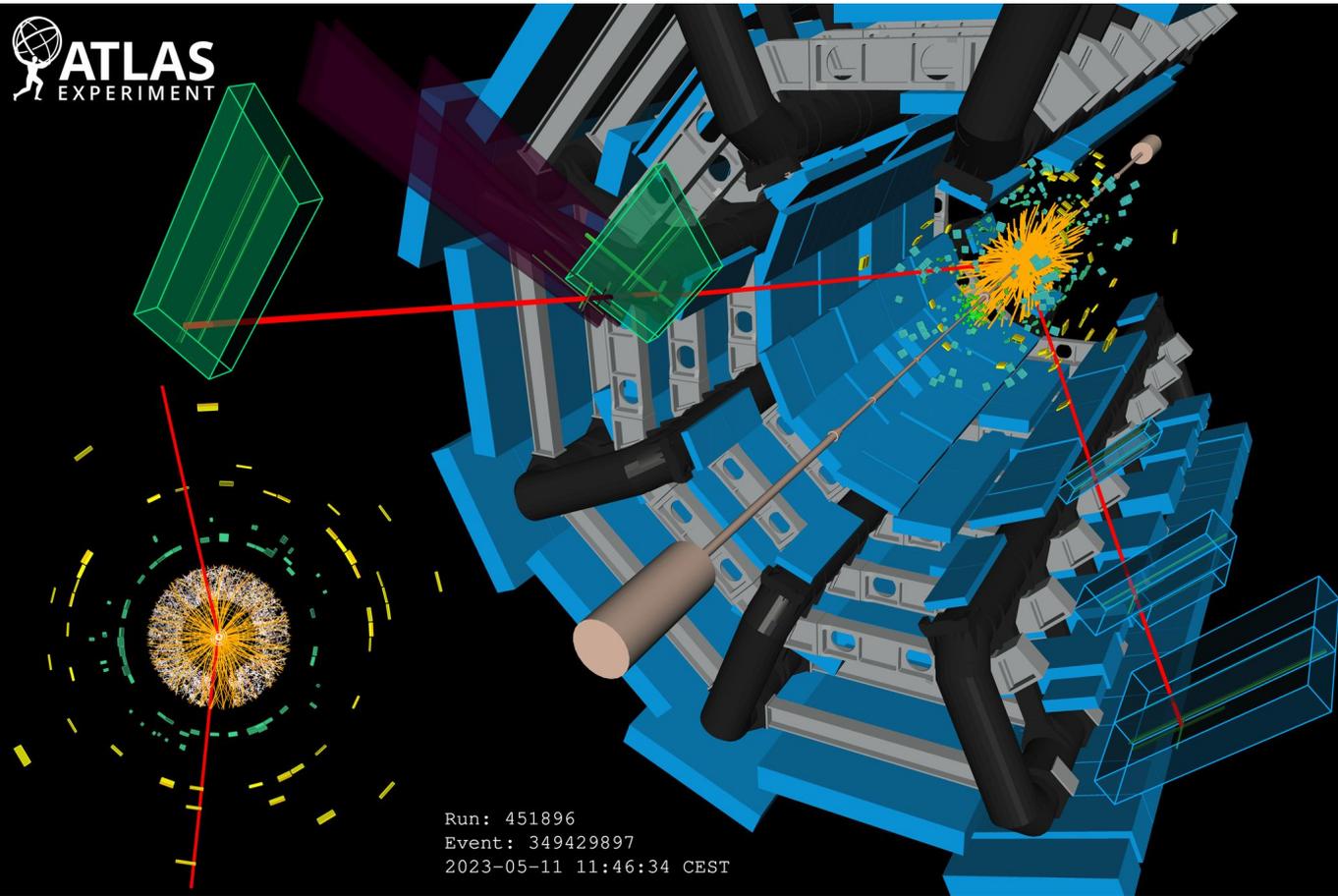


# Bublinová komora

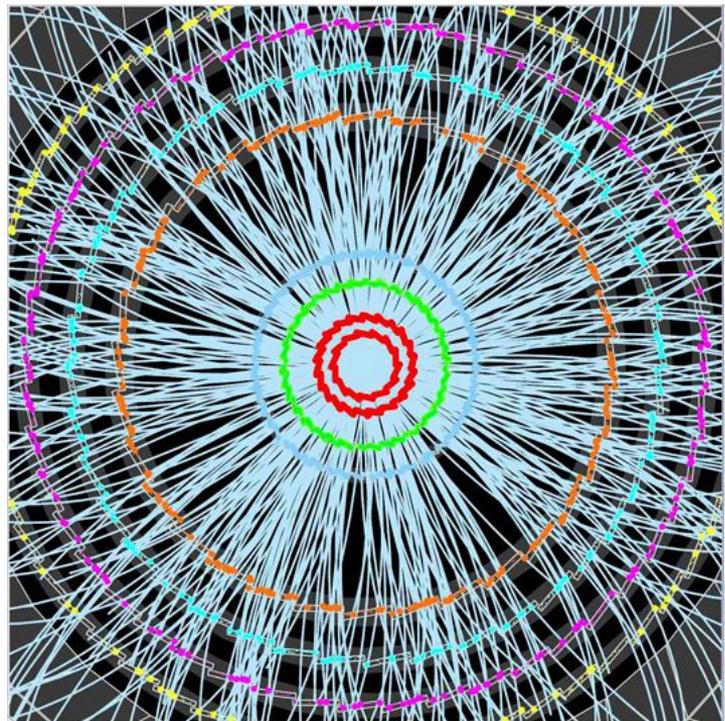
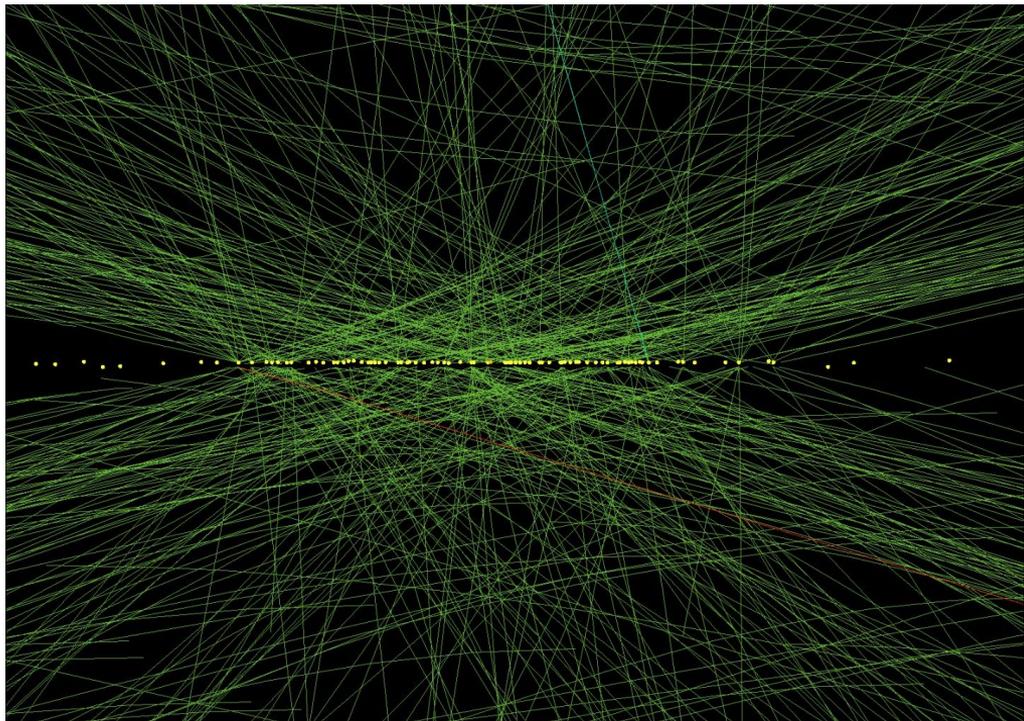
<https://en.wikipedia.org/wiki/Neutrino>



# Data ...



# Data data...



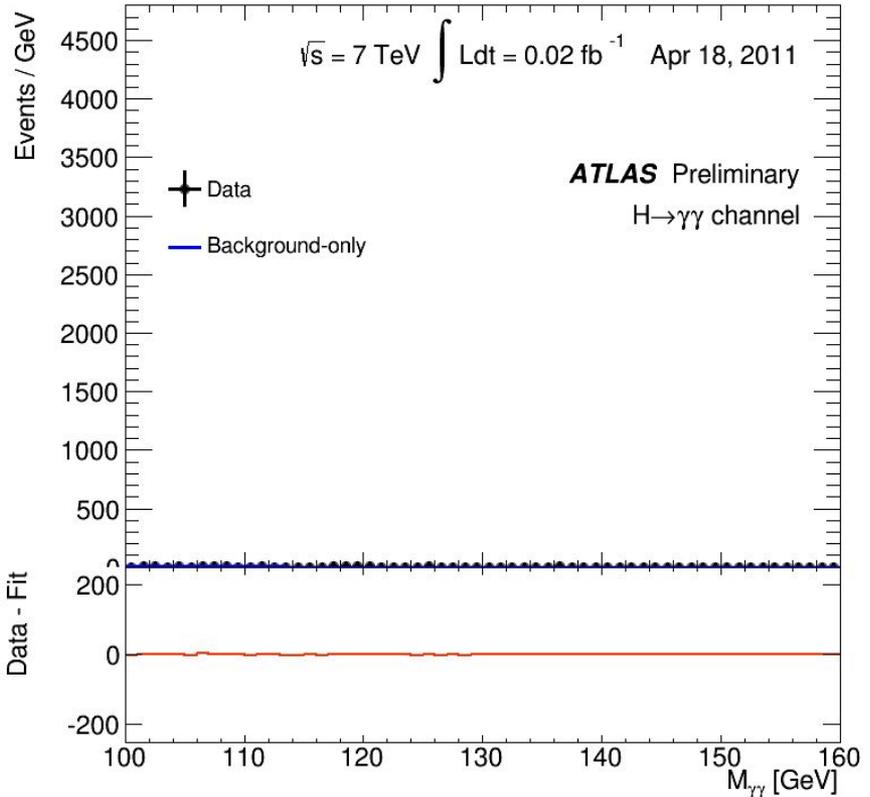
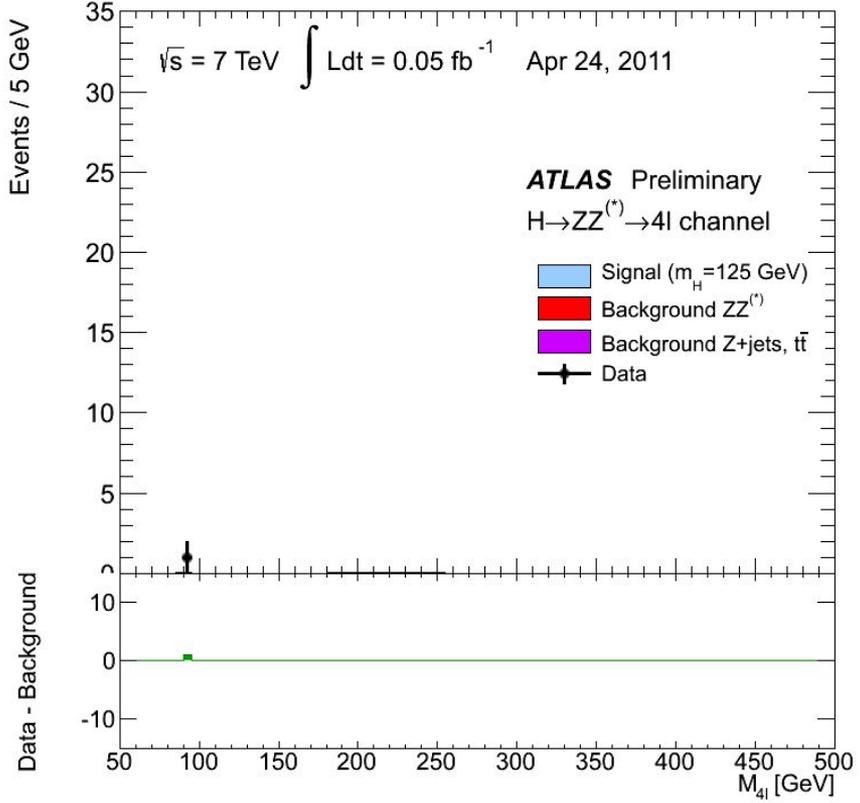
# Data ...



# Data ...

```
1414 cout << "*** Number of events in nominal and Alternative trees: " << nentries;
1415
1416
1417 // number of bytes read from TTrees
1418 Long64_t nbytes = 0;
1419 Long64_t nb_nominal = 0;
1420 Long64_t nb_nominalAlt = 0;
1421 Long64_t nb_ptcl = 0;
1422 Long64_t nb_truth = 0;
1423 m_ToRunOver = nentriesNominal;
1424 if (m_isMCsignal && m_runOverAllLevels) {
1425     // we need to go through all generated parton events
1426     // and find matching events in the detector nominal and particle trees
1427     m_ToRunOver = nentriesTruth;
1428 }
1429 if (m_isMCsignal && m_runOverPtclOnly) {
1430     m_ToRunOver = nentriesParticle;
1431 }
1432 // HACK!
1433 m_ToRunOver = 200000;
1434
1435 this -> InitLoop();
1436
1437 // +-----+
1438 // |  LOOP!  |
1439 // +-----+
1440
1441 if (m_isData) {
1442     cout << "Will run over " << m_ToRunOver << " entries." << endl;
1443 } else {
1444     cout << "Will run over " << m_ToRunOver << " entries while sumWeights is 1." << endl;
1445     m_h_sumWeights -> SetBinContent(1, m_sumWeights);
1446     m_h_sumWeightsSq -> SetBinContent(1, m_sumWeightsSq);
1447 }
1448
1449
1450
1451 for (Long64_t jentry = 0; jentry < m_ToRunOver; jentry++) {
1452     if (jentry % verbose == 0) {
1453         cout << "Processing " << jentry << "/" << m_ToRunOver << endl;
1454     }
1455     nb_nominal = 0;
```

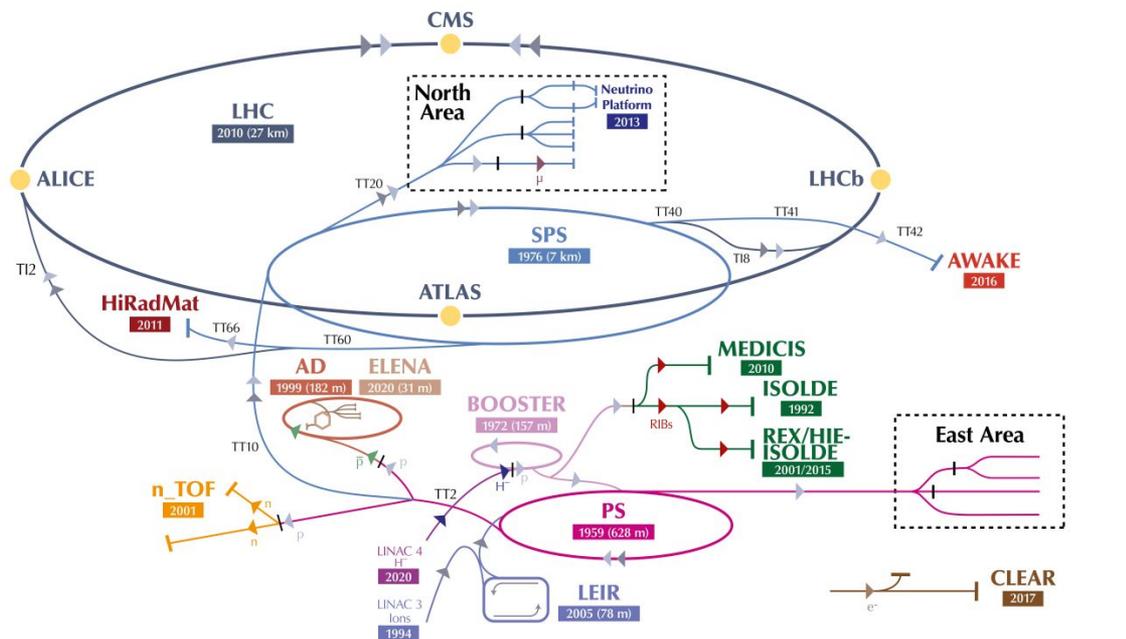
# Data ...



**And now...  
From Big Data  
and  
Big Experiments...  
To Table-Top Experiments!**

# Komplex urychlovačů CERN

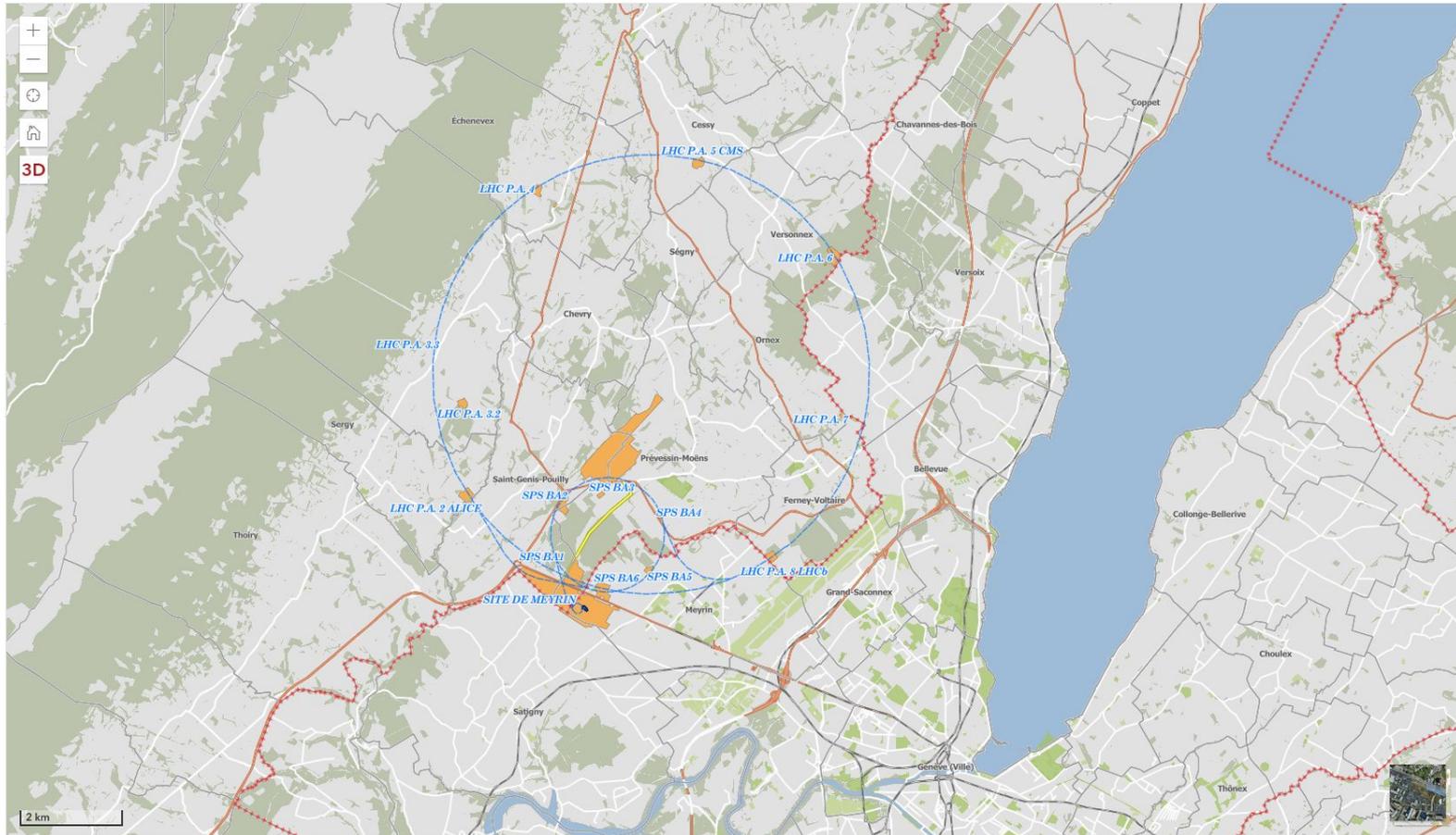
The CERN accelerator complex  
Complexe des accélérateurs du CERN



<https://home.cern/science/accelerators/accelerator-complex>

▶  $H^-$  (hydrogen anions) ▶ p (protons) ▶ ions ▶ RIBs (Radioactive Ion Beams) ▶ n (neutrons) ▶  $\bar{p}$  (antiprotons) ▶  $e^-$  (electrons) ▶  $\mu$  (muons)

LHC - Large Hadron Collider // SPS - Super Proton Synchrotron // PS - Proton Synchrotron // AD - Antiproton Decelerator // CLEAR - CERN Linear Electron Accelerator for Research // AWAKE - Advanced WAKEfield Experiment // ISOLDE - Isotope Separator OnLine // REX/HIE-ISOLDE - Radioactive Experiment/High Intensity and Energy ISOLDE // MEDICIS // LEIR - Low Energy Ion Ring // LINAC - LINear ACcelerator // n\_TOF - Neutrons Time Of Flight // HiRadMat - High-Radiation to Materials // Neutrino Platform





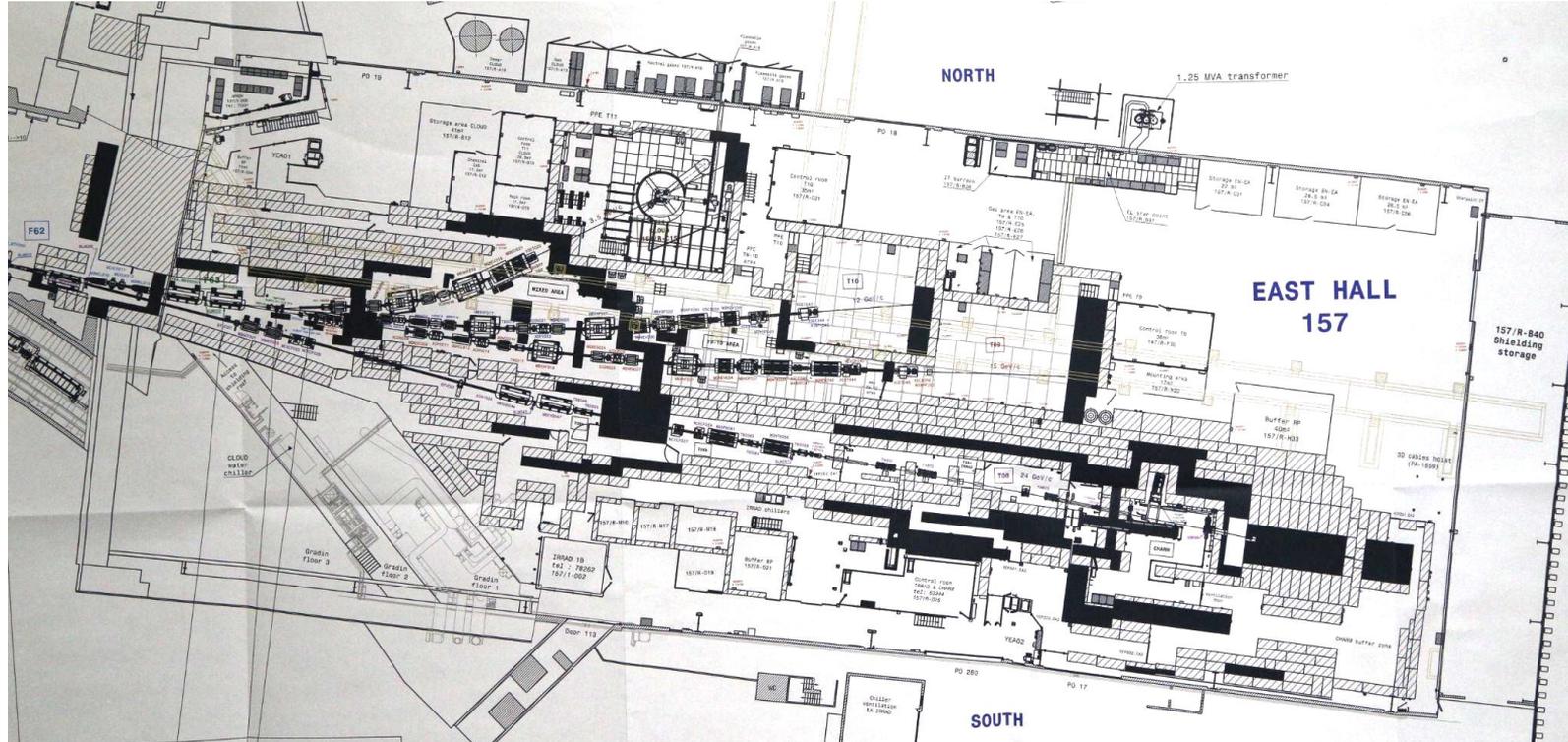
# East Area 2022

- Building 157 at CERN's Meyrin site, large experimental hall housing also the CLOUD experiment, T8, T9 and T10 beam lines.



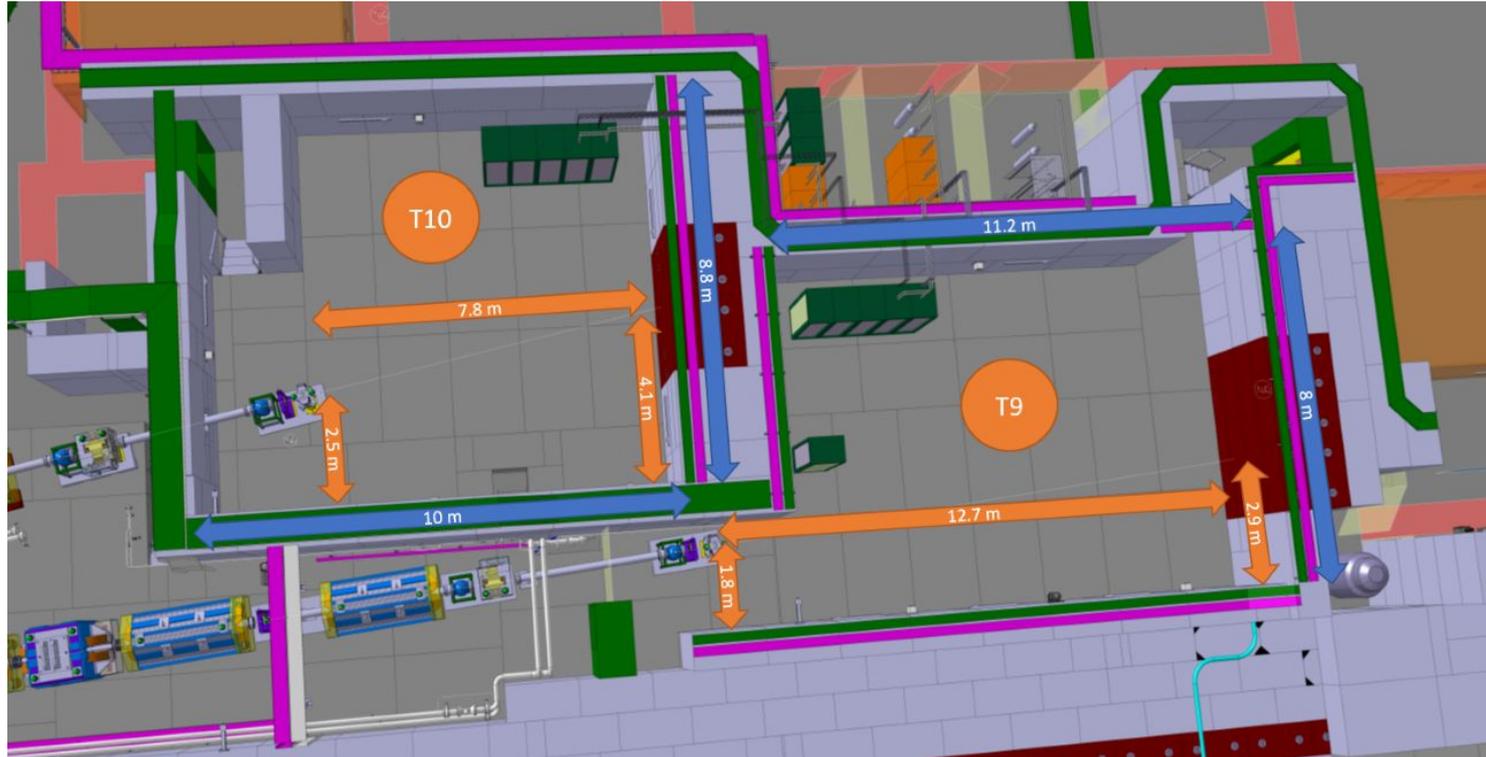
# East Area

- Building 157 at CERN's Meyrin site, large experimental hall housing also the CLOUD experiment, T8, T9 and T10 beam lines.



# East Area

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# East Area

- Building 157 at CERN's Meyrin site, large experimental hall housing also the CLOUD experiment, T8, T9 and T10 beam lines.

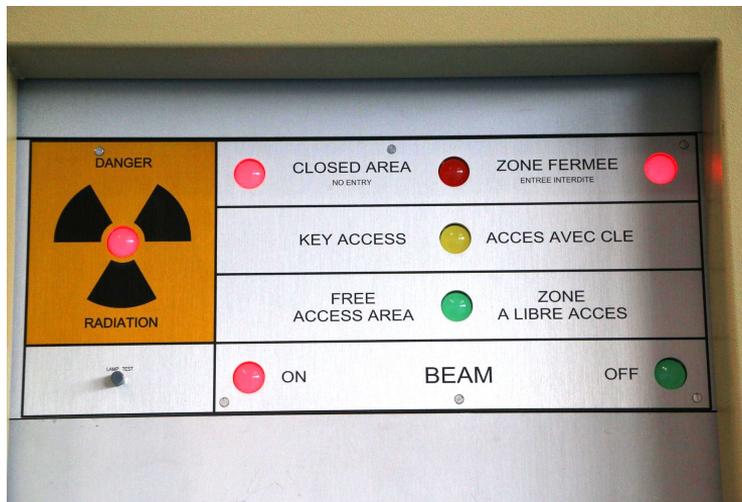


# East Area



# East Area 2022

- Experimentální hala v budově 157
- Svazkové linie T8, T9 and T10, CLOUD experiment...



# East Area 2023

- Experimentální hala v budově 157
- Svazkové linie T8, T9 and T10, CLOUD experiment...



# East Area 2023

- Experimentální hala v budově 157
- Svazkové linie T8, T9 and T10, CLOUD experiment...

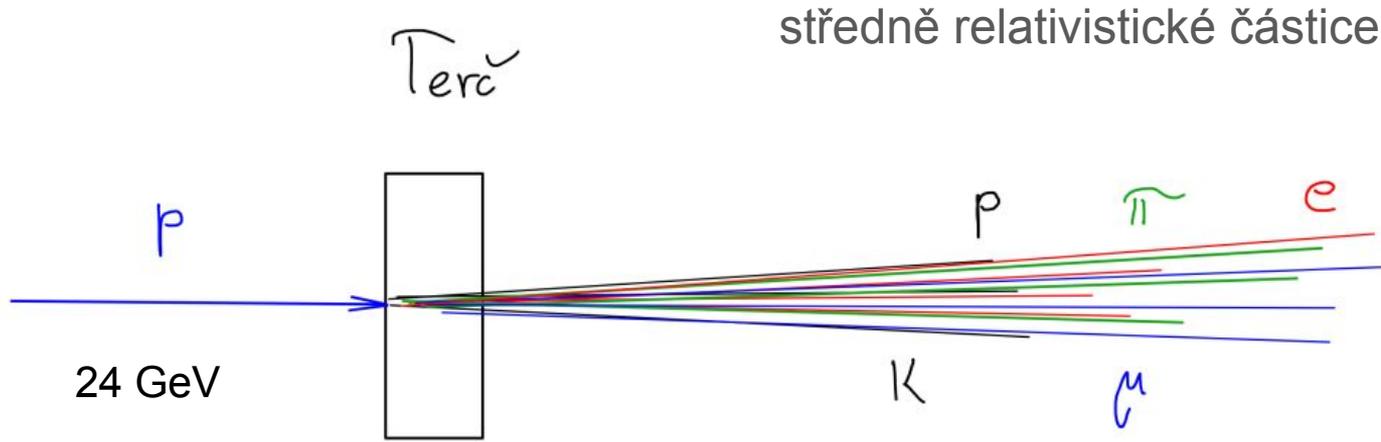


# East Area 2023

- Experimentální hala v budově 157
- Svazkové linie T8, T9 and T10, CLOUD experiment...



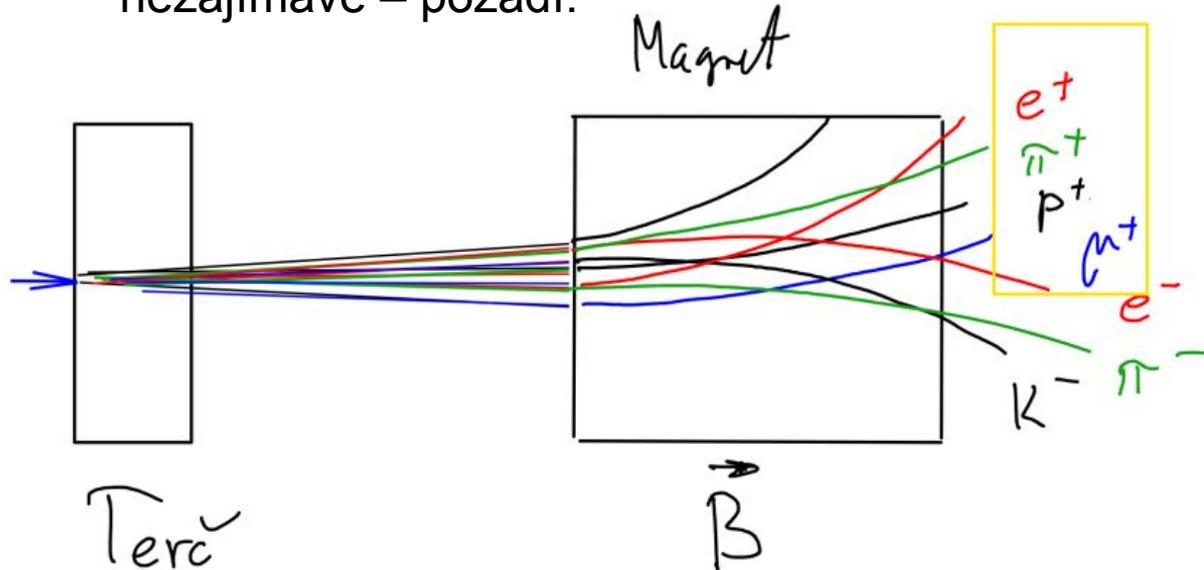
# East Area: Sekundární svazky částic pro testování



$$|\vec{p}| \quad 200 \text{ MeV} \approx 1 \text{ GeV}$$

# Analýza složení svazku

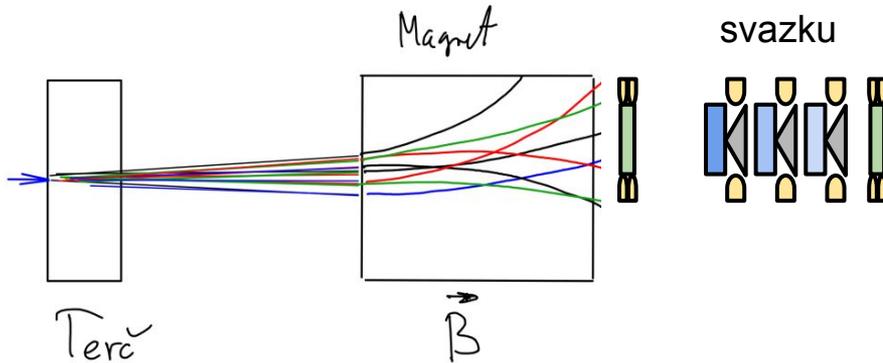
- Sekundární svazek částic o různých hybnostech
- Dominováno elektrony, ale ty mohou být nezajímavé – pozadí.



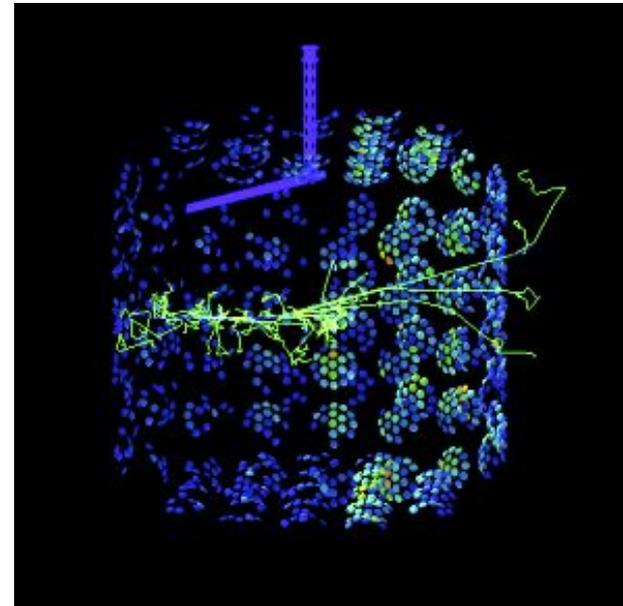
$$p = m v \quad \beta \equiv \frac{v}{c}$$
$$p = \gamma m v = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} m v$$
$$p = \beta \gamma m c = \frac{1}{\sqrt{1 - \beta^2}} \beta m c$$

# Analýza složení svazku

- Jaké je složení svazku?
- Bude v něm dost pionů a mionů na kalibraci odezvy detektoru na tyto částice?
- ⇒ Identifikace částic!
  - pro budovaný experiment.

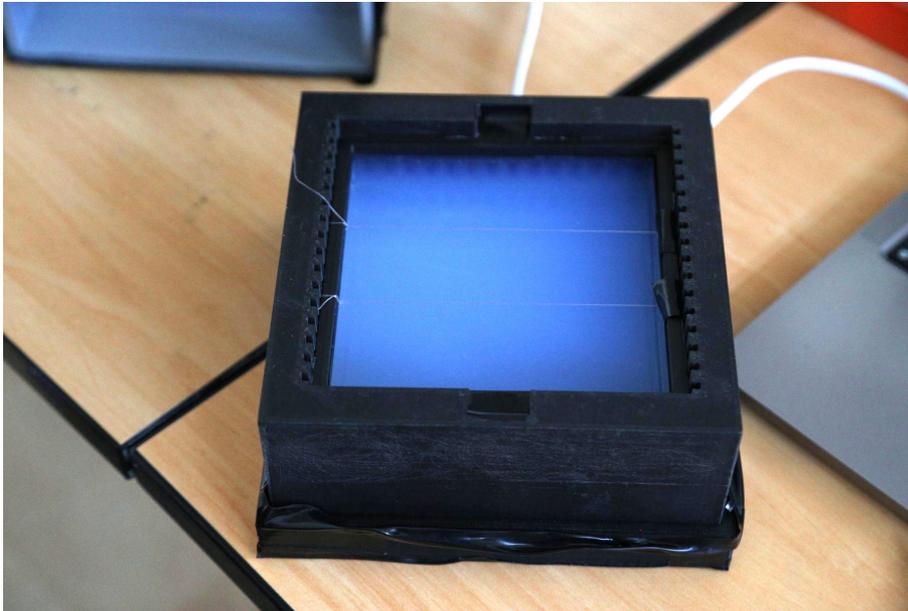


Budoucí experiment



# Installation :: Aerogelové Čerenkovské detektory

- Průhledný materiál, v kterém září částice, pokud se pohybují rychlostí větší než rychlost světla v daném prostředí: elektromagnetická rázová vlna.
- **Aerogel: Nejlehčí pevná látka!**
- Světlo je reflexní fólií odraženo do citlivých detektorů světla, fotonásobičů (PhotoMultiplier Tube, PMT):

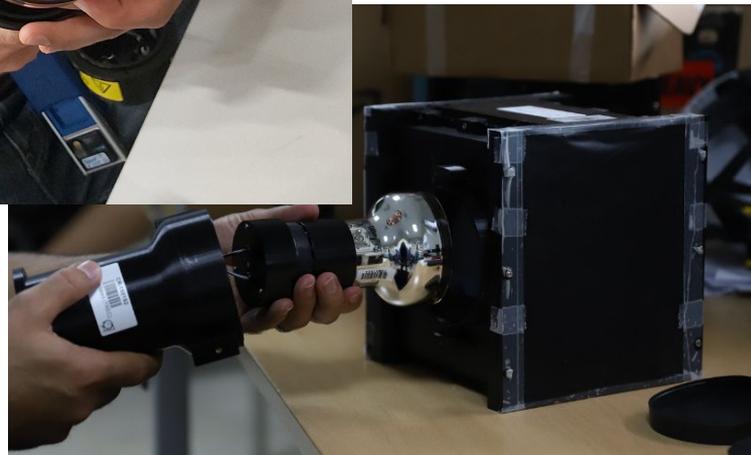


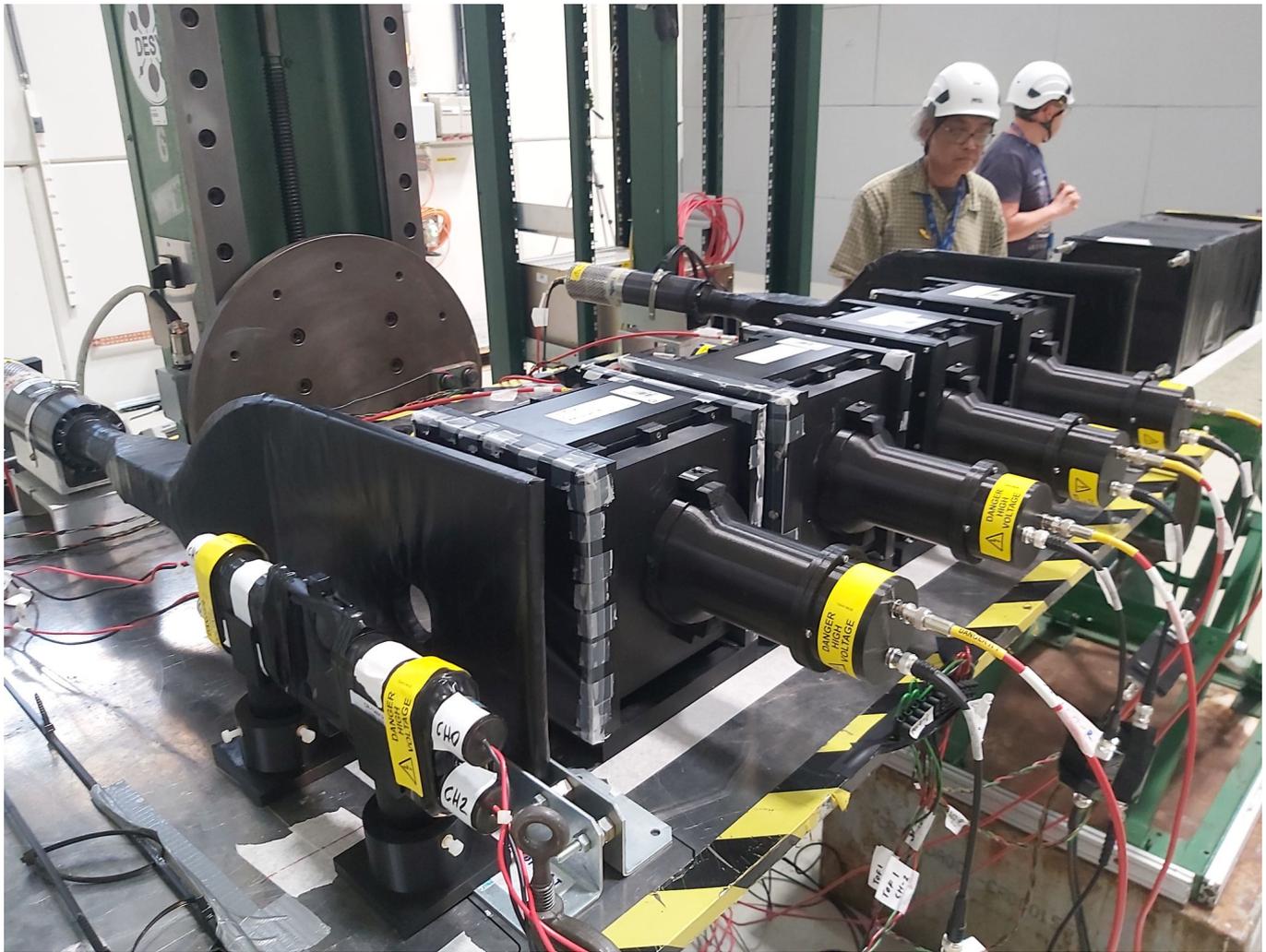
# Aerogel

- Nejlehčí pevná látka: World's Lightest Solid!
  - [https://www.youtube.com/watch?v=AeJ9q45PfD0&ab\\_channel=Veritasium](https://www.youtube.com/watch?v=AeJ9q45PfD0&ab_channel=Veritasium)



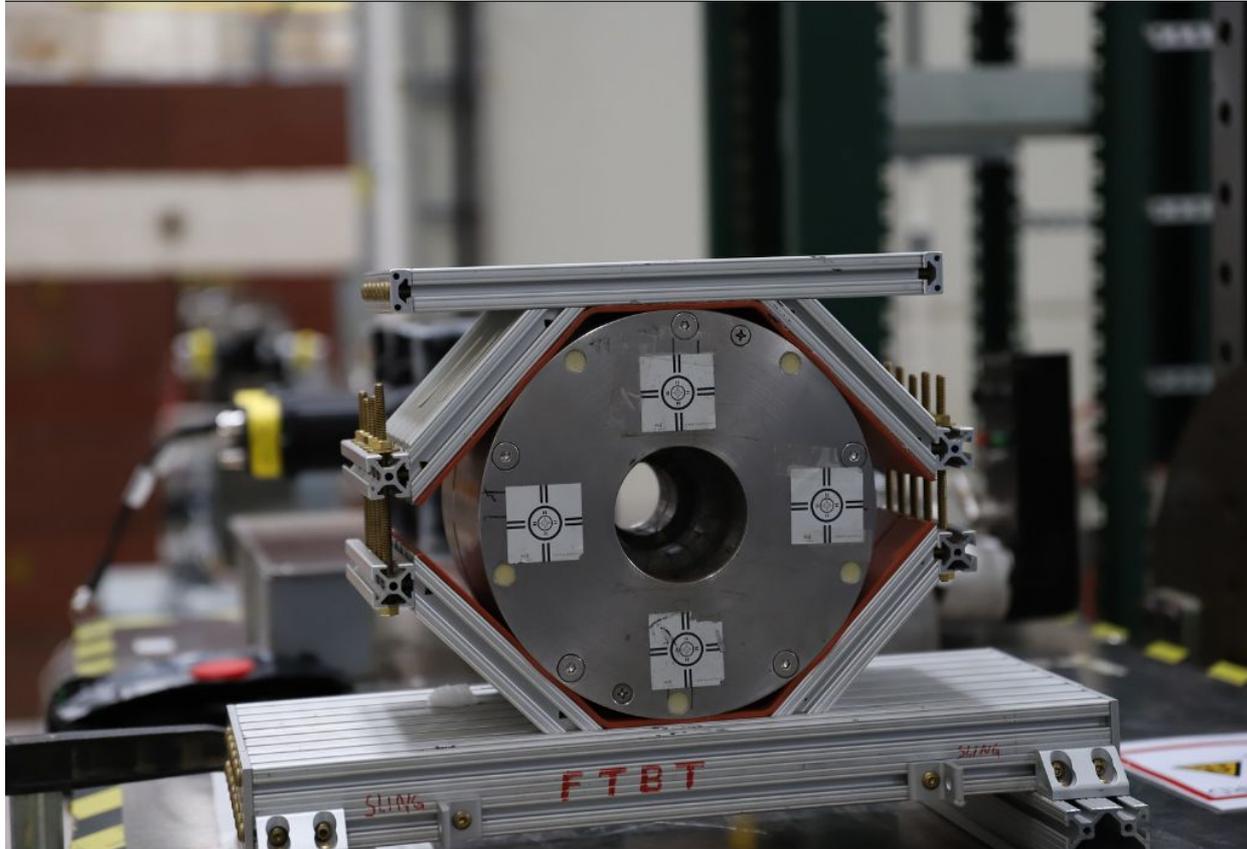
# Fotonásobič



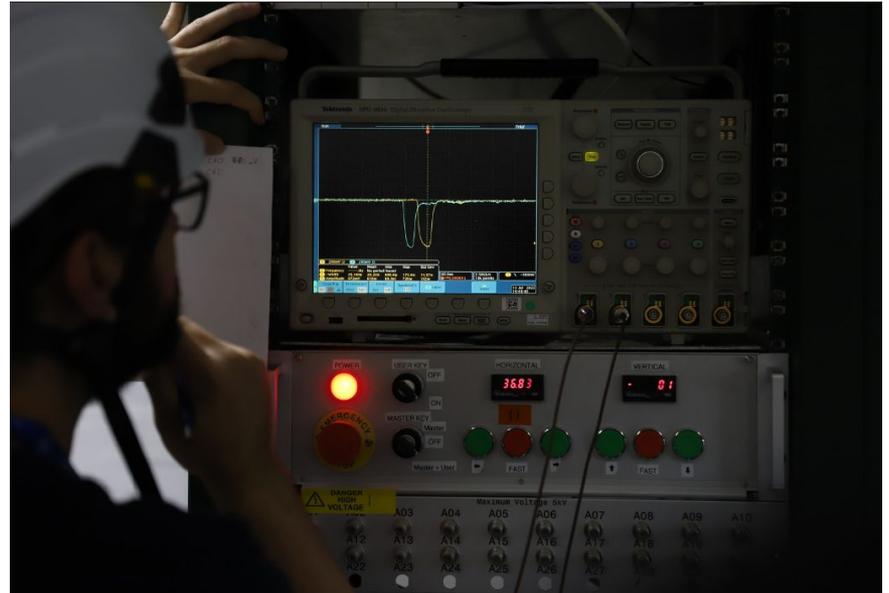




# Meet Mr. Permanent Magnet of 1.7T!

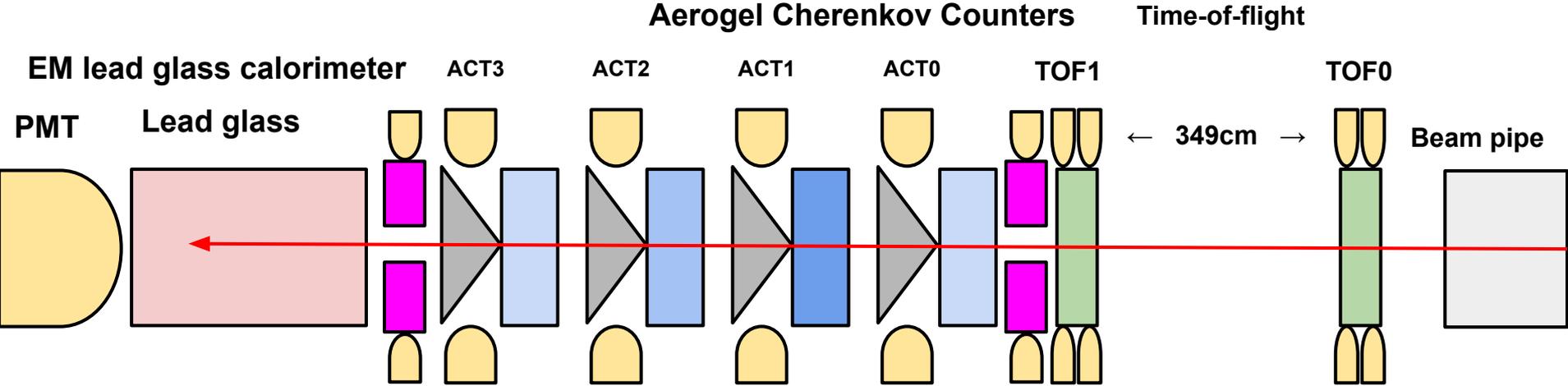








# Setup overview updated 2023



hole  
counters

## Legend

← Beam

PMT

Aerogel

Mirror

Trigger  
scintillators  
/ old ToF

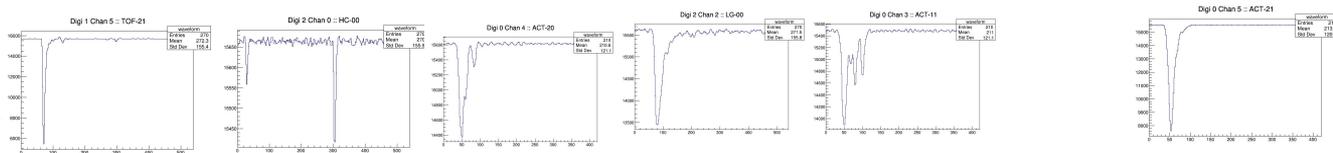
Lead glass

# Setup overview updated 2023

EM lead glass calorim

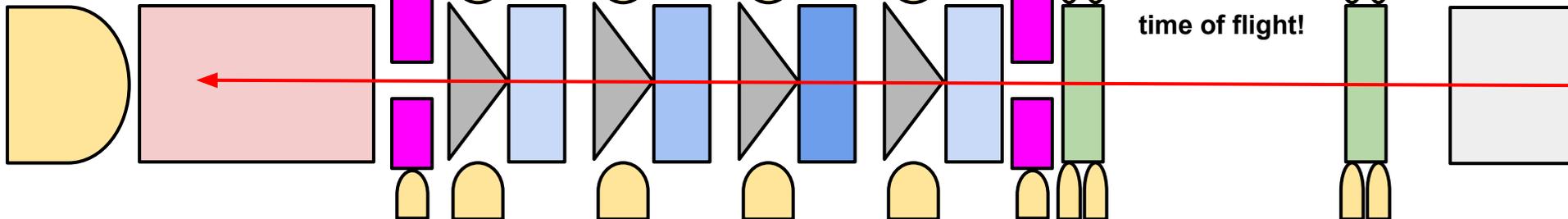
PMT

Lead glass

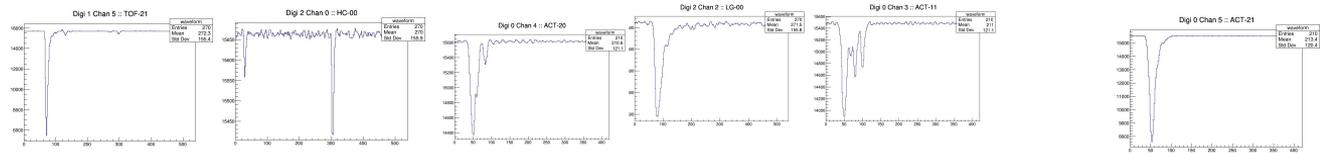


← 349cm →  
time of flight!

Beam pipe



hole  
counters



← Beam

PMT

Aerogel

Mirror

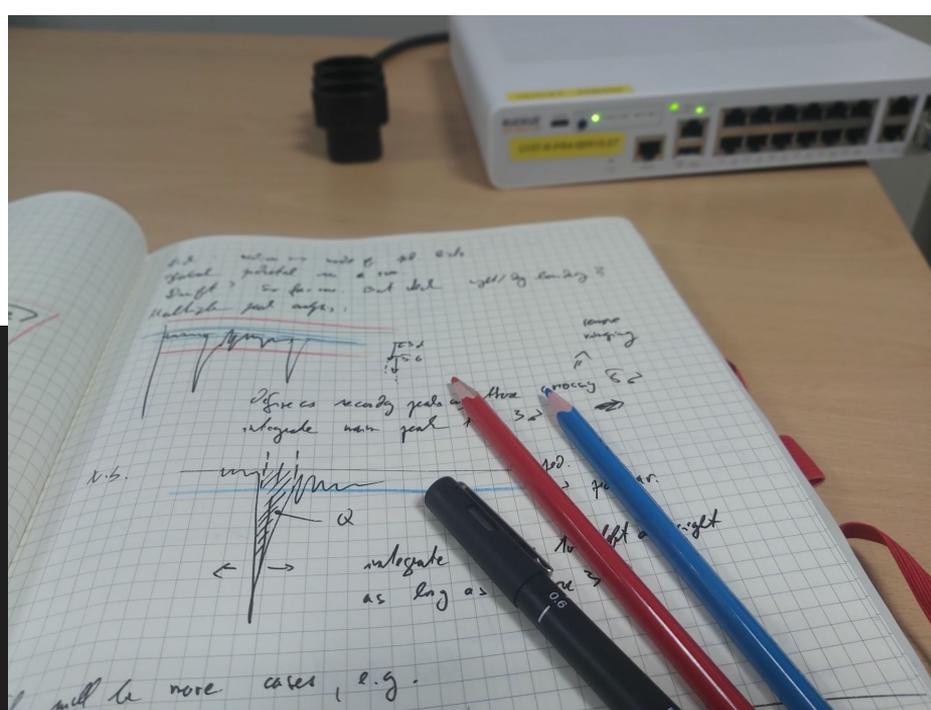
Trigger  
scintillators  
/ old ToF

Lead glass

```

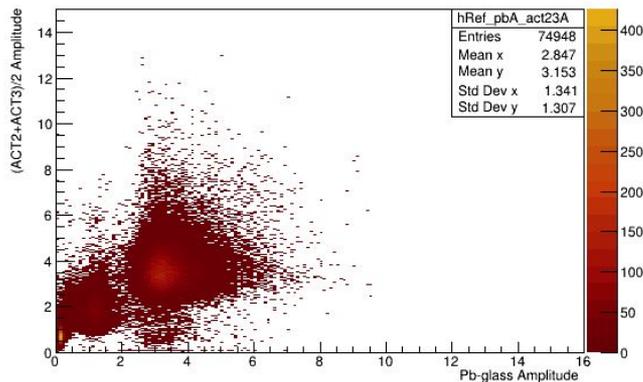
366 // +-----+
367 // |         event loop         |
368 // +-----+
369
370 cout << "Event loop!" << endl;
371
372 // TODO:
373 // check also the number of entries in the trees?
374
375 for(int ientry = 0; ientry < _ent[0]; ientry++) {
376
377     if (ientry % verbose == 0) {
378         cout << "processing " << ientry << " / " << _ent[0] << endl;
379     }
380
381     _eventInfo -> GetEntry(ientry);
382     Long64_t RunNumber = _eventInfo -> RunNumber;
383     Int_t EventNumber = _eventInfo -> EventNumber;
384     Int_t SpillNumber = _eventInfo -> SpillNumber;
385     // cout << " RunNumber=" << RunNumber << " EventNumber=" << EventNumber << " SpillNumber=" << SpillNumber << endl;
386
387     for (int ich = 0; ich < _nChannels; ++ich) {
388         if (_debug) cout << "getting entry for " << _treeNames[ich] << endl;
389         _reader[ich] -> GetEntry(ientry);
390     }
391     if (_debug) cout << "done" << endl;
392
393
394     this -> ReadChannels();
395
396     // peak cuts on demand
397     if (! _isHodoscopeRun) {
398         if (! this -> PassedPeakCuts())
399             continue;

```

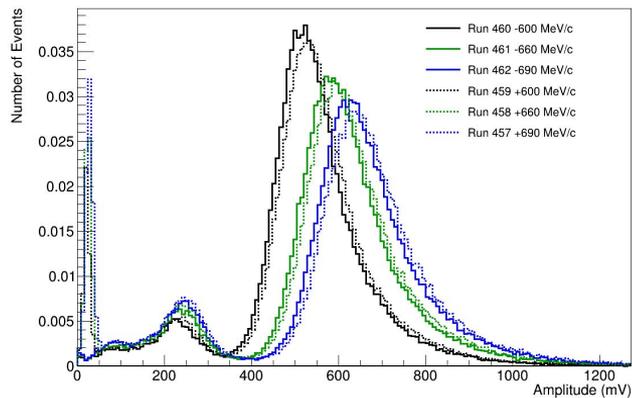


# Současná analýza dat z více fotonásobičů

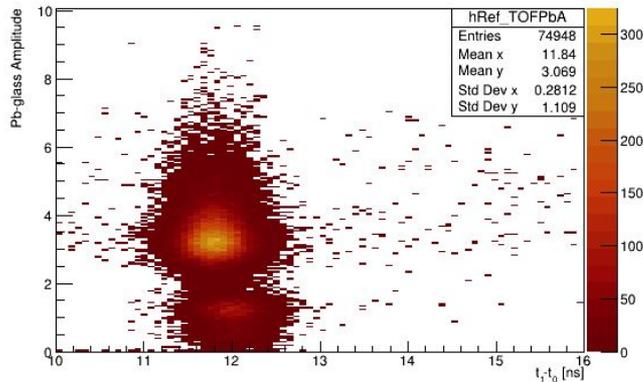
ACT2+3 vs Lead Glass 408



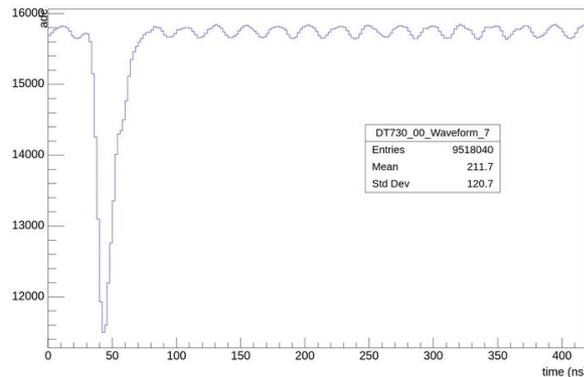
DT730 amplitude for digitizer 02 channel=2



Lead glass vs TOF 408

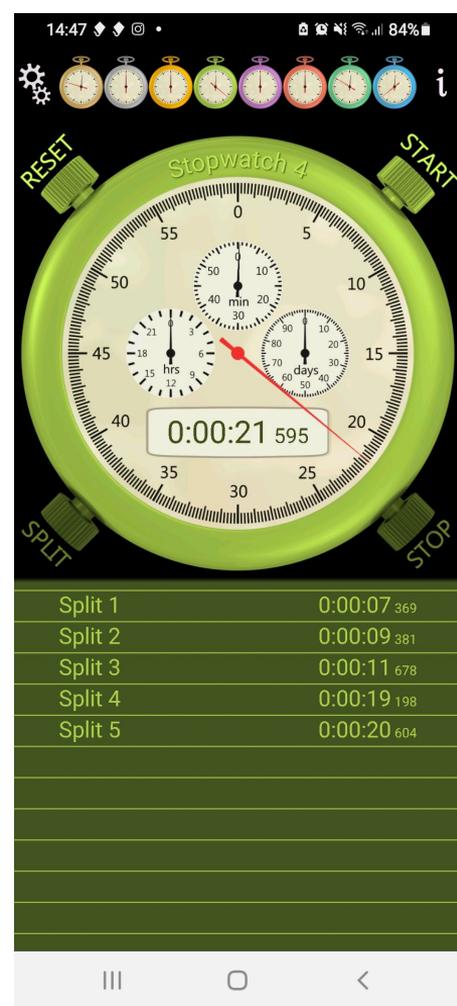


DT730 waveform for digitizer=00 ch=7



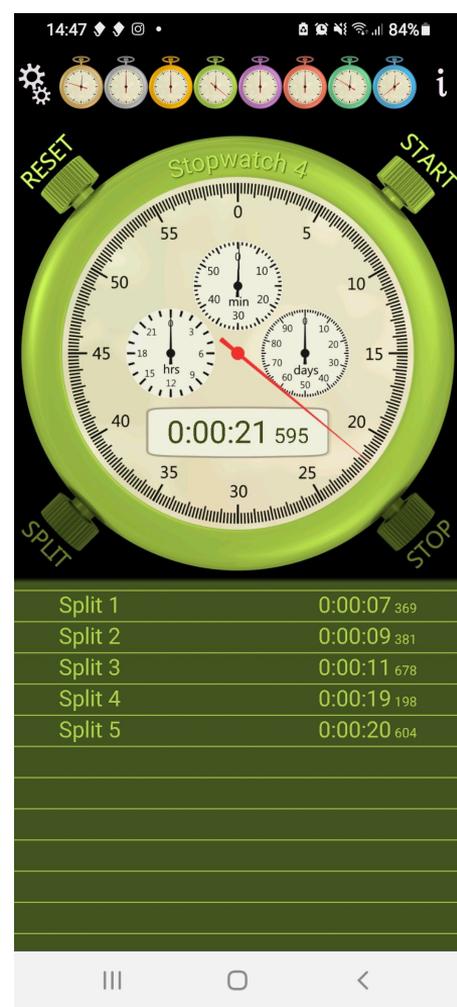
# Měření doby letu částic = time of flight!

- Lze použít k identifikaci částic.
- Time-of-flight scintillators, 349cm apart.



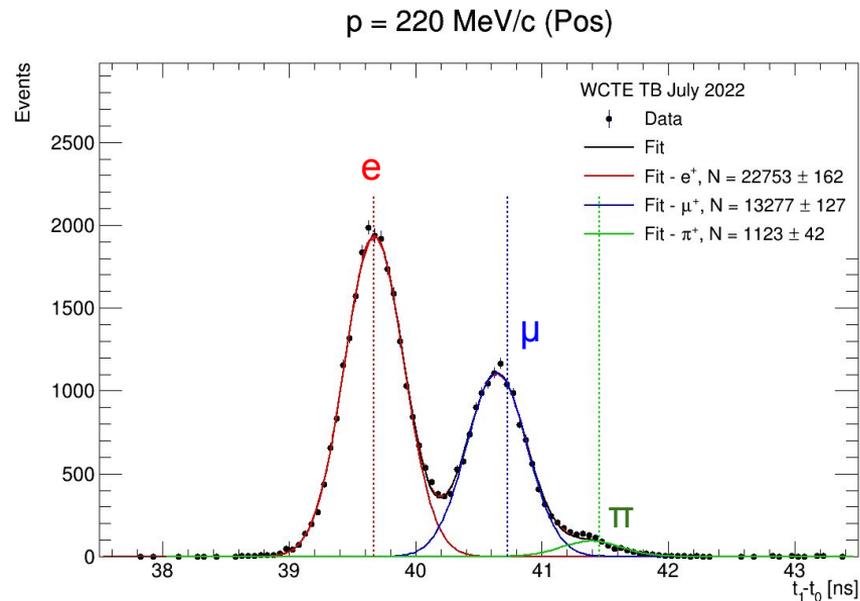
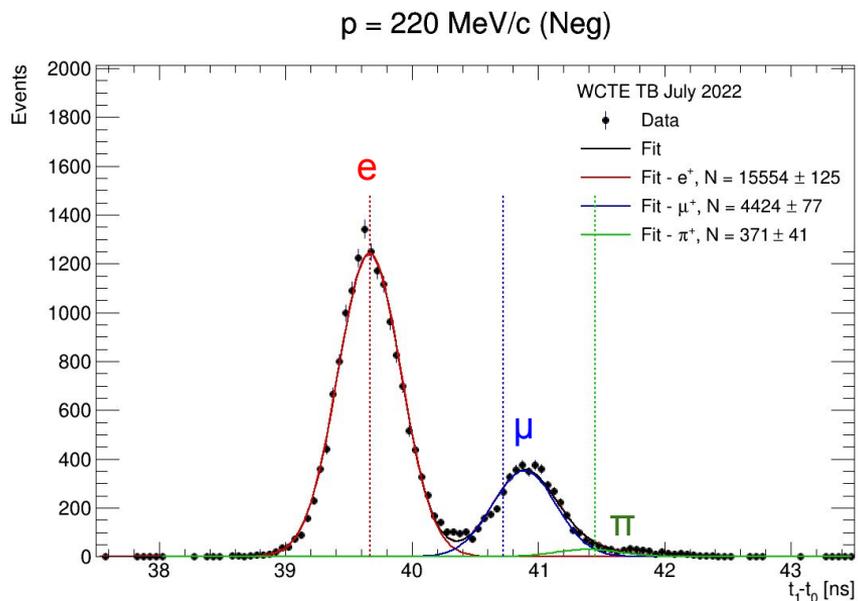
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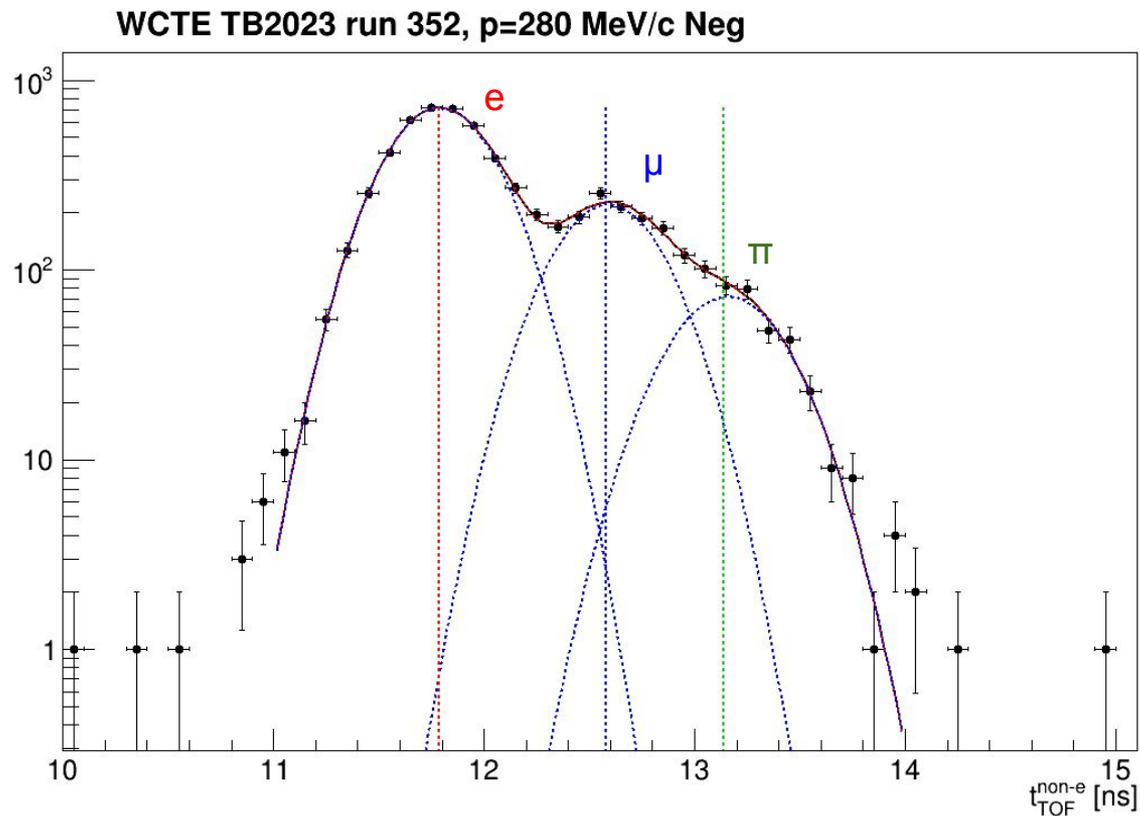


# Výsledky 2022

- Vertikální linky: teoretické časy přiletu  $\mu$  a  $\pi$  za peakem elektronů.

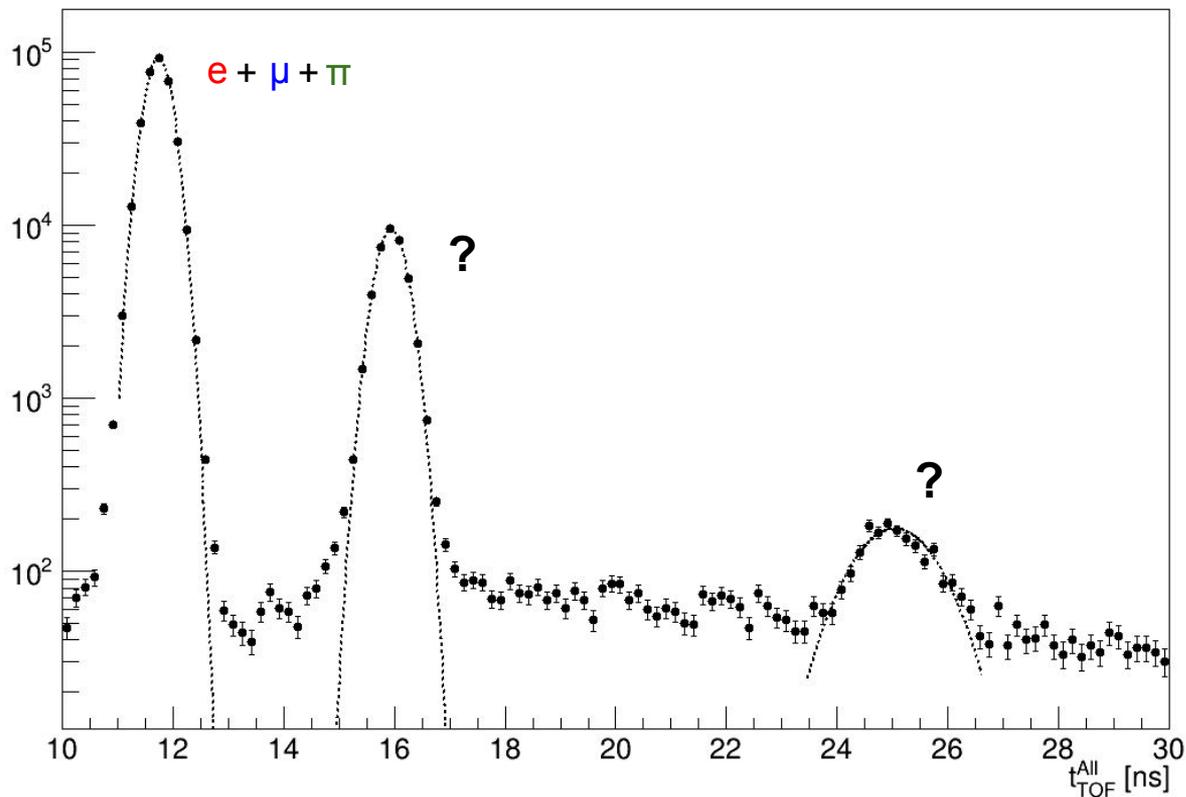


# Výsledky 2023



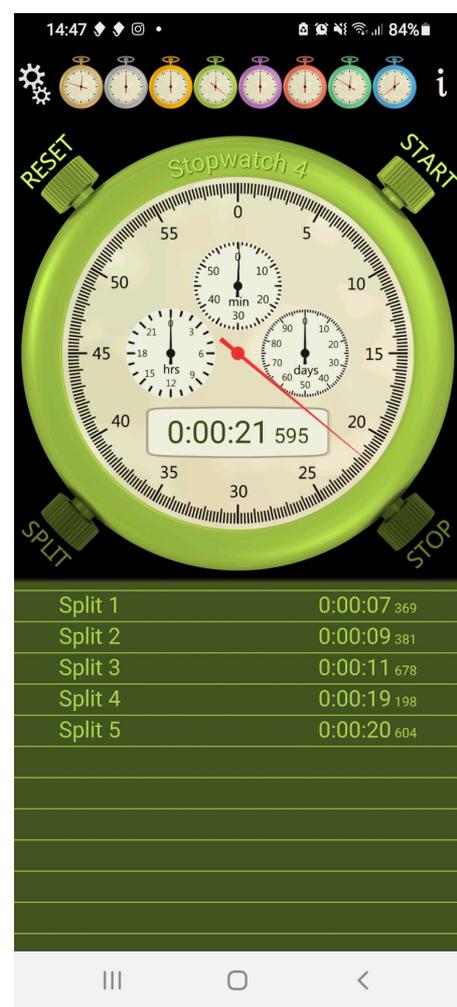
# Kladně nabité částice, větší hybnost 1000 MeV/c

- Při vyšších hybnostech časy příletu elektronů, mionů a pionů už splývají.
- Ale objevují se další peaky ve svazku pozitivních částic...

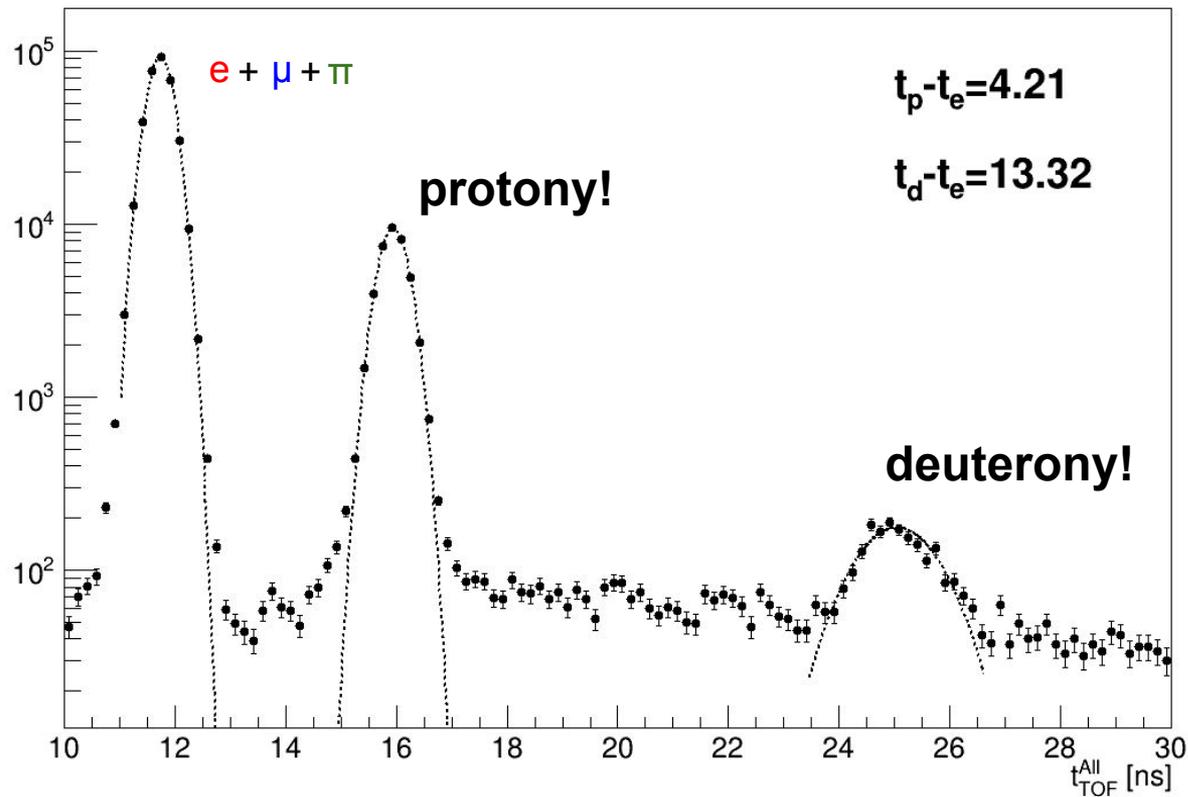


# Doba letu

- Pro známou hybnost svazku lze spočítat očekávané časy příletu
- Kdy doletí elektron, mion, pion...?
  - $m_e c^2 = 0.511 \text{ MeV}$
  - $m_\mu c^2 = 105.6 \text{ MeV}$
  - $m_\pi c^2 = 139.6 \text{ MeV}$
  - $m_p c^2 = 938.27 \text{ MeV}$
  - $m_n c^2 = 939.565 \text{ MeV}$
  - $m_D c^2 = 1875.613 \text{ MeV}$

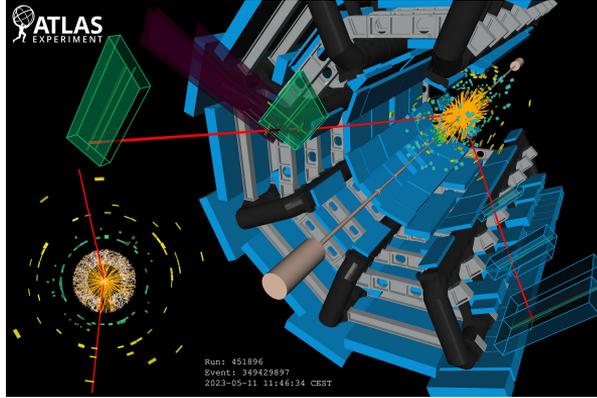


# Kladně nabité částice, větší hybnost 1000 MeV/c



**You identified and  
found your own  
particle!**

**The End**



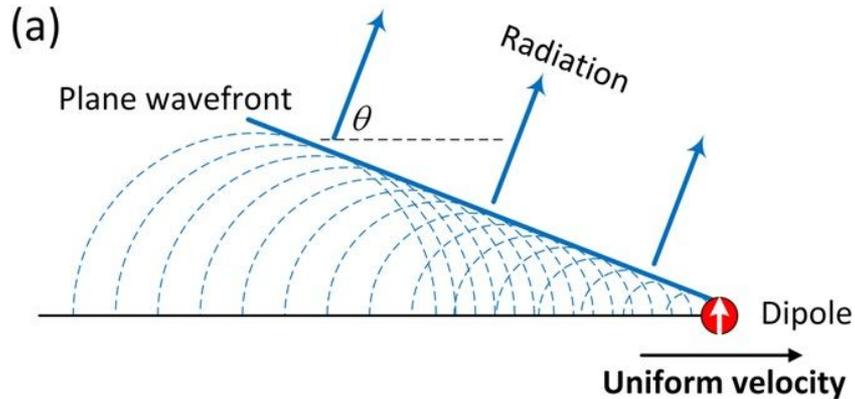
# Zdroje

- <https://home.cern/about/who-we-are/our-history>
- <https://home.cern/news/news/cern/year-celebrate-cerns-70th-anniversary>
- <https://visit.cern/science-gateway>
- <https://home.cern/about/who-we-are/our-history>
- <https://home.cern/news/news/cern/cern-highlights-2023>

# Backup

# Čerenkovovo záření

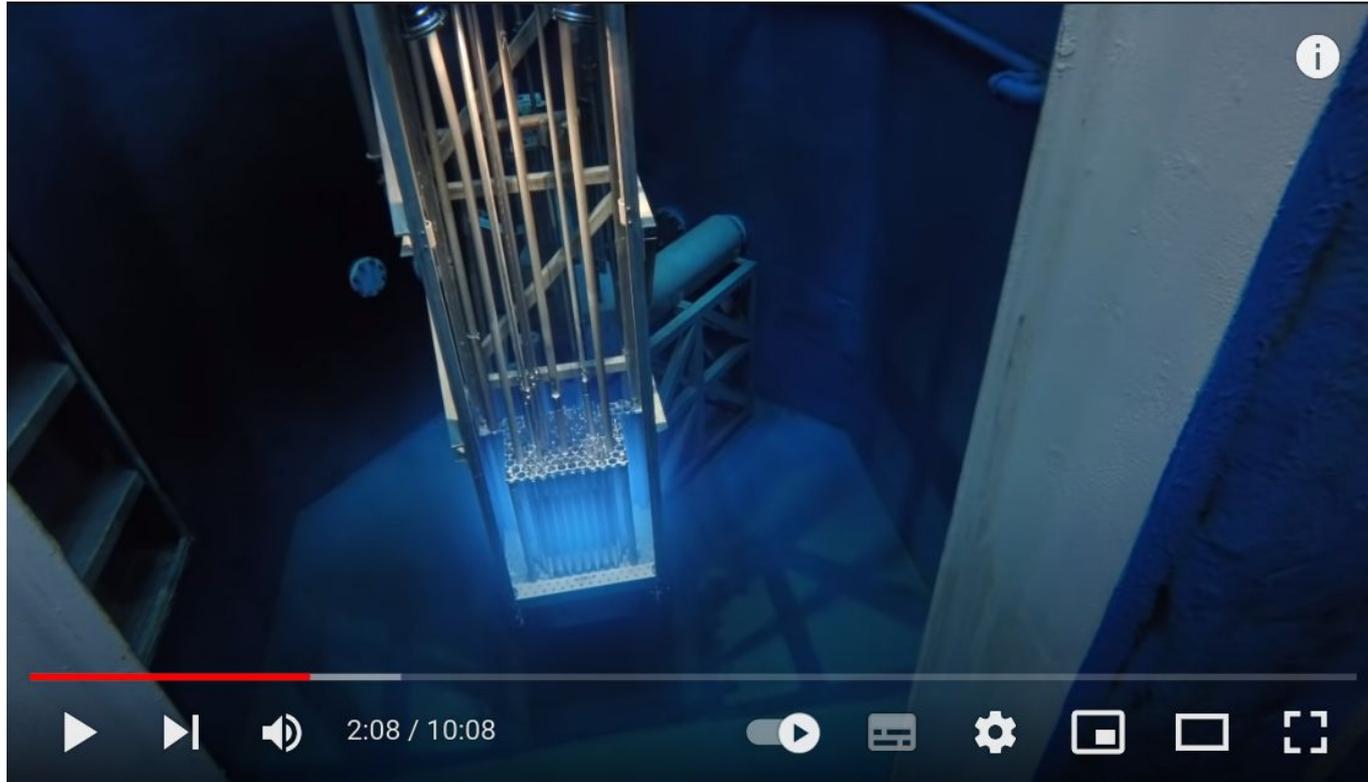
- Nabitá částice překračuje rychlost světla v hmotném prostředí
  - ideálně transparentní médium
- $v_{\text{světla}} = c / \text{index lomu} = c / n$



<https://www.nature.com/articles/s41598-017-08705-4>

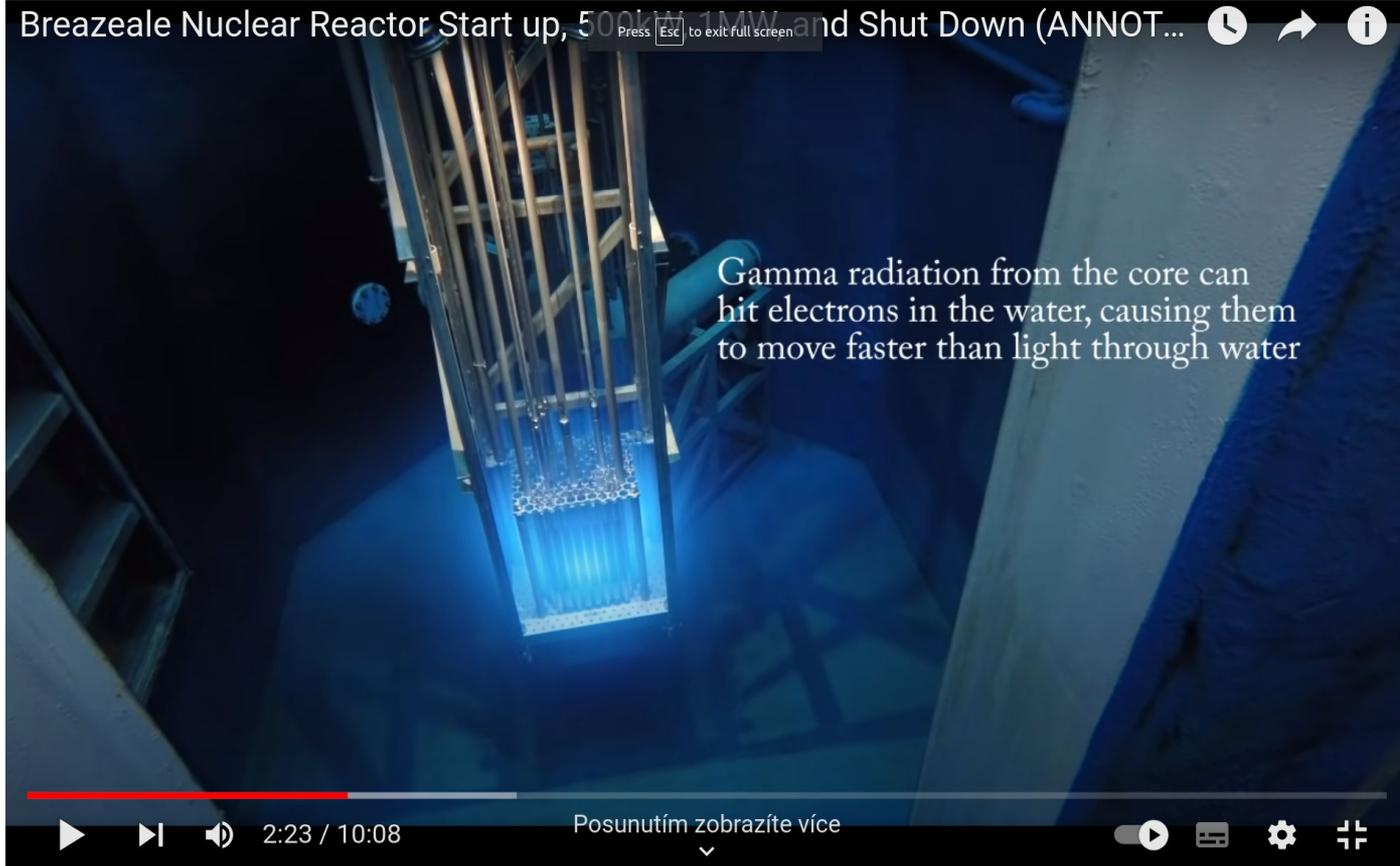
<https://www.eso.org/public/czechrepublic/images/eso1841i/>

# Čerenkovovo záření



[https://www.youtube.com/watch?v=uYrhWO\\_ZLYw&t=516s&ab\\_channel=AlexLandress](https://www.youtube.com/watch?v=uYrhWO_ZLYw&t=516s&ab_channel=AlexLandress)

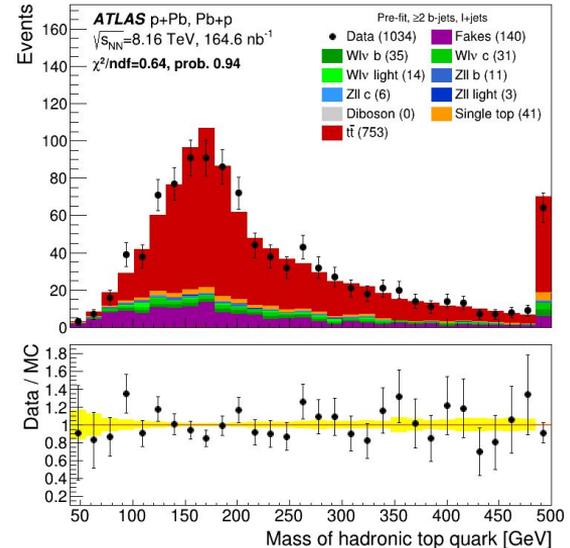
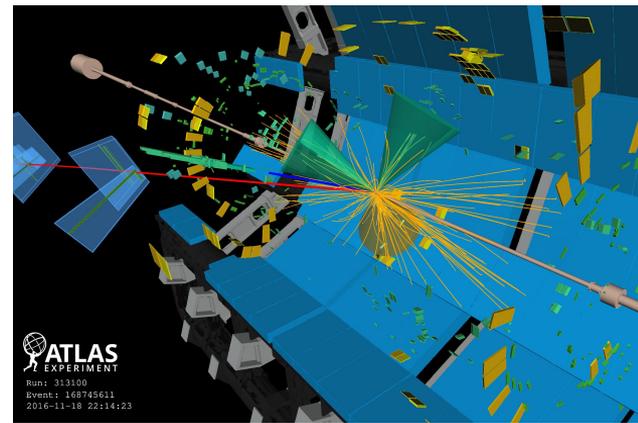
# Čerenkovovo záření



[https://www.youtube.com/watch?v=uYrhWO\\_ZLYw&t=516s&ab\\_channel=AlexLandress](https://www.youtube.com/watch?v=uYrhWO_ZLYw&t=516s&ab_channel=AlexLandress)

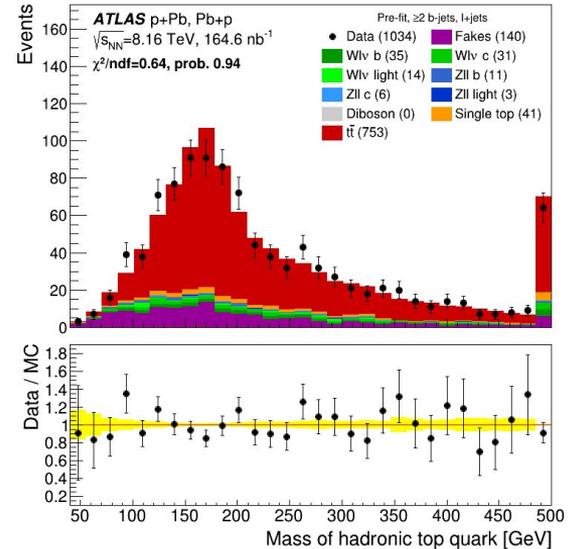
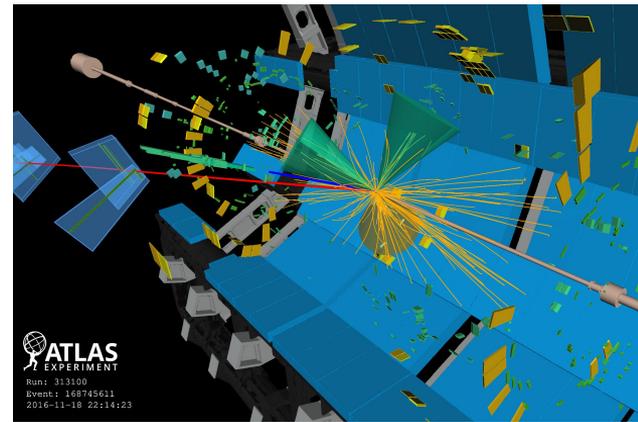
# A typical analysis

- Událost v detektoru obsahuje informace z jednotlivých poddetektorů
- Jsou zrekonstruovány a identifikovány fyzikální objekty
  - kandidáti na elektrony, miony, jety, gamma, tau, chybějící příčná energie...
- V rámci event selection jsou vybírány objekty na základě jejich požadovaného počtu a vlastností
  - podle očekávání toho, co by měl do detektoru přinášet signální proces
- Změřené spektrum musí být opraveno
  - kalibrace, linearita, rozlišení, efficiency (as function of...), data/simulation agreement



# A typical analysis

- Potřebuje
  - preselected data sample
  - data quality, luminosity measurement
  - objects calibrations, reconstruction algos
  - simulation samples
    - to optimize the analysis
    - estimate signal
    - estimate, understand and model backgrounds



# Invariantní hmota

$$E = mc^2$$

$$E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$E^2 = p^2 c^2 + m^2 c^4$$

$$m^2 = E^2 - p^2 \quad \text{v jednotkách } c=1$$

$$m^2 = \left( \sum_i E_i \right)^2 - \left( \sum_i \vec{p}_i \right)^2$$

$$E_i^2 = p_{ix}^2 + p_{iy}^2 + p_{iz}^2$$

$$E = \gamma mc^2$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \quad \beta \equiv \frac{v}{c}$$

$$E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$E = E_1 + E_2$$

$$\vec{p} = \vec{p}_1 + \vec{p}_2$$

$$\vec{p}^2 \equiv \vec{p} \cdot \vec{p} = \vec{p}_1^2 + 2\vec{p}_1 \cdot \vec{p}_2 + \vec{p}_2^2$$

$$m_1 \approx m_2 \approx 0 \Rightarrow$$

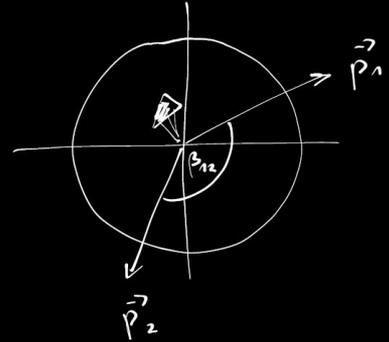
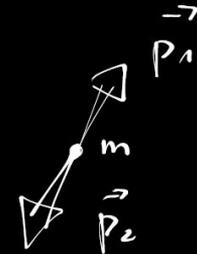
$$\doteq E_1^2 + 2\vec{p}_1 \cdot \vec{p}_2 + E_2^2$$

$$= E_1^2 + E_2^2 + 2E_1 E_2 \cos \beta_{12}$$

$$m^2 = (E_1 + E_2)^2 - (\vec{p}_1 + \vec{p}_2)^2$$

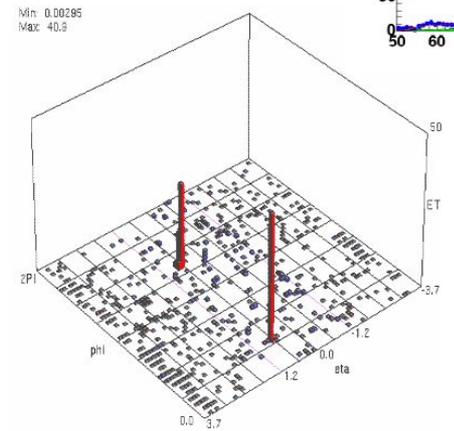
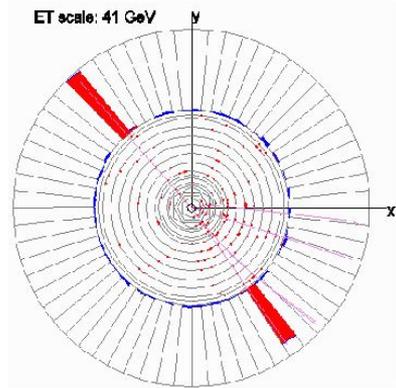
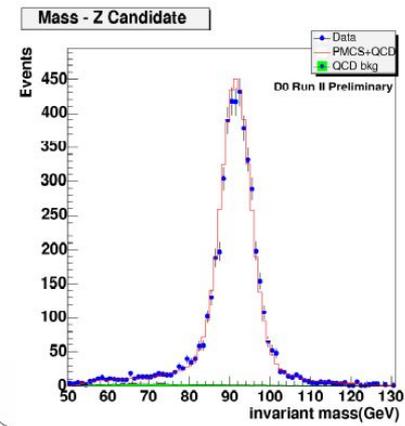
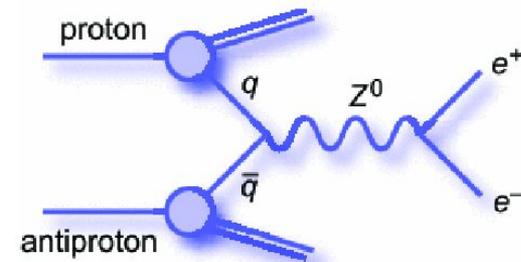
$$= \cancel{E_1^2} + 2E_1 E_2 + \cancel{E_2^2}$$

$$- \cancel{E_1^2} - \cancel{E_2^2} - 2E_1 E_2 \cos \beta_{12}$$

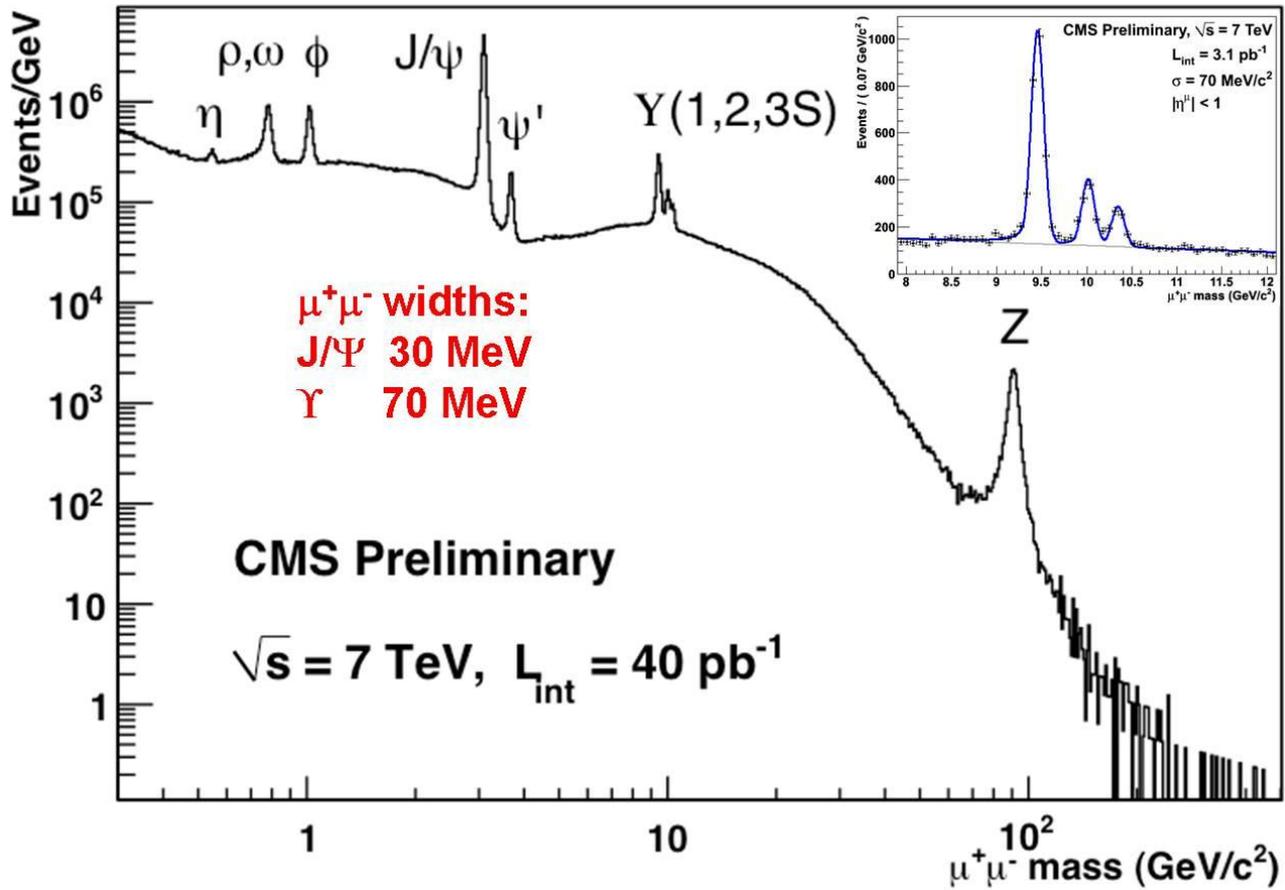


$$m^2 \approx 2E_1 E_2 (1 - \cos \beta_{12})$$

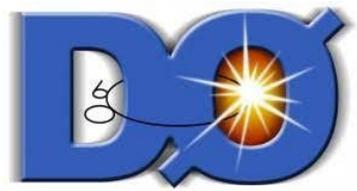
# Invariantní hmota



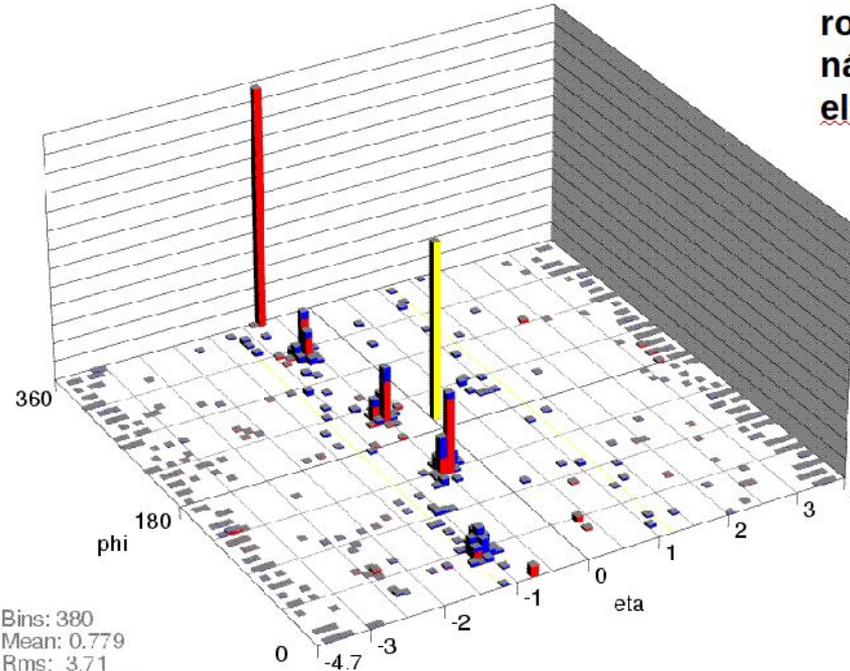
# Invariantní hmota



# Invariantní hmota



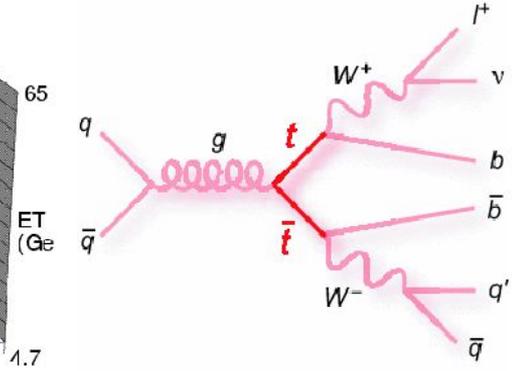
DZero experiment at Fermilab's Tevatron, a proton-antiproton collider.



Bins: 380  
 Mean: 0.779  
 Rms: 3.71  
 Min: 0.00966  
 Max: 63.9

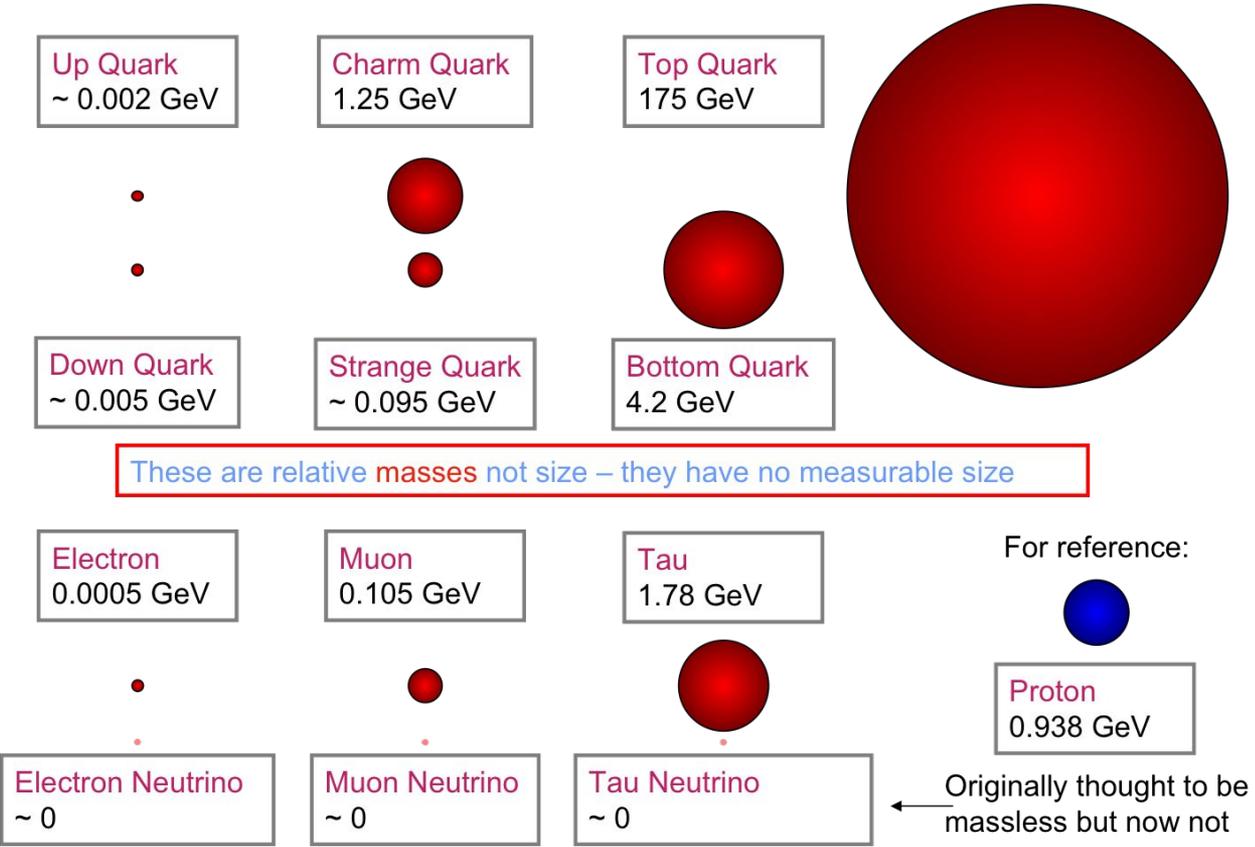
mE\_t: 47.2  
 phi\_t: 195 deg

Pár top-anti top rozpadající se na  $WbWb$  a následně na 4 jety, elektron a neutrino.



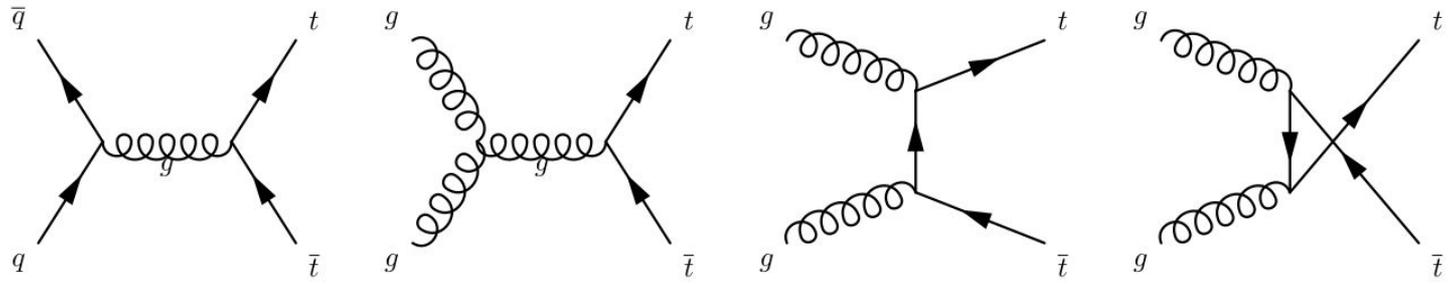
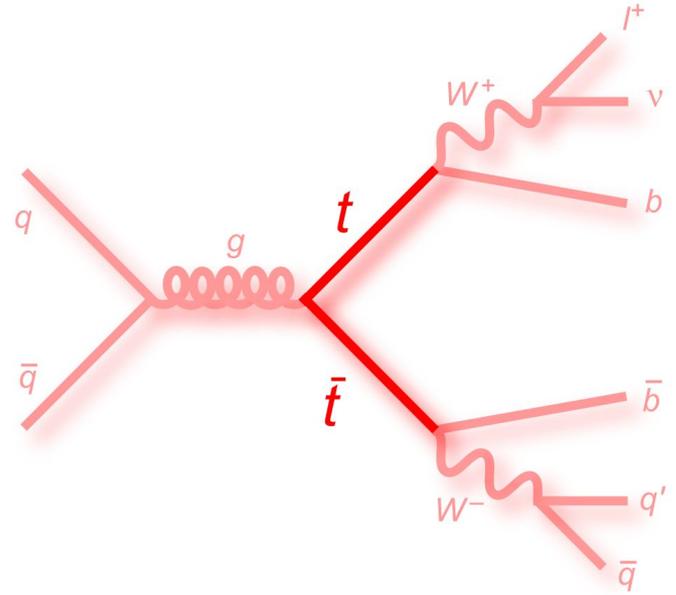
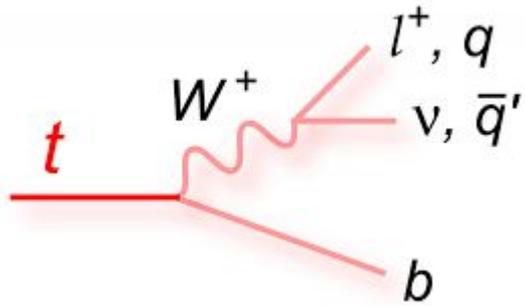
**Red:** EM Calorimeter  
**Blue:** Hadronic Calorimeter  
**Yellow:** Missing Energy (neutrino signature)

# Quarks, Leptons, Gauge Bosons and the BEH boson



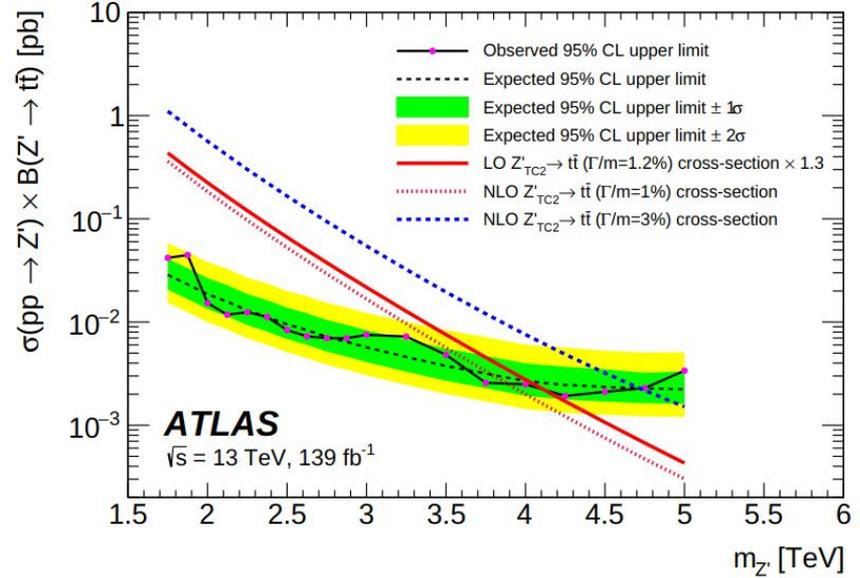
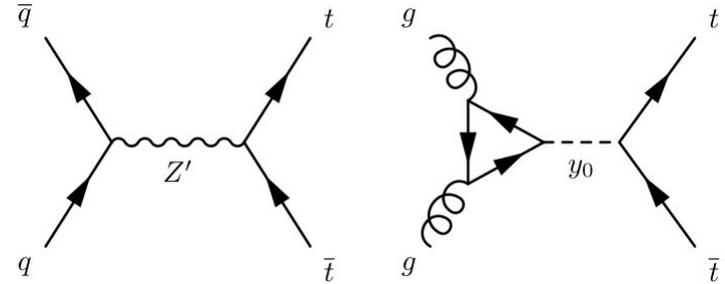
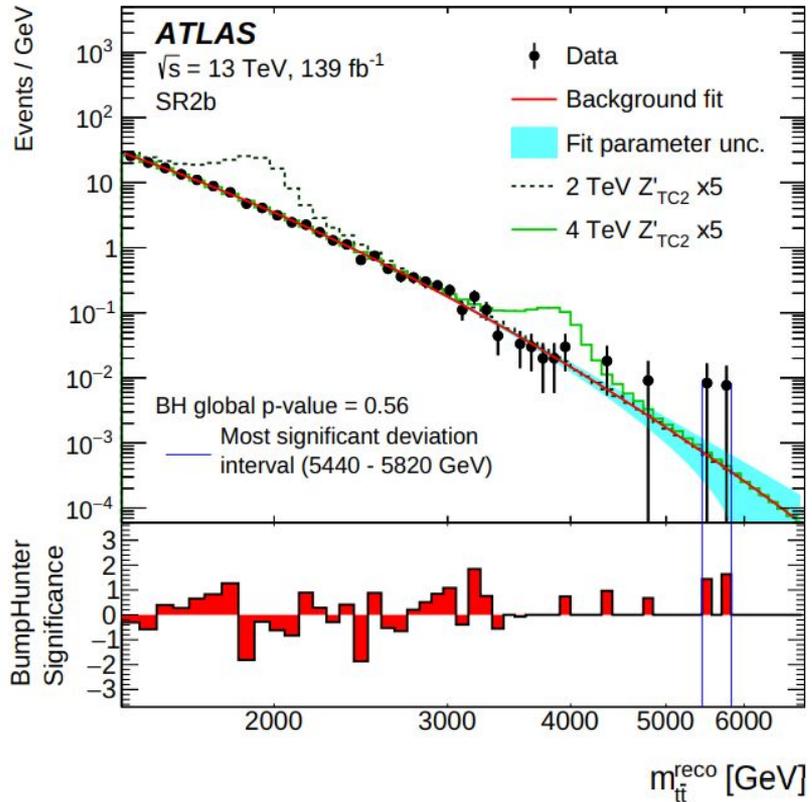
# Top quark

•



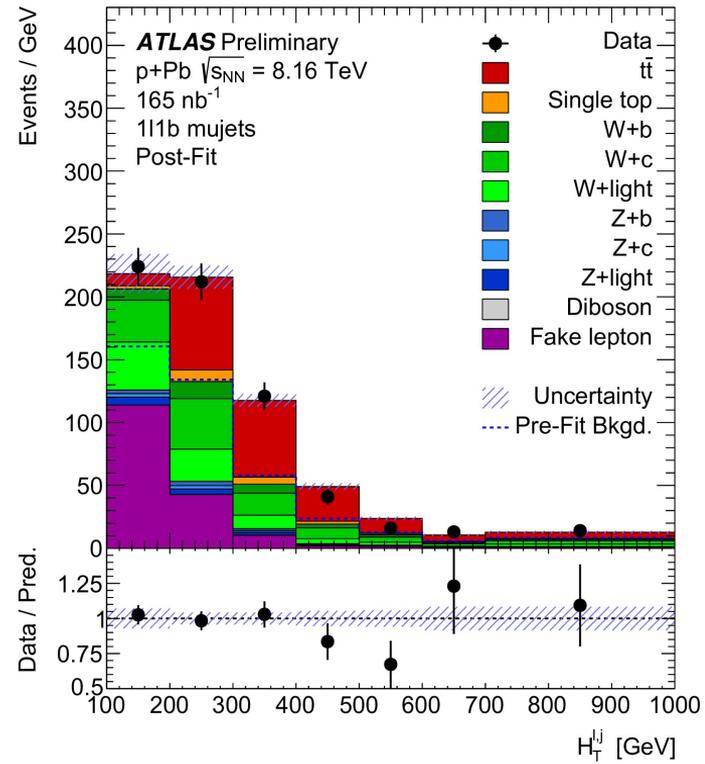
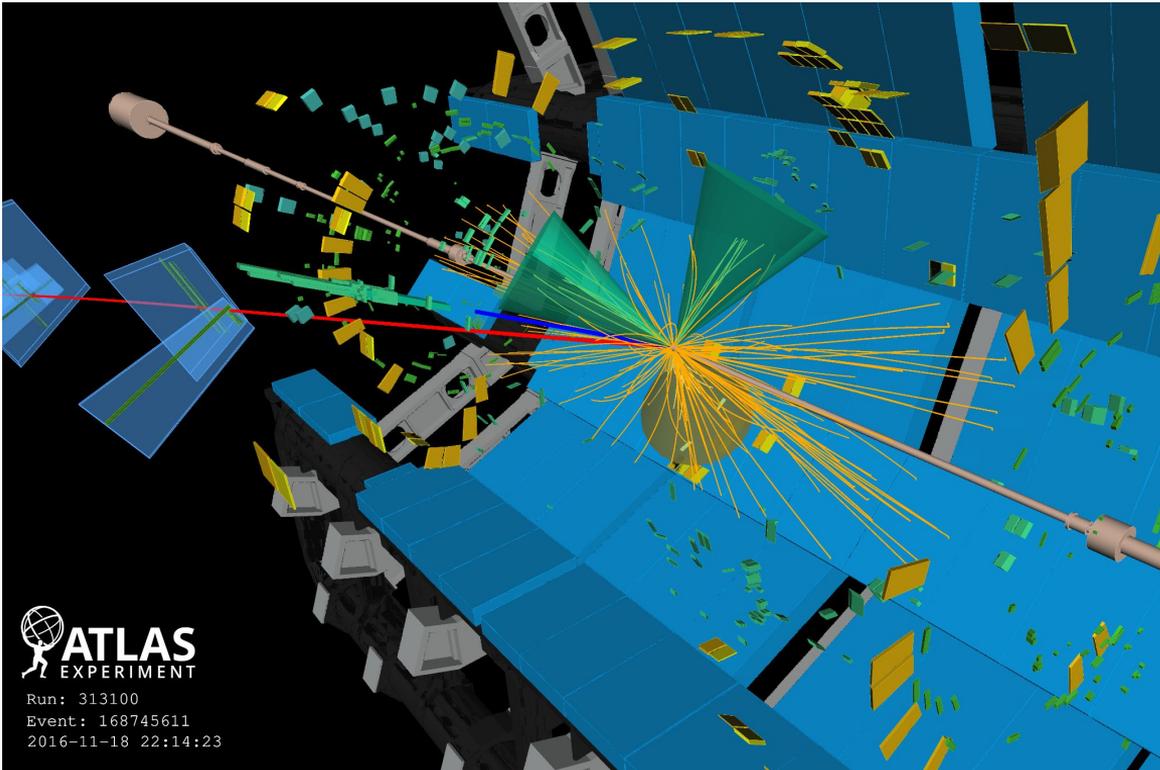
# Top quark

- $t\bar{t}$  resonances

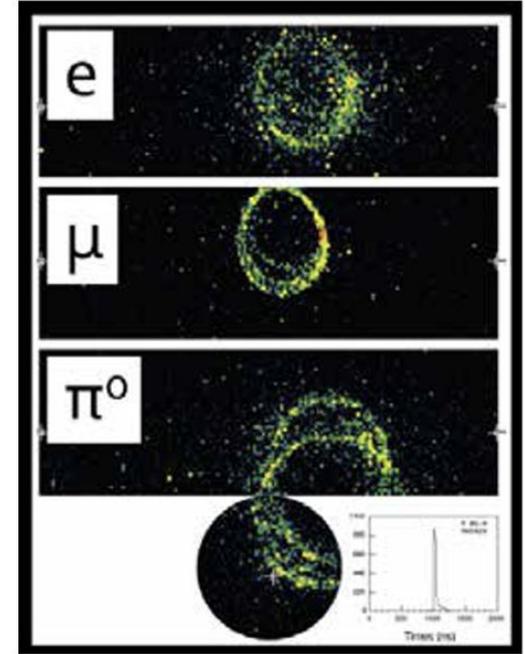
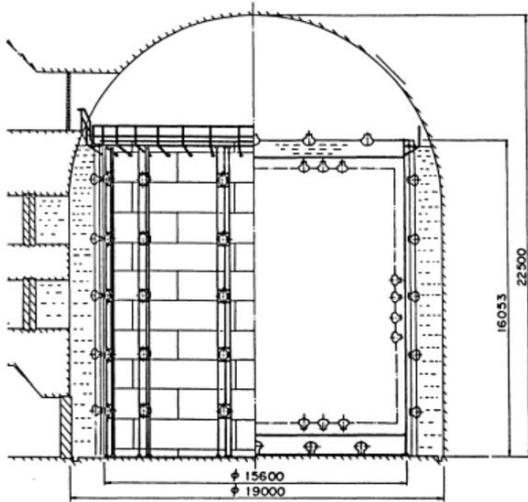


# A typical analysis

- <https://atlas.cern/Updates/Briefing/HI-Top-Observation>



# Experiment Kamiokande(-II) :: 3kt H<sub>2</sub>O



Neutrino identification

Figure 2: The detector of KAMIOKANDE -II. The dimensions are given in millimeters.

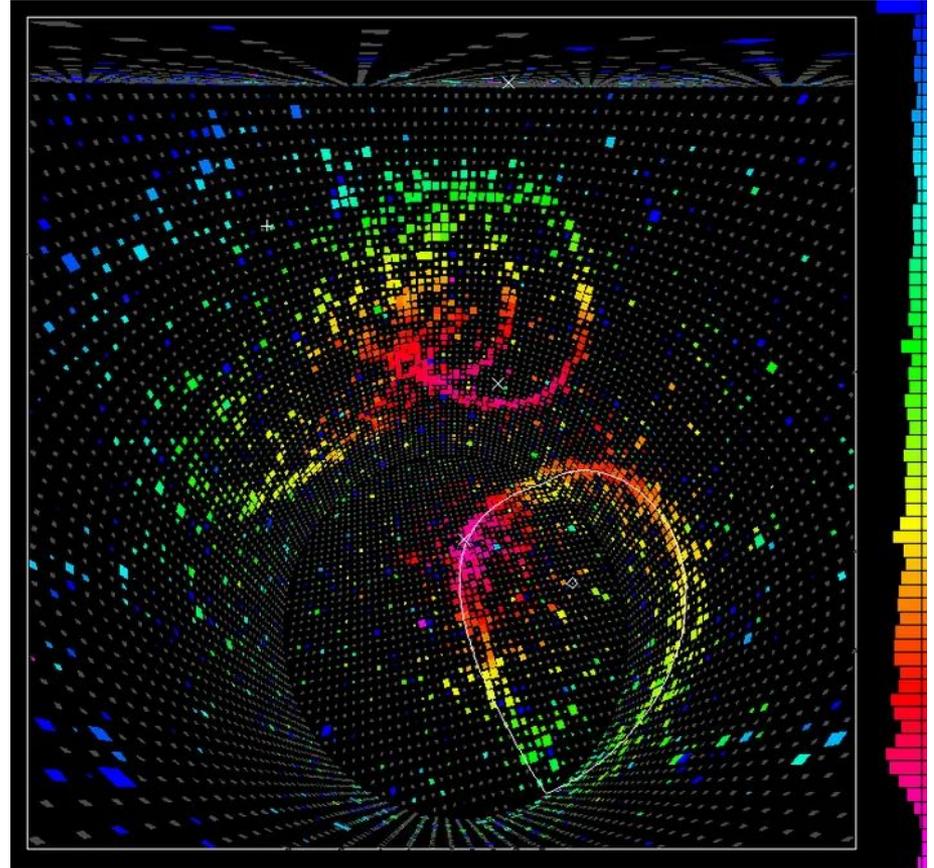
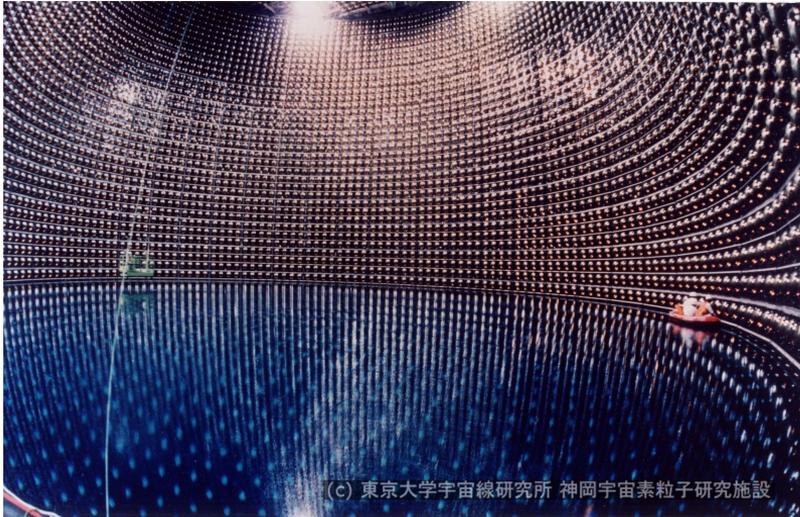
<https://articles.adsabs.harvard.edu//full/1987ESOC...26..219K/0000220.000.html>

# Super Kamiokande :: 50kt H<sub>2</sub>O

## Kamioka Nucleon Decay Experiment

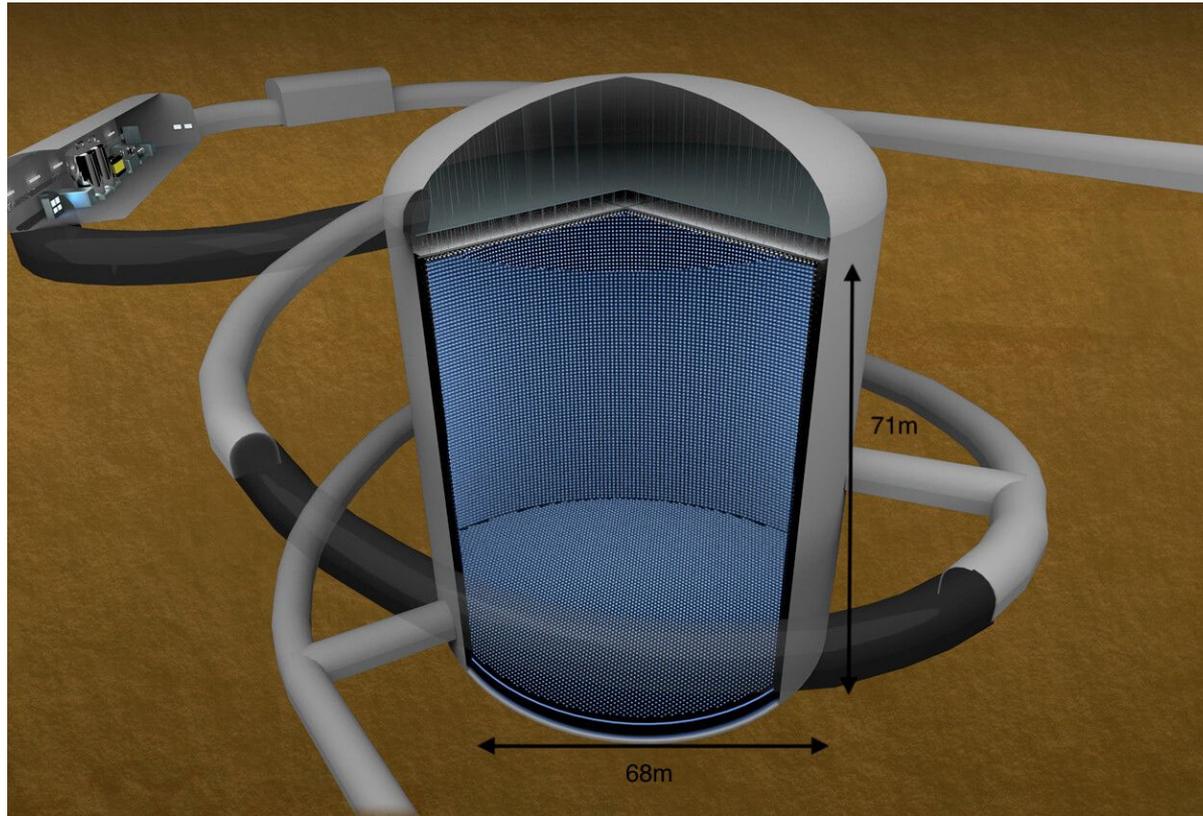
<https://www-sk.icrr.u-tokyo.ac.jp/en/sk/about/history/>

<https://www-sk.icrr.u-tokyo.ac.jp/en/sk/experience/gallery/>  
<https://www.forbes.com/sites/startswithabang/2018/12/05/is-there-really-a-fourth-neutrino-out-there-in-the-universe/>



# Hyper Kamiokande :: 250kt H<sub>2</sub>O

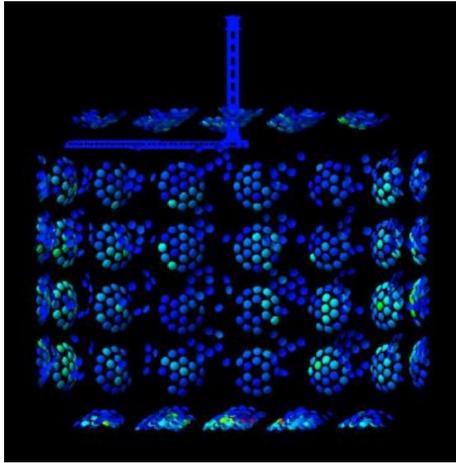
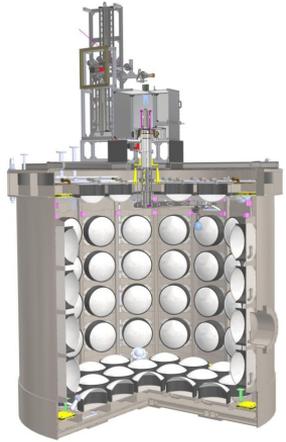
<https://www-sk.icrr.u-tokyo.ac.jp/en/hk/about/detector/>  
<https://www-sk.icrr.u-tokyo.ac.jp/en/hk/>



YT :: Hyper-Kamiokande  
[https://www.youtube.com/watch?v=JFOE3D2z7LM&t=12s&ab\\_channel=Hyper-Kamiokande](https://www.youtube.com/watch?v=JFOE3D2z7LM&t=12s&ab_channel=Hyper-Kamiokande)

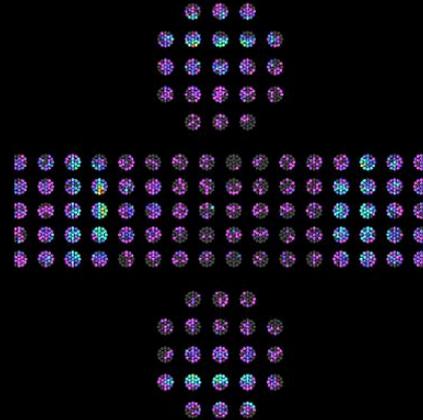
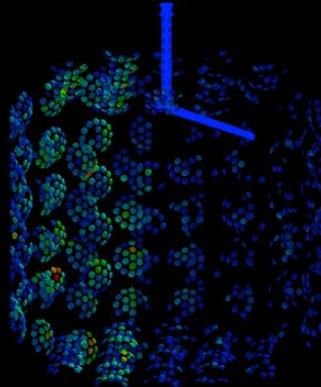
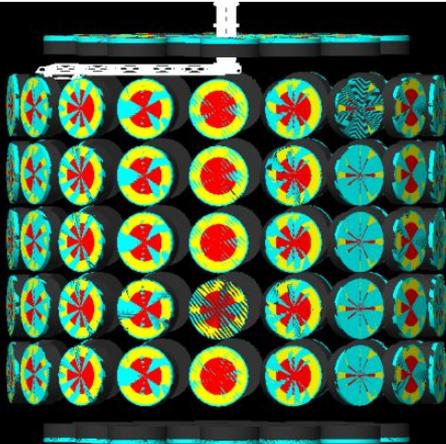


# Water Cherenkov Test Experiment :: 50t H<sub>2</sub>O @ CERN

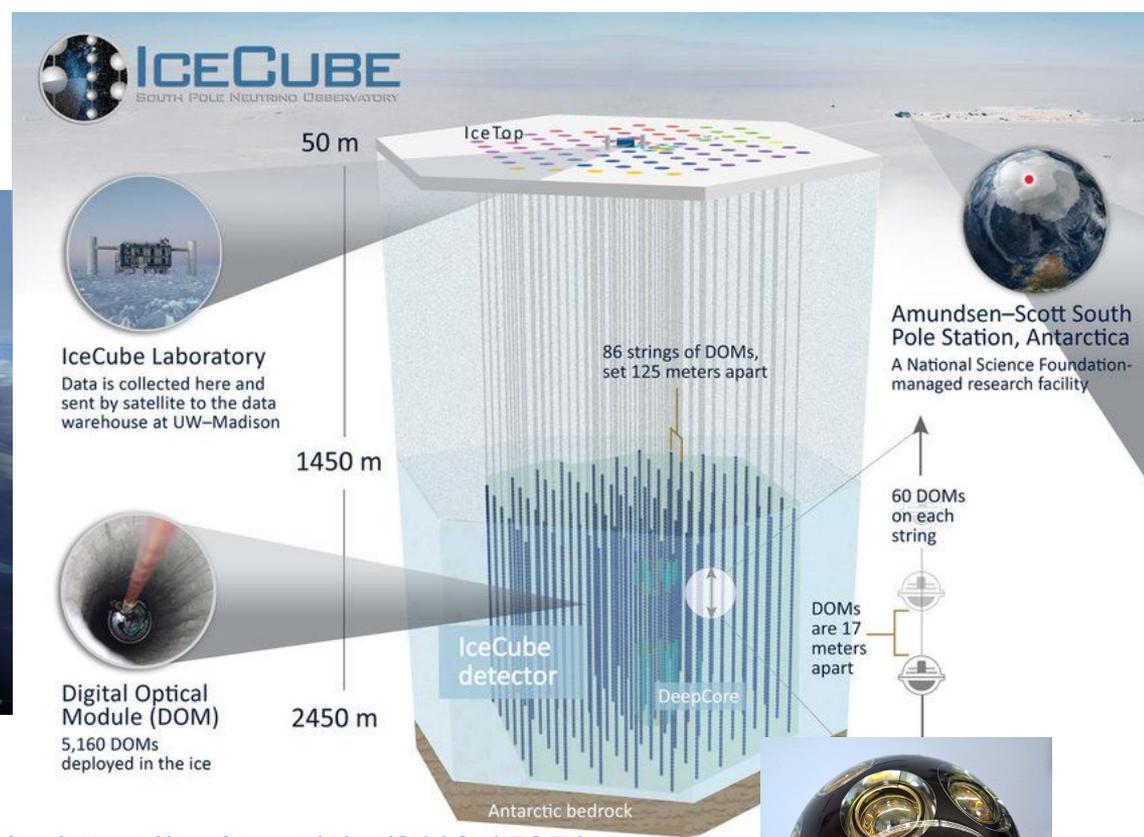
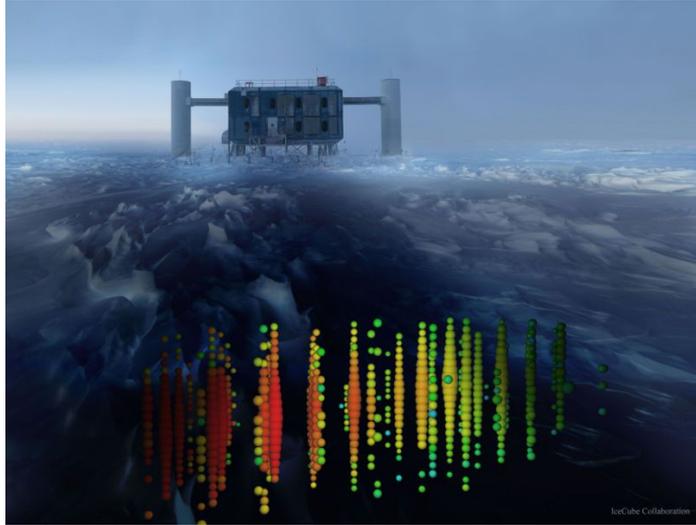


- Cíl: porozumět interakcím částic vznikajících v neutrinových reakcích.
- Pro kalibrace velkých neutrinových experimentů.

[https://indico.cern.ch/event/855372/contributions/4452155/attachments/2304577/3921520/WCTE@CERN\\_NuFACT-2021-09-08\\_LAnthony.pdf](https://indico.cern.ch/event/855372/contributions/4452155/attachments/2304577/3921520/WCTE@CERN_NuFACT-2021-09-08_LAnthony.pdf)



# IceCube



Francis Halzen

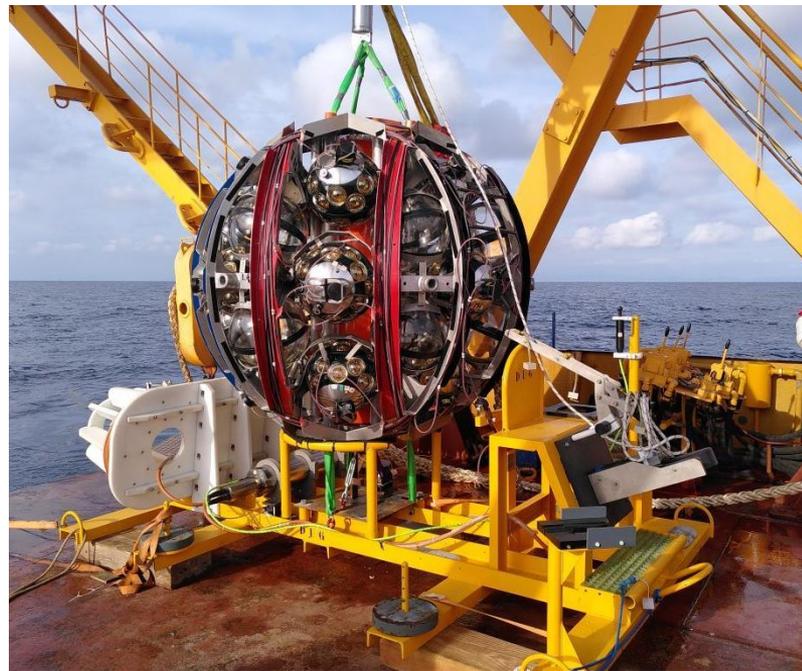
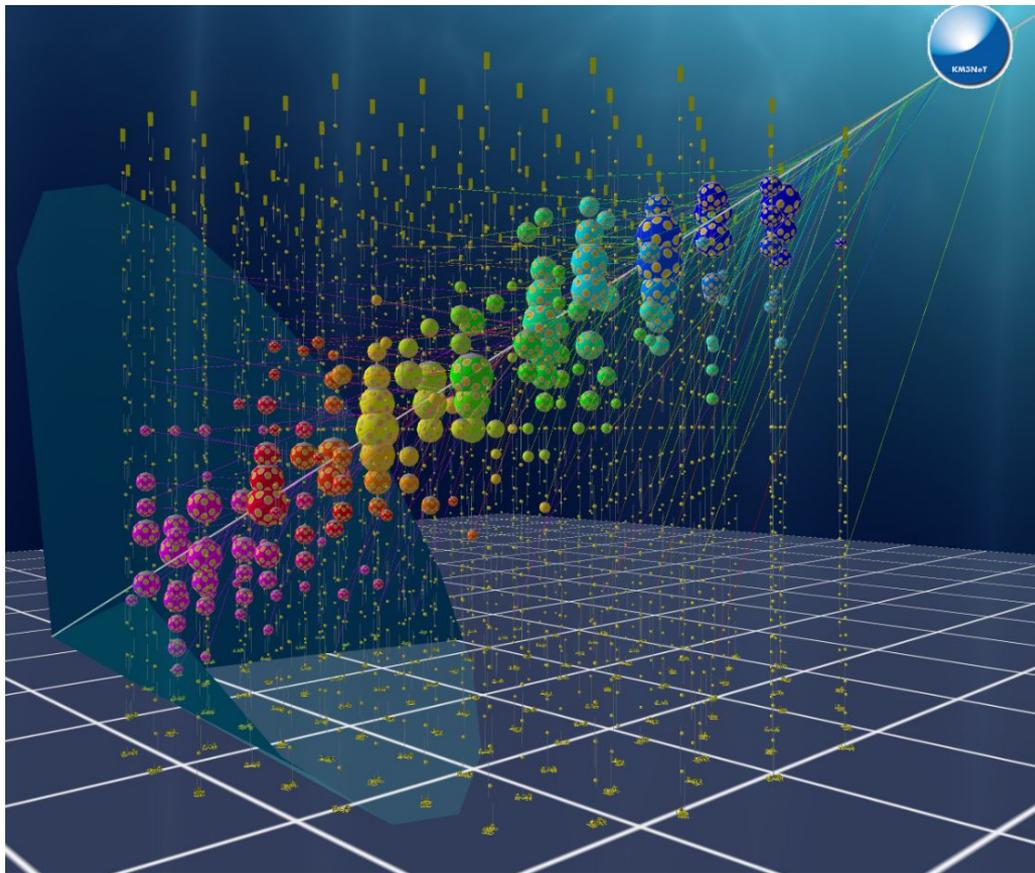
<http://jointlab.upol.cz/kvita/idpacs2022.html> :: <https://arxiv.org/abs/2110.15051>

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[https://www.researchgate.net/figure/Prototypes-of-an-mDOM-left-and-an-D-Egg-right-for-the-IceCube-Upgrade\\_fig8\\_351199668](https://www.researchgate.net/figure/Prototypes-of-an-mDOM-left-and-an-D-Egg-right-for-the-IceCube-Upgrade_fig8_351199668)

# KM3NeT



# Neutrino:)

Original - Photocopy of PLC 0373  
Abschrift/15.12.56 PW

Offener Brief an die Gruppe der Radioaktiven bei der  
Gauvereins-Tagung zu Tübingen.

Abschrift

Physikalisches Institut  
der Eidg. Technischen Hochschule  
Zürich

Zürich, 4. Des. 1930  
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich huldvollst  
ansuhören bitte, Ihnen das Näheres auseinandersetzen wird, bin ich  
angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie  
des kontinuierlichen beta-Spektrums auf einen verzweigten Ausweg  
verfallen um den "Wechselgatz" (1) der Statistik und den Energiegatz  
zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale  
Teilchen, die ich Neutronen nennen will, in den Kernen existieren,  
welche dem Spin 1/2 haben und das Ausschliessungsprinzip befolgen und  
sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie  
nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen  
müsste von derselben Grössenordnung wie die Elektronenmasse sein und  
sogarfalls nicht grösser als 0,01 Protonenmasse. Das kontinuierliche  
beta-Spektrum wäre dann verständlich unter der Annahme, dass beim  
beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert  
wird, derart, dass die Summe der Energien von Neutron und Elektron  
konstant ist.

Nun handelt es sich weiter darum, welche Kräfte auf die  
Neutronen wirken. Das wahrscheinlichste Modell für das Neutron scheint  
mir aus wellenmechanischen Gründen (näheres weiss der Ueberbringer  
dieser Zeilen) dieses zu sein, dass das ruhende Neutron ein  
magnetischer Dipol von einem gewissen Moment  $\mu$  ist. Die Experimente  
verleihen wohl, dass die ionisierende Wirkung eines solchen Neutrons  
nicht grösser sein kann, als die eines gamma-Strahls und darf dann  
 $\mu$  wohl nicht grösser sein als  $e \cdot (10^{-13} \text{ cm})$ .

Ich traue mich vorläufig aber nicht, etwas über diese Idee  
zu publizieren und wende mich erst vertrauensvoll an Buch, liebe  
Radioaktive, mit der Frage, wie es um den experimentellen Nachweis  
eines solchen Neutrons stände, wenn dieses ein ebensolches oder etwa  
10mal grösseres Durchdringungsvermögen besitzen würde, wie ein  
gamma-Strahl.

Ich gebe zu, dass mein Ausweg vielleicht von vornherein  
wenig wahrscheinlich erscheinen wird, weil man die Neutronen, wenn  
sie existieren, wohl schon längst gesehen hätte. Aber mir wer sagt,  
genügt und der Ernst der Situation beim kontinuierlichen beta-Spektrum  
wird durch einen Ausspruch meines verehrten Vorgängers im Amt,  
Herrn Debye, beleuchtet, der mir kürzlich in Brüssel gesagt hat:  
"O, daran soll man am besten gar nicht denken, sowie an die neuen  
Steuern." Darum soll man jeden Weg zur Rettung ernstlich diskutieren.-  
Also, liebe Radioaktive, prüfet, und richtet.- Leider kann ich nicht  
persönlich in Tübingen erscheinen, da ich infolge eines in der Nacht  
vom 6. zum 7. Des. in Zürich stattfindenden Balles hier unabkömmlich  
bin.- Mit vielen Grüssen an Buch, sowie an Herrn Baek, Buer  
untertänigster Diener

ges. W. Pauli



Wolfgang Pauli

# SN1987A 24.2.1987 :: 170kly



[https://en.wikipedia.org/wiki/SN\\_1987A](https://en.wikipedia.org/wiki/SN_1987A)

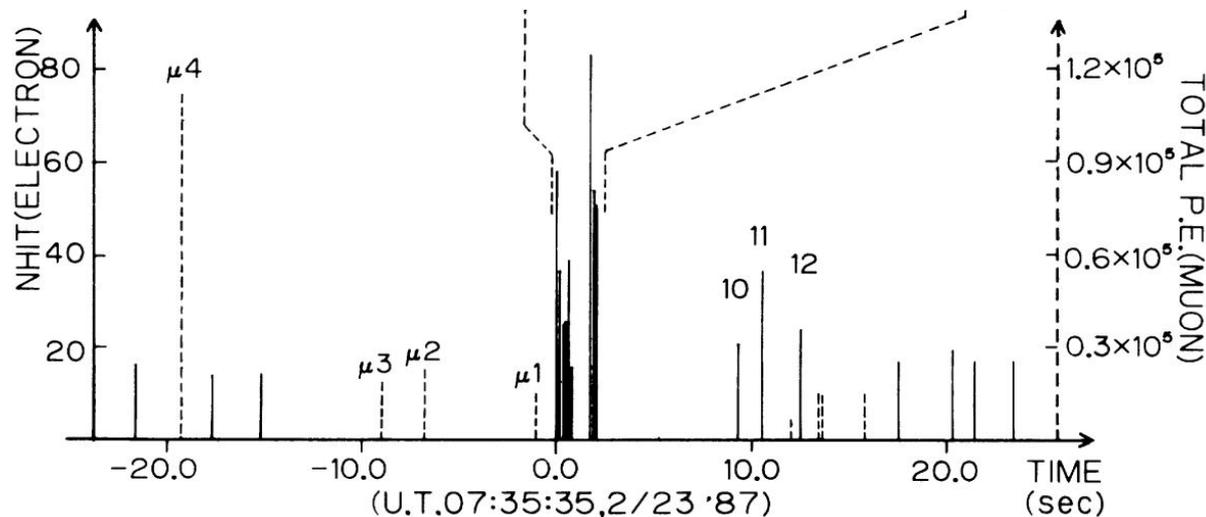


<https://www-sk.icrr.u-tokyo.ac.jp/en/news/detail/324>

## Observation of a Neutrino Burst from the Supernova SN1987A

K. Hirata,<sup>(a)</sup> T. Kajita,<sup>(a)</sup> **M. Koshiba,**<sup>(a,b)</sup> M. Nakahata,<sup>(b)</sup> Y. Oyama,<sup>(b)</sup>  
N. Sato,<sup>(c)</sup> A. Suzuki,<sup>(b)</sup> M. Takita,<sup>(b)</sup> and Y. Totsuka<sup>(a,c)</sup>

*University of Tokyo, Tokyo 113, Japan*



# Experiment Kamiokande(-II) :: 3kt H<sub>2</sub>O

- Počátky neutrinové astronomie! :-)
- 1 Erg is a unit of energy equal to  $10^{-7}$  Joules (100 nJ)

form volume distribution. Additional support is provided by the correlation in angle of the first two observed events with the direction to SN1987A. The event burst occurred roughly 18 h prior to the first optical sighting.<sup>1</sup>

Correcting for energy-dependent detection efficiency, and assuming that nine of the twelve events are due to  $\bar{\nu}_e p \rightarrow e^+ n$ , we obtain an integral flux of  $1.0 \times 10^{10} \bar{\nu}_e \text{ cm}^{-2}$  for the burst, where the  $\bar{\nu}_e$  energy (the observed electron energy plus 1.3 MeV) is above 8.8 MeV. This, in turn, leads to the  $\bar{\nu}_e$  output of SN1987A of  $8 \times 10^{52}$  ergs for an assumed average energy of 15 MeV.

This observation is the first direct observation in neutrino astronomy, and coincides remarkably well with the current model of supernova collapse and neutron-star formation.<sup>6</sup> In that model an aged, massive star, having exhausted its nuclear fuel, undergoes a supernova explosion.

In supernovae of Type II almost all of the gravitational binding energy of the resultant neutron star,  $\sim 3 \times 10^{53}$  ergs, is radiated within a few seconds in the form of  $10^{58}$  neutrinos of all flavors with average energy in the vicinity of 10–15 MeV.