

Od Newtona ke Schrödingerovi aneb Jak spočítat tvar atomu





a photograph of isaac newton sitting in grass under an apple tree with a single **red** apple falling on his head, with dark sky full of huge planets

Midjourney

Newtonovy zákony

- Newtonovy pohybové zákony
 - Nobelova cena 1687

$$\vec{F} = m \frac{d^2}{dt^2} \vec{x}(t)$$

$$\vec{F} = \frac{d\vec{p}}{dt} = \frac{d}{dt} (m\vec{v})$$

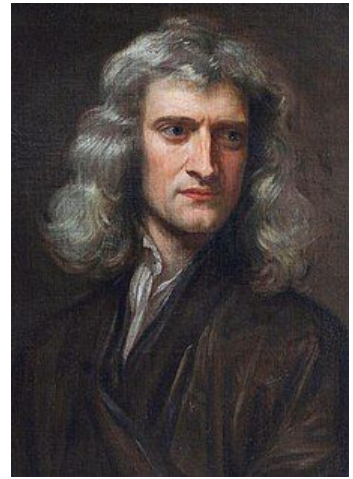


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 - Nobelova cena 1687
- A co z nich?
 - Předpověď trajektorie! Lze ověřit, aplikovat.
 - Plynou z nich např. Keplerovy “zákony”.
- Současně: definice síly i hmotnosti.
- Zajímavost: hmotnost setrvačná a gravitační jsou stejné:)



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- **Univerzalita gravitace** (jablko, planety)!

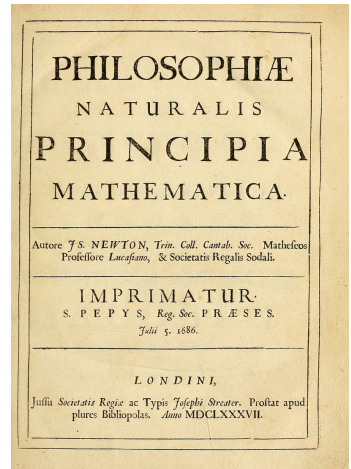
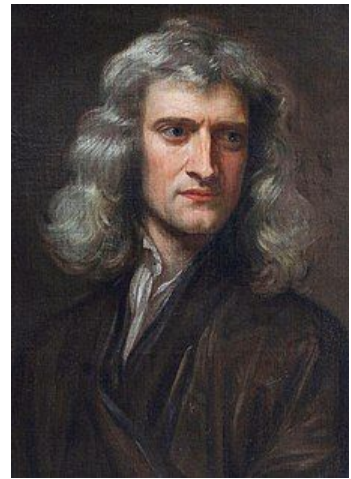
$$F_G = \gamma \frac{m_1 m_2}{|\vec{r}_1 - \vec{r}_2|^2}$$

$$V_G = -\gamma \frac{m_1 m_2}{|\vec{r}_1 - \vec{r}_2|}$$

$$\vec{x}(t) = \begin{pmatrix} x(t) \\ y(t) \\ z(t) \end{pmatrix}$$

$$F = -\text{grad} V$$

$$F = -\nabla V$$





a photograph of isaac newton sitting in grass under an apple tree with a **read** apple falling on his head, with dark sky full of huge planets

Midjourney



a photograph of isaac newton sitting in grass under an apple tree with a **single red** apple falling on his head, with large dark sky with huge planet Saturn on it

Midjourney



red apple falling on the head of isaac newton standing on an asteroid with planet saturn and its rings in the background

Midjourney



Midjourney

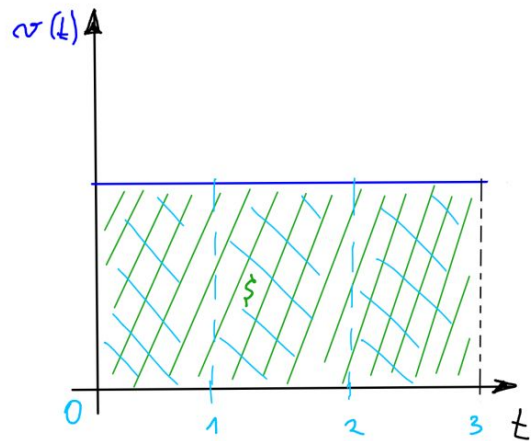
Matematika Newtonových zákonů?

$$\vec{F} = m \frac{d^2}{dt^2} \vec{x}(t)$$

$$\vec{F} = \frac{d\vec{p}}{dt} = \frac{d}{dt} (m \vec{v})$$

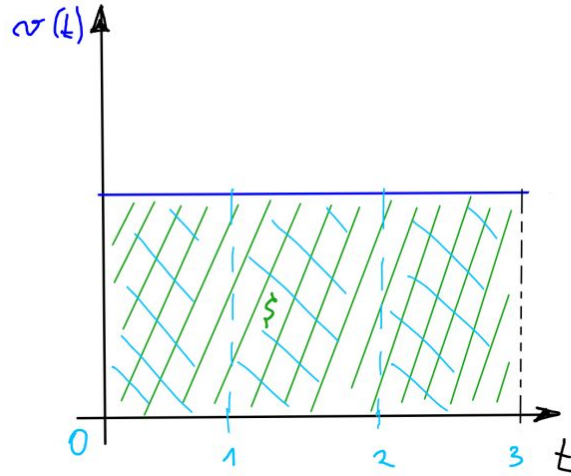


Dráha a rychlost :: rovnoměrný pohyb



$$\xi = v \cdot t$$

Dráha a rychlost :: rovnoměrný pohyb

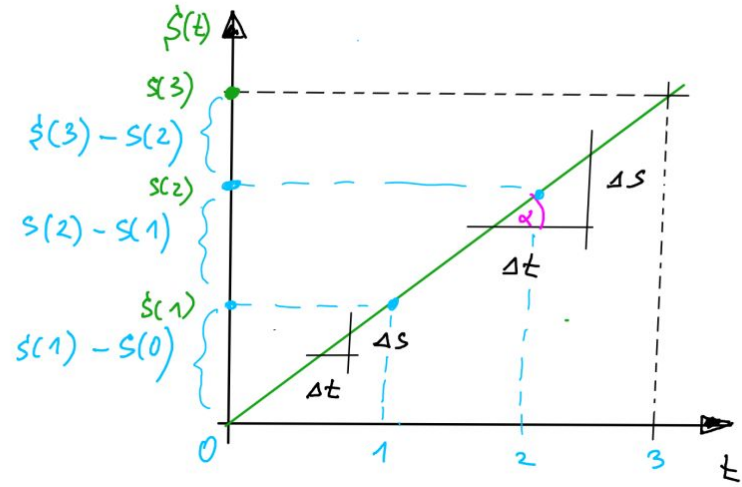


$$\zeta = v \cdot t$$

Rovnoměrný pohyb

Uražená dráha

Rychlost

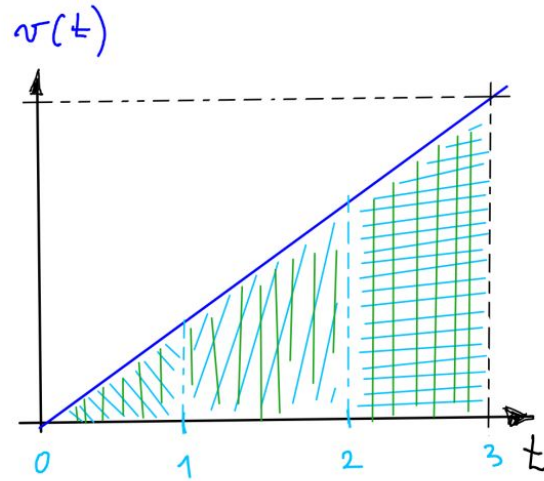


průměrná $v = \frac{\Delta s}{\Delta t} = \text{tg } \alpha$

okamžitá $v = \frac{ds}{dt} = \dot{s}$

přesněji $v(t) = \frac{ds(t)}{dt} = \dot{s}(t)$

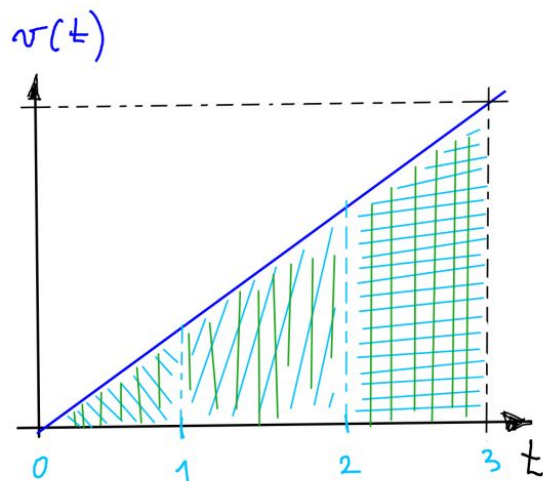
Dráha a rychlost :: volný pád



$$v = g \cdot t$$

Volný pád

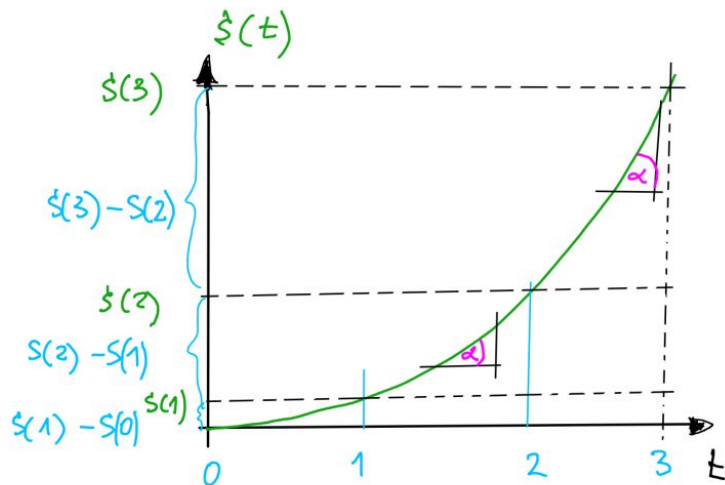
Dráha a rychlost :: volný pád



$$v = g \cdot t \quad \text{Volný pád}$$

$$a = \frac{\Delta v}{\Delta t} \quad \text{zrychlení}$$

$$a(t) = \frac{dv(t)}{dt}$$



celková dráha: plocha

$$s = \frac{1}{2} v \cdot t = \frac{1}{2} g \cdot t \cdot t$$

$$s = \frac{1}{2} g t^2$$

$$v(t) = \frac{ds(t)}{dt} = \dot{s}(t)$$

Rychlost → dráha: Integrace

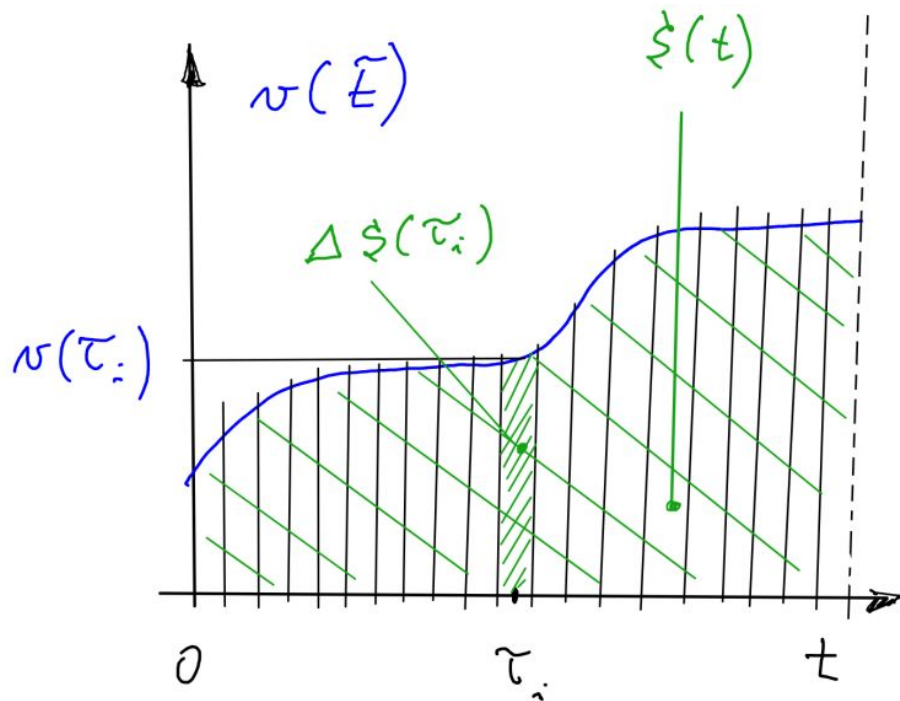
$$\Delta s = v(\tilde{t}) \Delta \tilde{t}$$

$$\Delta s_i = v(\tilde{t}_i) \Delta \tilde{t}$$

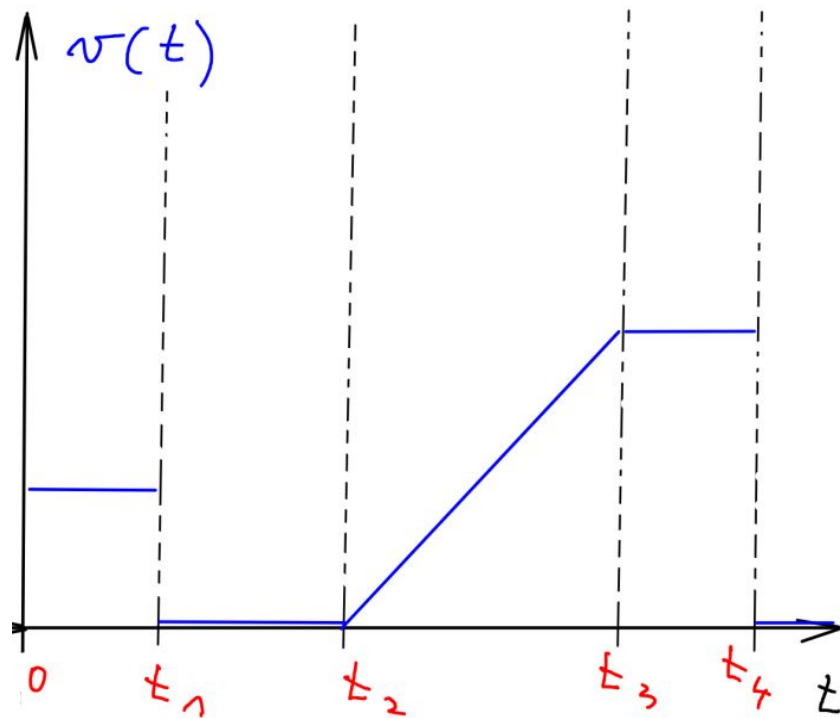
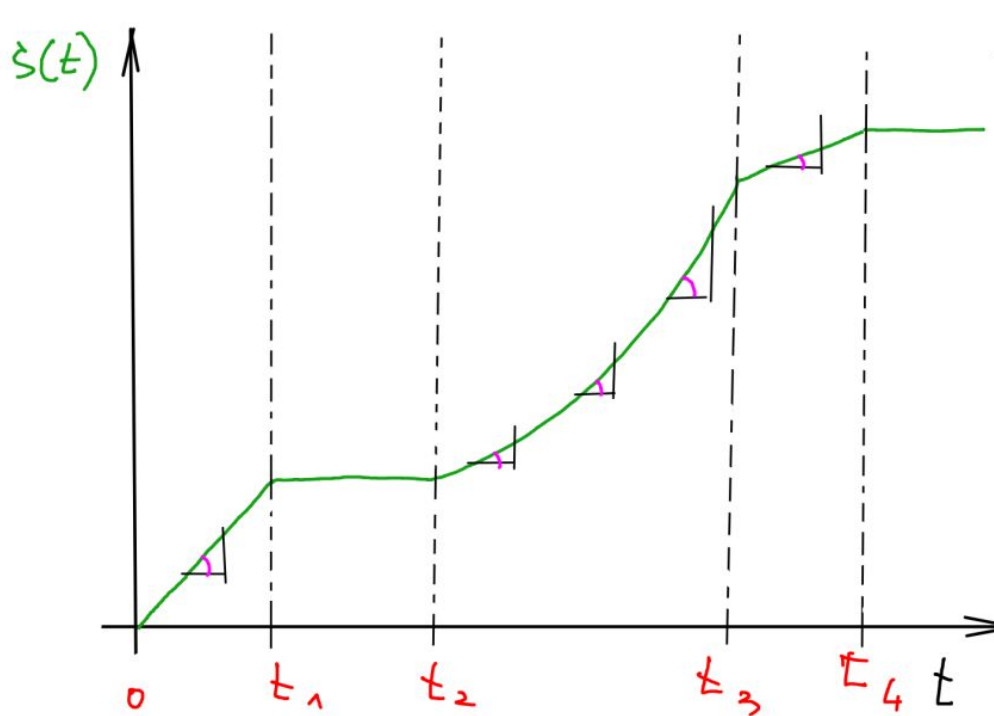
$$s = \sum_i \Delta s_i = \sum_i v(\tilde{t}_i) \Delta \tilde{t}_i$$

$$ds = v(\tilde{t}) d\tilde{t}$$

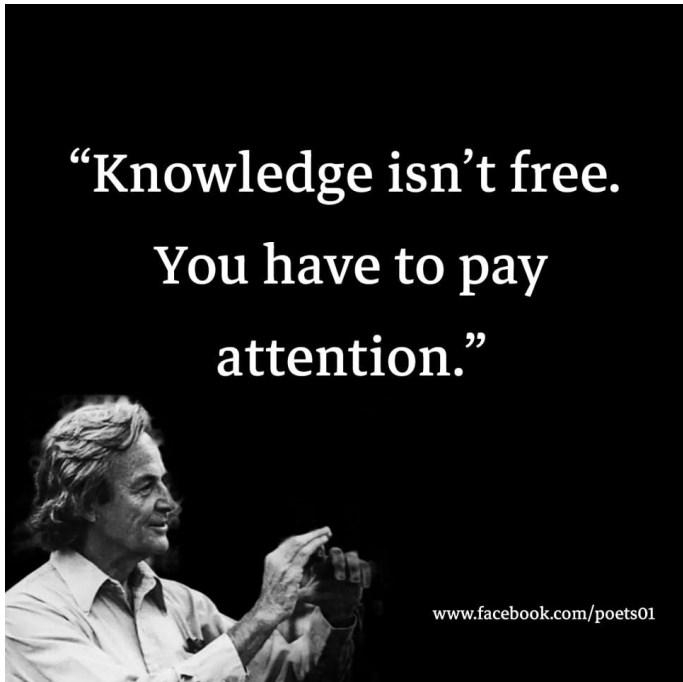
$$s(t) = \int_0^t v(\tilde{t}) d\tilde{t}$$



Dráha → Rychlost: Derivace



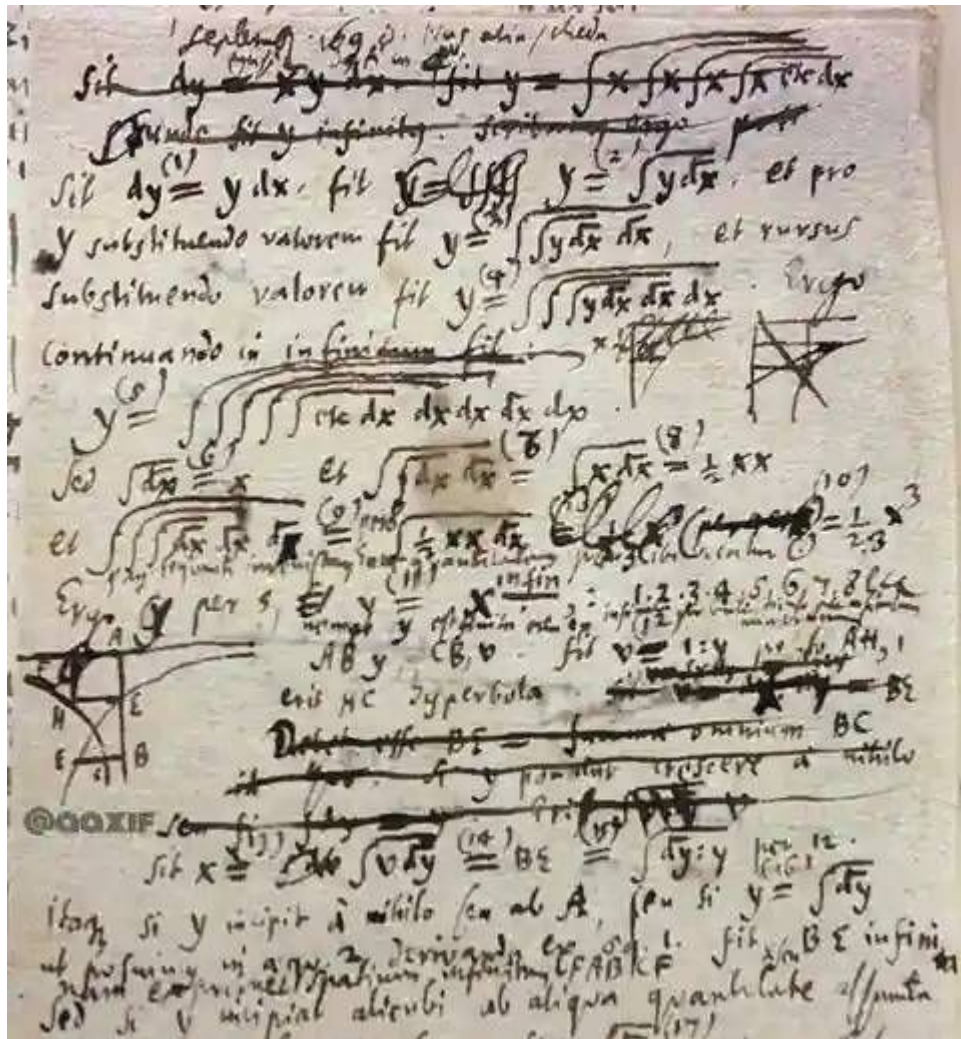
Integrály



“Knowledge isn’t free.
You have to pay
attention.”

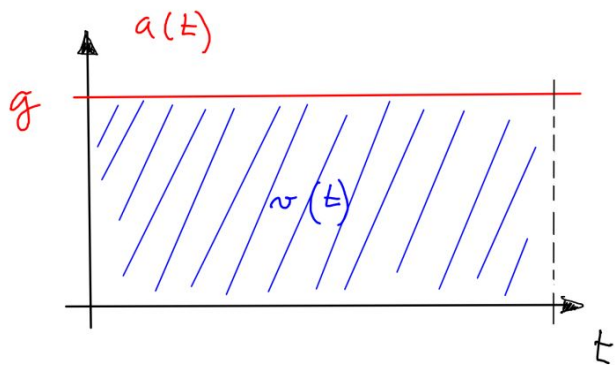
www.facebook.com/poets01

Leibnitz



Zrychlení :: volný pád

-



$$v(t) = \int_0^t a(\tilde{t}) d\tilde{t}$$

$$a(t) = \dot{v}(t) = \ddot{s}(t)$$

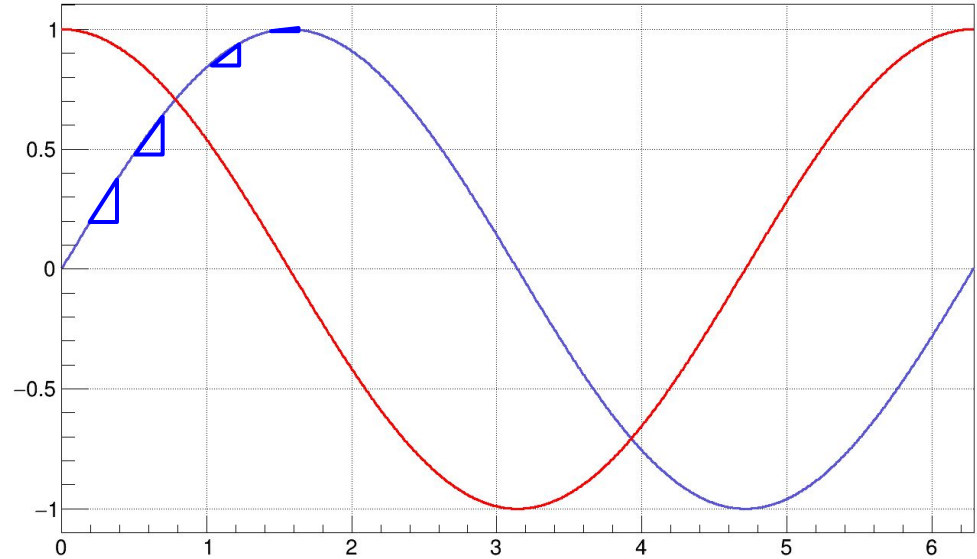
$$a(t) = \frac{dv}{dt} = \frac{d^2s}{dt^2}$$

$$\vec{F} = m \frac{d^2}{dt^2} \vec{x}(t)$$

$$\vec{F} = \frac{d\vec{p}}{dt} = \frac{d}{dt} (m\vec{v})$$

Sinus a Cosinus

$$\frac{d}{dx} \sin x = \cos x$$

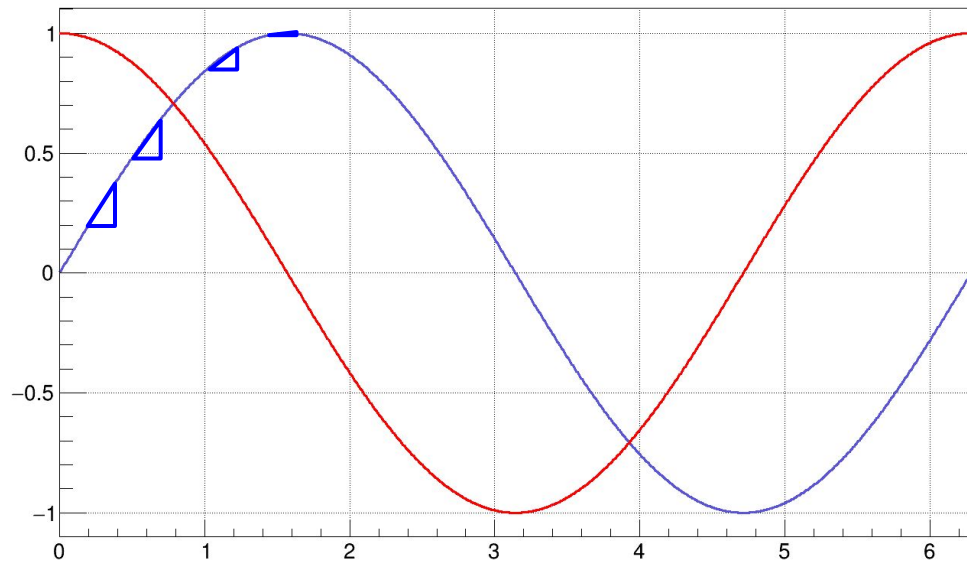


Sinus a Cosinus

- první derivace je rychlost růstu funkce

$$\frac{d}{dx} \sin x = \cos x$$

$$\frac{d}{dx} \cos x = -\sin x$$



Sinus a Cosinus

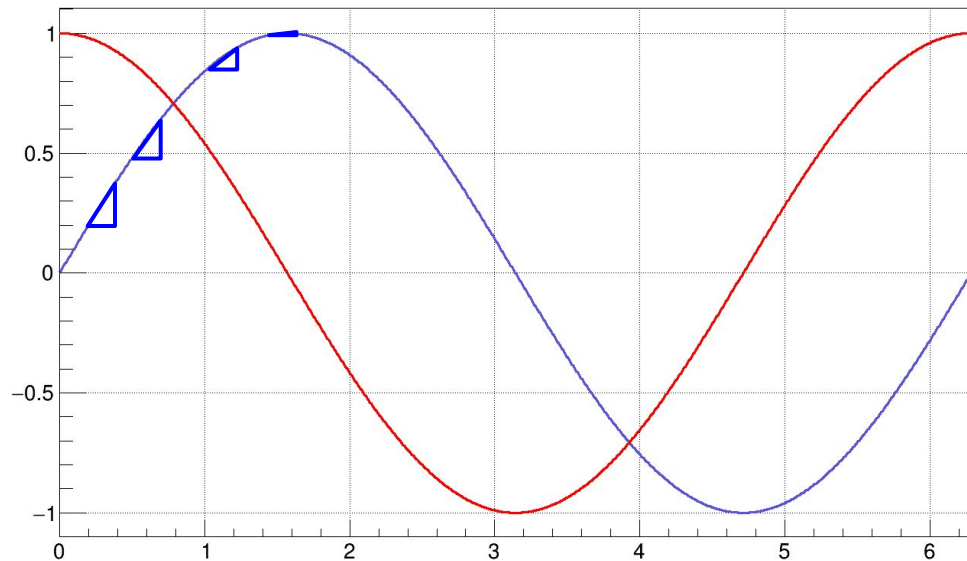
- druhá derivace je křivost funkce

$$\frac{d}{dx} \sin x = \cos x$$

$$\frac{d}{dx} \cos x = -\sin x$$

$$\frac{d^2}{dx^2} \sin x = -\sin x$$

$$\frac{d^2}{dx^2} \cos x = -\cos x$$



Sinus a Cosinus

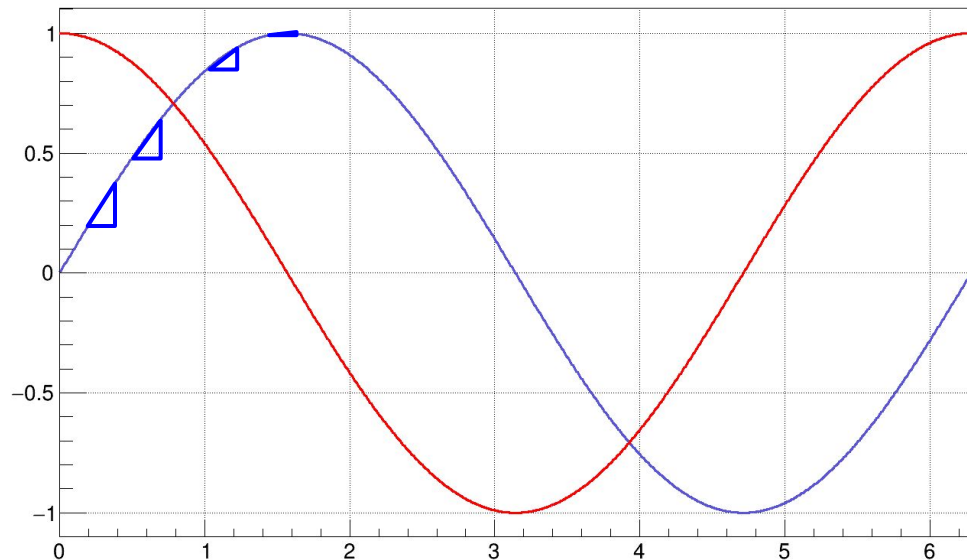


$$(\sin x)' = \cos x$$

$$(\cos x)' = -\sin x$$

$$(\sin x)'' = -\sin x$$

$$(\cos x)'' = -\cos x$$



Sinus a Cosinus

$$(\sin kx)' = k \cos kx$$

$$(\cos kx)' = -k \sin kx$$

$$(e^x)' = e^x$$

$$(e^{ax})' = a e^{ax}$$

$$(e^{ikx})' = ik e^{ikx} \quad (e^{ikx})'' = -k^2 e^{ikx}$$

Vlna

- Zobecnění / Povýšení pozorování na princip
- Necht' diferenciální rovnice určuje řešení, kterému budeme říkat vlna

$$\frac{d^2}{dx^2} f(x) = -f(x)$$

$$\Rightarrow f(x) = A \sin x + B \cos x$$

$$f''(x) = -f(x)$$

Vlna

Postupující harmonická vlna

Např. $f(x,t) = A \cos[k(x-vt)]$

$$\frac{\partial f}{\partial x} = -kA \sin[k(x-vt)]$$

$$\frac{\partial^2 f}{\partial x^2} = -k^2 A \cos[k(x-vt)] = -k^2 f \quad \nabla \quad \therefore)$$

$$\frac{\partial f}{\partial t} = -kvA \sin[k(x-vt)]$$

$$\frac{\partial^2 f}{\partial t^2} = -k^2 v^2 A \cos[k(x-vt)] = -k^2 v^2 f \quad \nabla \quad \therefore)$$

Vlnová rovnice

$$1D: \left(\frac{1}{v^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} \right) f(x, t) = 0$$

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$$1D: \frac{d^2}{dx^2} \quad 3D: \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \equiv \Delta$$

Vlnová rovnice

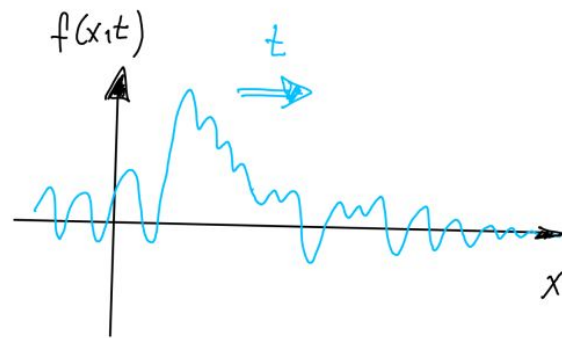
$$1D: \left(\frac{1}{v^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} \right) f(x, t) = 0$$

Vlnová rovnice 3D

$$\left(\frac{1}{v^2} \frac{\partial^2}{\partial t^2} - \Delta \right) f(x, t) = 0$$

$$\square f = 0$$

Posunují se oběma směry



$$1D: \frac{d^2}{dx^2} \quad 3D: \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \equiv \Delta$$

Vlnová rovnice

Jean le Rond d'Alembert



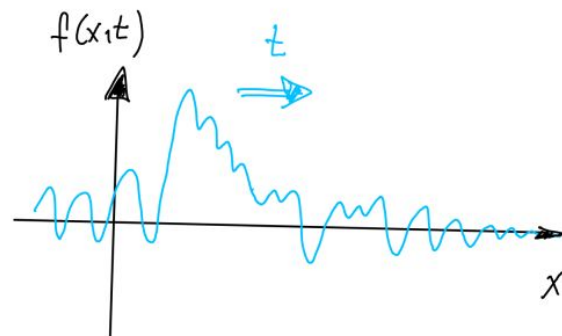
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Vlnová rovnice 3D

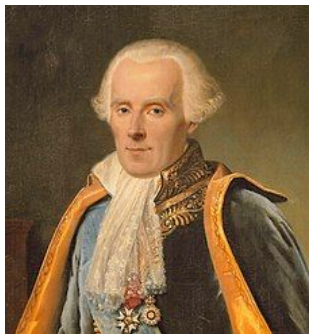
$$\left(\frac{1}{v^2} \frac{\partial^2}{\partial t^2} - \Delta \right) f(x, t) = 0$$

$$\square f = 0$$

Posunují se oběma směry



Pierre-Simon Laplace



$$1D: \frac{d^2}{dx^2}$$

$$3D: \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \equiv \Delta$$

<https://www.surfertoday.com/surfing/chicama-the-worlds-longest-wave>



Louis de Brogli

- Částice mají i vlnové vlastnosti:
- \Rightarrow Myšlenka: hmotné vlny
- Částici s hybností má i vlnovou délku:

$$p = \frac{h}{\lambda}$$



Louis de Broglie

Schrödingerova rovnice

- Ústřední nerelativistická rovnice kvantové mechaniky
- Inspirace vedla k postulátu pohybové rovnice

$$\left[-\frac{\hbar^2}{2m} \Delta + V(\vec{x}, t) \right] \Psi(\vec{x}, t) = i\hbar \frac{\partial \Psi(\vec{x}, t)}{\partial t}$$

$$\Psi : \mathbb{R}^3 \times \mathbb{R} \rightarrow \mathbb{C} \quad |\Psi|^2 \equiv \bar{\Psi} \Psi^* = \rho$$

$$\left[\hat{T} + \hat{V} \right] \bar{\Psi} = i\hbar \partial_t \bar{\Psi} \quad \hat{H} \Psi = E \Psi$$
$$V \neq V(t)$$

Schrödinger



Erwin Schrödinger



**erwin schrodinger with round glasses
in light blue suit playing with a pink
hairy cat with background of a black
board covered with equations**

Midjourney



erwin schrodinger with round glasses
in light blue suit playing with a pink
hairy cat with background of a black
board covered with the schrodinger
equation

Midjourney

Schrödingerova rovnice: řešení pro částici v krabici

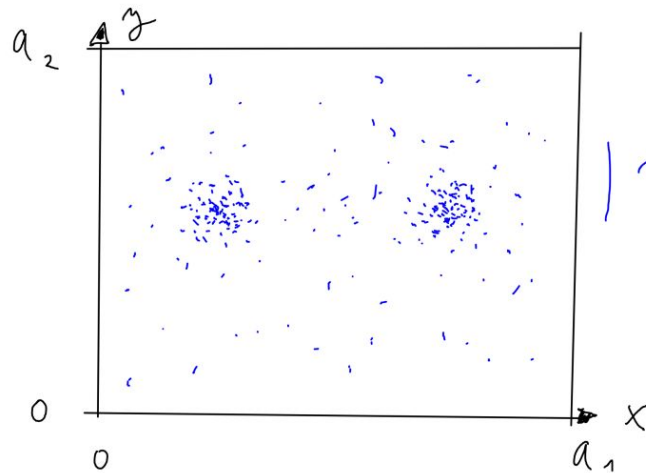
- Vlnová funkce částice je řešením bezčasové Schrödingerovy rovnice pro nulový potenciál, ale částice je omezena na určitou oblast.
- **Řešením jsou stojaté hmotné částicové vlny a diskrétní možné hodnoty energie: kvantování!**

$$-\frac{\hbar^2}{2m} \Delta \psi_n = E_n \psi_n \quad + \text{obn. podm.}$$

\Rightarrow částice v krabici:
 \Rightarrow stojaté vlny

$$E_n = \frac{(n\pi\hbar)^2}{2ma^2} = \frac{(n\pi)^2 (\hbar c)^2}{2mc^2 a^2}$$

$$\psi_{n_1 n_2}(x, y) \sim \sin \frac{n_1 \pi x}{a_1} \cdot \sin \frac{n_2 \pi y}{a_2}$$



$$\rho \sim |\psi(x, y)|^2$$

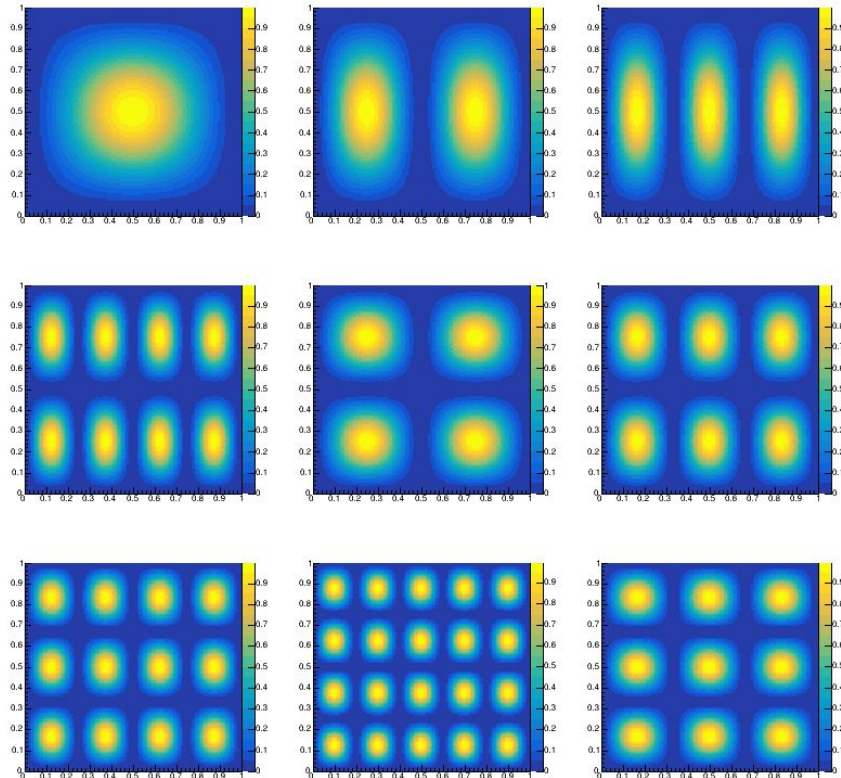
Schrödingerova rovnice: řešení pro částici v krabici

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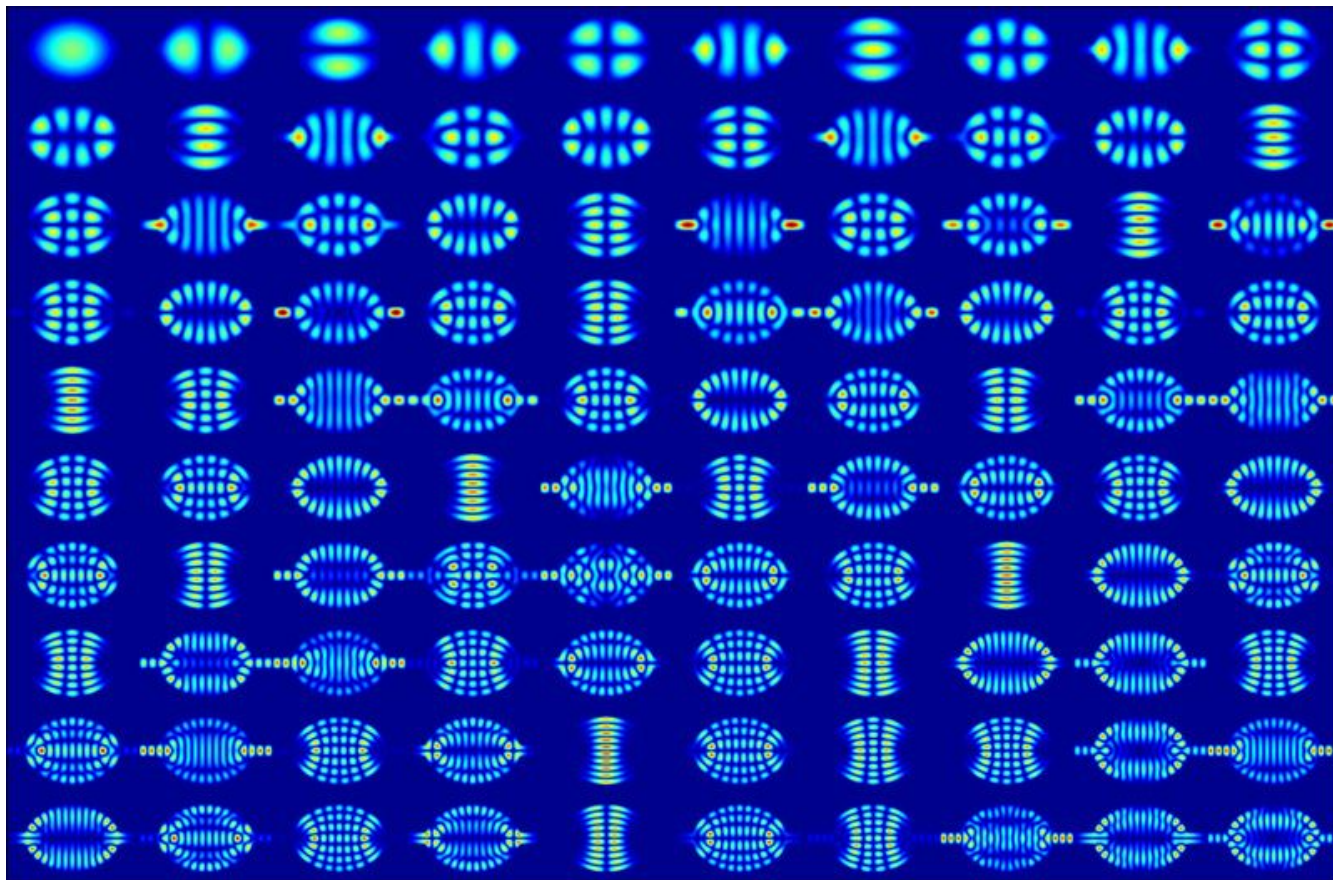
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Částice v krabici

- Kvantový billiard na elipse



Atom vodíku

- Bohrov model atomu
- Vycházel z heuristického kvantování momentu hybnosti
- Rozměr Planckovy konstanty h je rozměr momentu hybnosti:)
- Vedlo na kvantování energie elektronu v atomu a na vysvětlení spekter atomů.
- Ale nešlo zobecnit na další systémy.



Niels Bohr

$$L_n \equiv n \hbar$$

$$p = \frac{h}{\lambda}$$

$$E = E_p + E_k$$

$$p_n = \frac{h}{\lambda_n}$$

$$L_n = r_n p_n$$

Atom vodíku

- Vlnová funkce elektronu je řešením bezčasové Schrödingerovy rovnice pro Coulombický potenciál.
- Možná matematicky pěkná a fyzikální řešení jsou velmi specifická.
- Jdou ruku v ruce s omezenými a diskrétními možnými hodnotami energie: kvantování!

$$\left[-\frac{\hbar^2}{2m} \Delta + V(\vec{x}) \right] \psi(\vec{x}) = E \psi(\vec{x})$$

$$V = -\frac{1}{4\pi\epsilon_0} \frac{e^2}{r}$$

$$V = -\frac{\alpha \hbar c}{r}$$

$$\alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c} \doteq \frac{1}{137}$$

$$\hbar c \doteq 197 \text{ MeV} \cdot \text{fm}$$

Atom vodíku

- Řešit zvlášť rovnice jako funkce poloměru a úhlů
- Vyřešit kvantování momentu hybnosti.

$$\Delta = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$$

$$\Delta = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2} \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \varphi^2}$$

$$\hat{\Delta} = \Delta_r - \frac{1}{r^2} \frac{\hat{L}^2}{\hbar^2}$$

$$L^2 Y_l^m(\theta, \varphi) = \hbar^2 l(l+1) Y_l^m(\theta, \varphi)$$

$$L_z Y_l^m = m \hbar Y_l^m$$

$$Y_l^m = L_l^{|m|}(\cos \theta) e^{im\varphi}$$

Atom vodíku – poznámky k momentu hybnosti

		#	+ spin x 2	
$l = 0$... s	1	2	
$l = 1$... p	3	6	$2l + 1$
$l = 2$... d	5	10	$m = -l, \dots, l$
$l = 3$... f	7	14	

$$E_n \mapsto E_{n,l}$$

1s 2s 2p 3s 3p 4s 3d

Atom vodíku

- Rovnice nesoucí závislost na poloměru se podobá 1D Sch. rovnici a vede na kvantování energií.

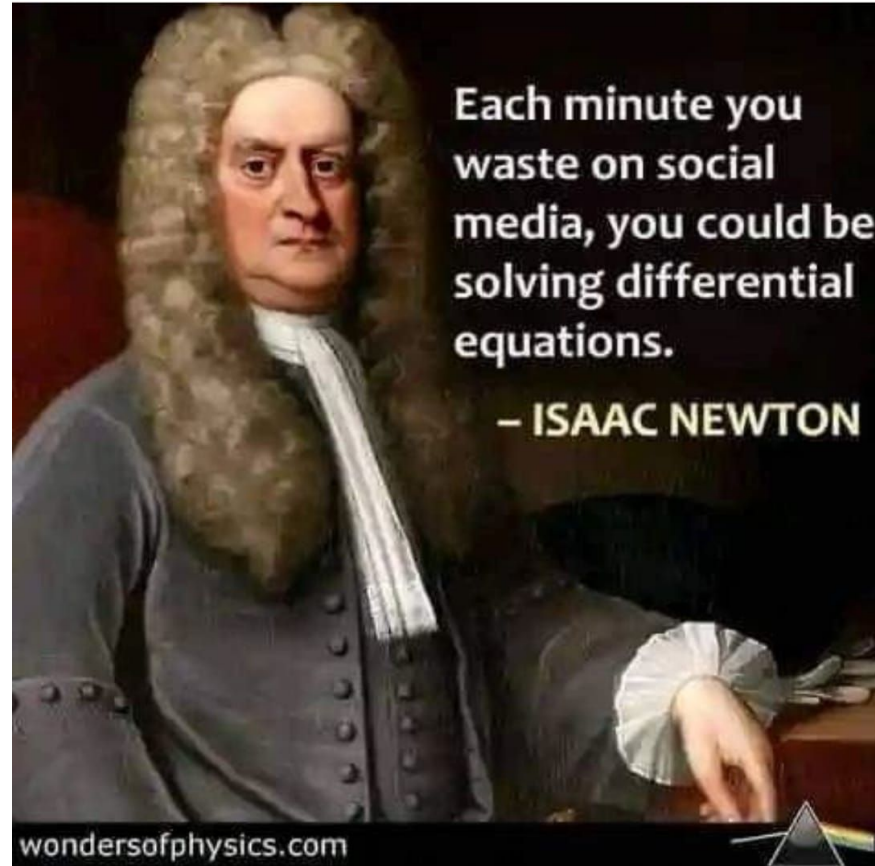
$$\left[-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + V(x) \right] \psi(x) = E \psi(x) \quad \dots \text{1D}$$

$$\left[-\frac{\hbar^2}{2m} \frac{d^2}{dr^2} + V(r) + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2} \right] u(r) = E u(r)$$
$$u(r) = \frac{R(r)}{r}$$

$$\psi(r, \theta, \varphi) \equiv R(r) Y(\theta, \varphi)$$

$$\psi_{n\ell m}(r, \theta, \varphi) \equiv R_{n\ell}(r) Y_{\ell}^m(\theta, \varphi)$$

A gentle reminder...



Each minute you
waste on social
media, you could be
solving differential
equations.

– ISAAC NEWTON

ADM no 2.2.2 23.11.2020 KA7

- apr. najpog. uporablj. pte; Coulomb'ski potencial:
 $V(r) = -\frac{e^2}{4\pi\epsilon_0} \frac{1}{r}$ $V(r) \sim -\frac{1}{r}$

Podob. E u QM pro V(r)
 → določ. radialni del: $\mu = E \mu$
 $-\frac{\hbar^2}{2m} \mu'' + \left[V(r) + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2} \right] \mu = E \mu$
 $\mu(r) \sim R = \frac{u}{r}$

Radialni del Schrödinger: $E \in \mathbb{R}$
 $[X] = [r]$ ⇒ $S = \hbar r$ $S \dots$ brezdimenzionalni

$$\mu'' + \frac{2m}{\hbar^2} \left(\frac{e^2}{4\pi\epsilon_0 r} - \left[\frac{l(l+1)}{r^2} + \chi^2 \right] \mu \right) = 0 \quad / \quad \frac{1}{\chi^2}$$

hca $-\frac{e^2}{4\pi\epsilon_0 \hbar^2} \approx \alpha$

$$\frac{1}{\chi^2} \mu'' + \frac{2m\alpha}{\hbar^2} \frac{\mu}{r} - \left[\frac{l(l+1)}{\chi^2 r^2} + 1 \right] \mu = 0$$

$S = \hbar r$

$\mu = u(r) \equiv u(r(S))$
 $u' = \frac{du}{dr} ; \frac{1}{\hbar} u'(r(S)) = \frac{du}{dr} \frac{dr}{dS} = \frac{1}{\hbar} \frac{du}{dr} = \frac{1}{\hbar} \frac{du}{dr}$
 $\rightarrow u' = \frac{1}{\hbar} \frac{du}{dr} \rightarrow u'' = \frac{1}{\hbar^2} \left(\frac{d}{dS} \right)^2$

(ob. s. asympt. ob. : $u(S) \sim e^{-S} S^{l+1}$)

Združimo radialni del in rešimo: $u(S) = e^{-S} u(S)$
 hca $u(S)$ je polinom: $u(S) = \sum_{j=0}^{\infty} c_j S^j$ $c_j \in \mathbb{C}/\mathbb{R}$

Nato določimo rešitev, ki se vedno približuje rešitvi $u(S)$
 Rešimo vs. pro $u(S)$; določimo c_j
 Navede: lastne vrednosti skrajnih polinomov ⇒ konstante!
 n.s.: $u' = \left[1 - \frac{S_0}{S} + \frac{l(l+1)}{S^2} \right] u$

Najd: u' dle (0)
 $u'' - 11 = \dots$

$$+ S_0 c_j \frac{S^{j-1}}{S} - 2(l+1) c_j S^{j-1} \} = 0$$

$$\Rightarrow \sum_{j=0}^{\infty} S^{j-1} \{ \dots \} = 0$$

skrajni del: $c_j S^{j-1}$
 $c_{j+1} (j+1) + 2(l+1) c_{j+1} (j+1) - 2 c_j + S_0 c_j$
 $- 2(l+1) c_j = 0$
 $c_{j+1} = \frac{2(l+1+j) - S_0}{(j+1)(j+2l+2)} c_j$

... rešujemo najdalj pro $c_{j+1} = 0$ c. c_j
 Asympt. $j \rightarrow \infty$: $c_{j+1} \approx \frac{2j}{(j+1)j} c_j$

$$u' = e^{-S} S^{l+1} u$$

$$u'' = e^{-S} \left[-S^{l+1} u' + (l+1) S^l u' + S^{l+1} u'' \right]$$

$$= e^{-S} S^l \left[(l+1-S) u' + S u'' \right]$$

$$u'' = e^{-S} \left[-S^{l+1} u' - S^l e^{-S} \dots \right] + \left[S^l e^{-S} \left[-u''(l+1-S) u' + u' + S u'' \right] \right]$$

$$= e^{-S} \left[\frac{2-l}{S} \left\{ (l-S) \left[(l+1-S) u' + S u'' \right] - S u'' + (l+1-S) S u'' \right. \right. \right.$$

$$\left. \left. + S u'' + S^2 u'' \right\} \right]$$

$$= e^{-S} \left[\frac{2-l}{S} \left\{ (l+1-S) \left[(l-S) - 1 \right] u' + u'' \left[-l \cdot S + l+1-S+1 \right] + S u'' \right\} \right]$$

$$= \frac{l(l+1) - l - l+1+S - 1}{S} = \frac{l(l+1) - 2l - 2 + S}{S}$$

Asympt. $j \rightarrow \infty$, l odd. navede $S \sim \nu$ vs. $\nu = \sum c_j S^j$
 $c_{j+1} = c_j \frac{2}{j} \Rightarrow c_j = c_0 \frac{2^j}{j!}$

hca: $u \approx c_0 \sum_{j=0}^{\infty} \frac{2^j S^j}{j!} = c_0 \sum_{j=0}^{\infty} (2S)^j = c_0 e^{2S}$

hca: $u(S) = e^{-S} S^{l+1} u$
 potrdi asympt. ob. in rešimo radialni del in rešimo vs. pro $S \rightarrow \infty$
 določimo u dle polinomov: $c_{j+1} = 0$

$u(S) = u(S)$
 $u' + \frac{S_0}{S} u - \left(\frac{l(l+1)}{S^2} + 1 \right) u = 0$
 $S \in (0, \infty)$

Hermitjeva simetrija: $S \rightarrow 0$; $S \rightarrow \infty$ določimo rešitve
 n. o. d. r.: $\frac{1}{S} \sim \frac{1}{S^2}$
 $u(S) = A e^{-S} + B e^{-S}$
 $u \sim e^{-S}$

Najd: $u'' = \left[\frac{l(l+1)}{S} - 2l - 2 + S \right] u' + 2(l-S+1) u'' + S u''$
 Najd: Rešimo dle: $u' = \left[1 - \frac{S_0}{S} + \frac{l(l+1)}{S^2} \right] u$
 + določimo e^{-S} ; določimo S^l

$$\left(\frac{l(l+1)}{S} - 2l - 2 + S \right) u' + 2(l-S+1) u'' + S u'' = \left[1 - \frac{S_0}{S} + \frac{l(l+1)}{S^2} \right] S u'$$

$$u'' + 2 \frac{l+1-S}{S} u' + \left[\frac{-2l-2+S}{S} - 1 + \frac{S_0}{S} \right] u = 0$$

$$u'' + 2 \frac{l+1-S}{S} u' + \left(\frac{S_0}{S} - 2 \frac{l+1}{S} \right) u = 0$$

$$c_{j+1} = c_j \frac{2(l+j+1) - S_0}{(j+1)(j+2l+2)}$$

$$c_{j+1} = 0 \quad ; \quad c_j \neq 0$$

$$2(l+j+1) - S_0 = 0$$

$$\Rightarrow n = 1, 2, \dots$$

ob.: $S_0 = \frac{2m\epsilon_0}{\hbar^2} \alpha$ $2n = S_0$
 ob.: $\chi^2 = -\frac{2m\epsilon_0}{\hbar^2} \alpha$ $\Rightarrow 4n^2 = S_0^2$

$S \rightarrow 0$: $u \sim \frac{l(l+1)}{S^2} u$
 določimo: $u \sim S^{l+1}$: $u' = (l+1) S^l \sim \frac{l+1}{S} u$
 $u'' = l(l+1) S^{l-1} \sim \frac{l(l+1)}{S^2} u$
 ob.: $u \sim S^l$ $u' \sim l S^{l-1} \sim -l \frac{u}{S}$
 $u'' \sim (-l)(-l-1) S^{l-2} \sim l(l+1) \frac{u}{S^2}$

dle: $S \rightarrow 0$: $u(S) = C S^{l+1} + D S^{-l}$
 ob.: $l = 0, 1, \dots$
 $R \sim \frac{u}{r}$

Najd: $u = \sum_{j=0}^{\infty} c_j S^j$
 $u' = \sum_{j=0}^{\infty} j c_j S^{j-1} = \sum_{j=1}^{\infty} c_{j-1} (j-1) S^{j-1}$
 $u'' = \sum_{j=0}^{\infty} c_j j(j-1) S^{j-2} = \sum_{j=2}^{\infty} c_{j-2} (j-2)(j-1) S^{j-2}$

Najdimo: n. o. d. r. pro u določimo lastne vrednosti u in določimo u
 Določimo dle (0)
 $\sum_{j=0}^{\infty} \left\{ c_{j+1} j(j+1) S^{j-1} + 2(l+1) c_{j+1} (j+1) S^{j-1} - 2 c_j S^{j-1} + \dots \right\}$

$2n = S_0$ / (1)
 $4n^2 = S_0^2 = \left(\frac{2m\epsilon_0}{\hbar^2} \alpha \right)^2$ $\chi^2 = -\frac{2m\epsilon_0}{\hbar^2} \alpha$
 $4n^2 = \frac{4m^2 \epsilon_0^2}{\hbar^4} \alpha^2$ $\frac{-\hbar^2}{2m\epsilon_0}$

$E_n = -\frac{1}{2} \frac{m\epsilon_0}{\hbar^2} \cdot \alpha^2$
 ... določimo lastne vrednosti u (n = 1, 2, ...)
 rešimo radialni del in rešimo vs. pro $S \rightarrow \infty$
 določimo u dle polinomov: $c_{j+1} = 0$

1

2

3

4

5

6

7

9

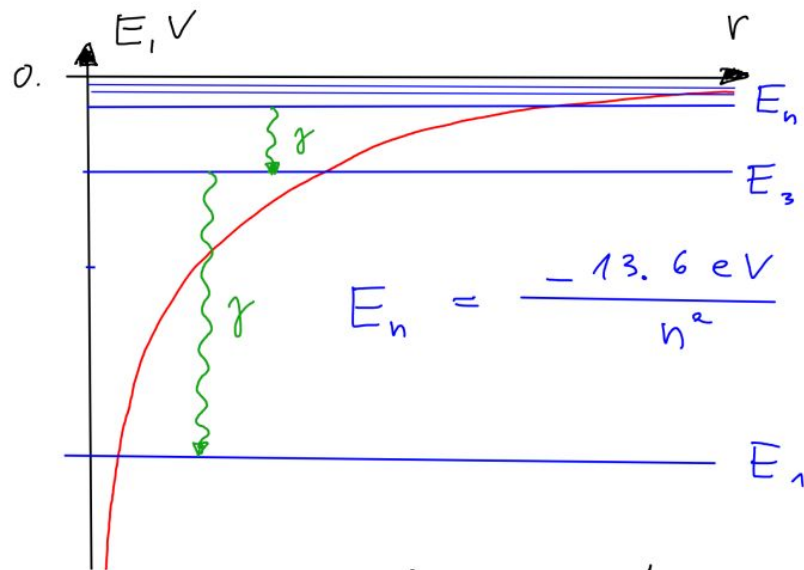
10

11

12

Atom vodíku

- Schr. rovnice řešitelná jen pro určité energie: kvantování!
- Vysvětlení diskretních spekter atomů a molekul! – přechody mezi hladinami.



$$V(r) \sim \frac{-e^2}{r}$$

vázaný stav

$$E = E_k + E_p < 0$$

$$E_\gamma = E_m - E_n = \frac{13.6 \text{ eV}}{n^2 - m^2}$$

$$\alpha \equiv \frac{e^2}{4\pi\epsilon_0 \hbar c} \doteq \frac{1}{137}$$

$$E_n = -\frac{1}{n^2} \frac{1}{2} \alpha^2 m_e c^2$$

$$m_e c^2 \doteq 511 \text{ keV}$$

Atom vodíku

$$\left[-\frac{\hbar^2}{2m} \Delta + V(\vec{x}) \right] \psi(\vec{x}) = E \psi(\vec{x})$$

$$\psi(\vec{x}) \equiv \psi_{n\ell m}(\vec{x}) = R_{n\ell}(r) Y_{\ell m}(\theta, \varphi)$$

$$\equiv \psi_{n\ell m}(r, \theta, \varphi)$$

- Určitým energiím přísluší určité vlnové funkce.
- Ale k čemu je vlnová funkce?

Atom vodíku

- Určitým energiím přísluší určité vlnové funkce.
- Ale k čemu je vlnová funkce?
- Vlnová funkce \Rightarrow pravděpodobnost, s jakou elektron nalezneme okolo jádra.



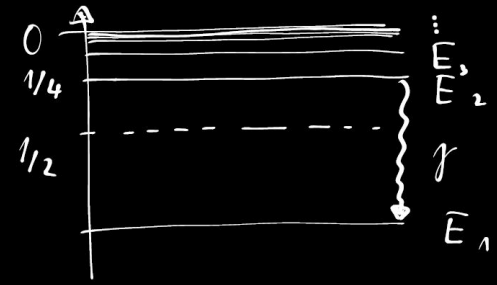
Max Born

$$\psi = \psi(x, y, z) = \psi(r, \theta, \varphi)$$

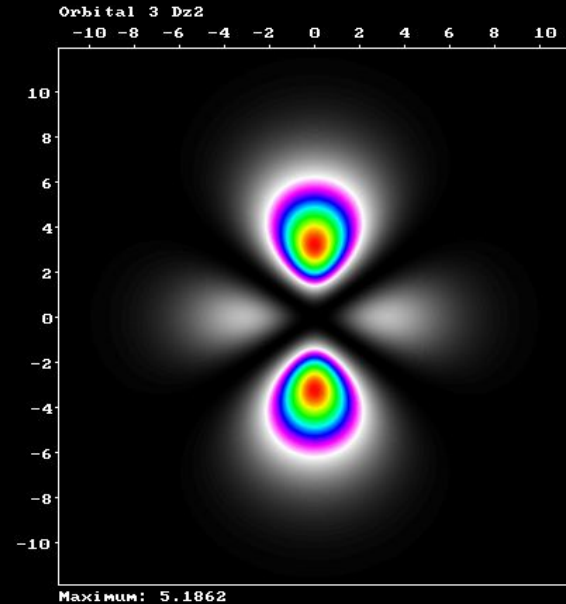
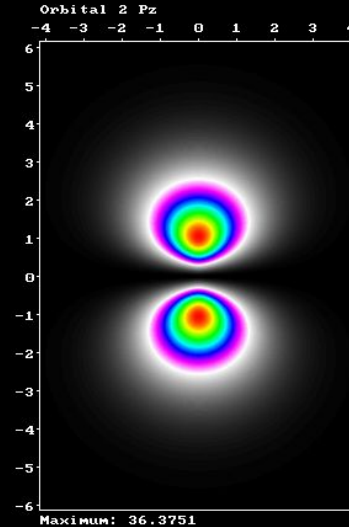
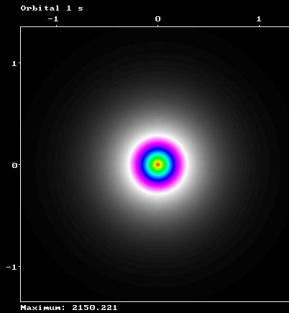
$$\rho(\vec{x}) = |\psi(\vec{x})|^2 \quad \text{hustota pravděpodobnosti}$$

Atom vodíku

- Různé pravděpodobnosti nalezení elektronu okolo jádra
 - pro různé energie, hodnoty dalších kvantových čísel
 - jde o různé vlnové funkce, tj. stavy elektronu v atomu.
 - Elektron je stojatá částicová vlna okolo jádra.
 - Mikrosvět a kvantová mechanika jsou pravděpodobnostní!
 - Albert Einstein: Bůh nehraje v kostky!



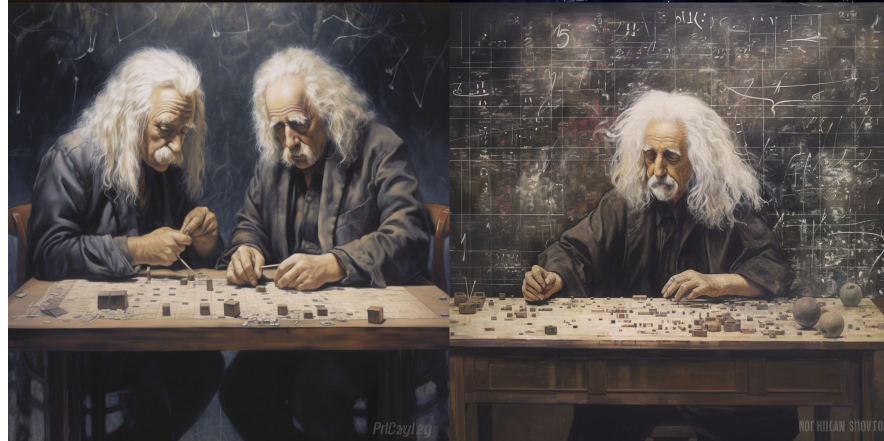
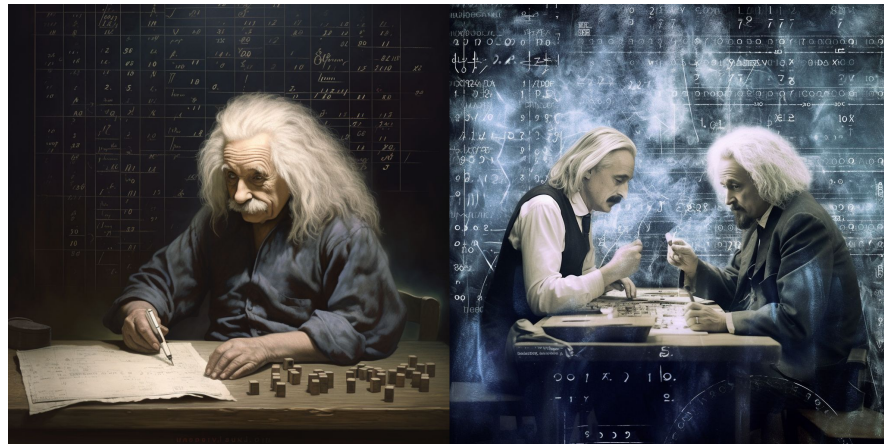
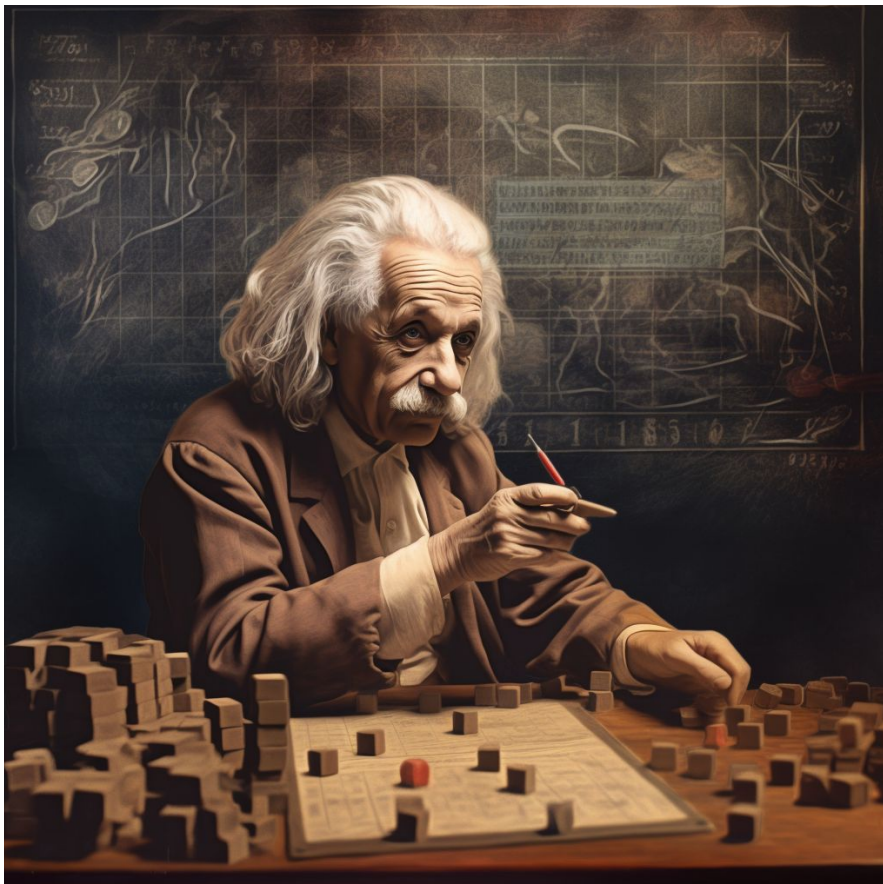
$$E_n = - \frac{13.6 \text{ eV}}{n^2}$$

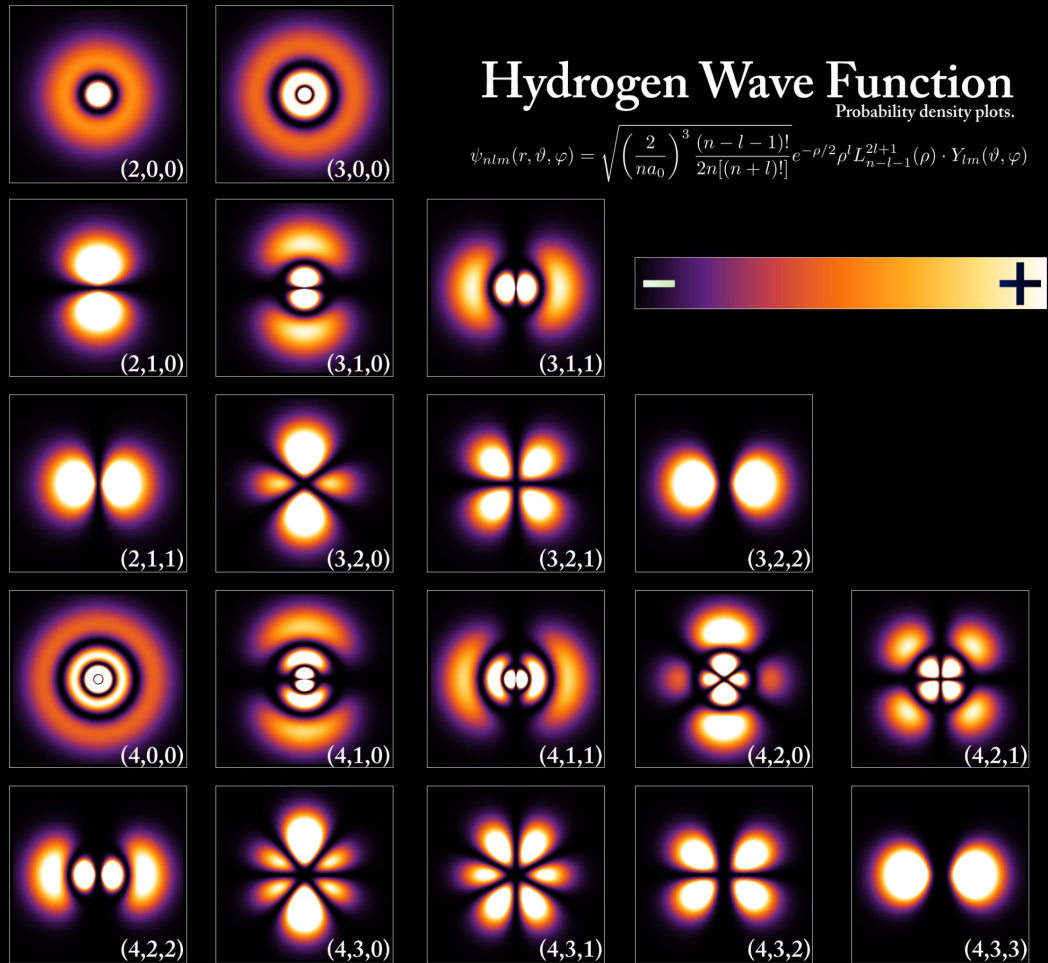


“Bůh nehraje v kostky”, Albert Einstein

Midjourney

Albert Einstein playing dice with the God at a table with background of black boards covered with equations





$$1s = \frac{1}{\pi^{1/2}} \left(\frac{Z}{a} \right)^{3/2} e^{-Zr/a}$$

$$2s = \frac{1}{4(2\pi)^{1/2}} \left(\frac{Z}{a} \right)^{3/2} \left(2 - \frac{Zr}{a} \right) e^{-Zr/2a}$$

$$2p_z = \frac{1}{4(2\pi)^{1/2}} \left(\frac{Z}{a} \right)^{5/2} r e^{-Zr/2a} \cos \theta$$

$$2p_x = \frac{1}{4(2\pi)^{1/2}} \left(\frac{Z}{a} \right)^{5/2} r e^{-Zr/2a} \sin \theta \cos \phi$$

$$2p_y = \frac{1}{4(2\pi)^{1/2}} \left(\frac{Z}{a} \right)^{5/2} r e^{-Zr/2a} \sin \theta \sin \phi$$

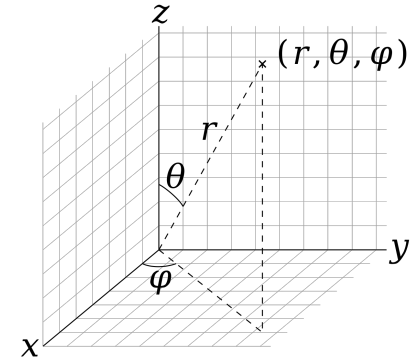
$$3s = \frac{1}{81(3\pi)^{1/2}} \left(\frac{Z}{a} \right)^{3/2} \left(27 - 18 \frac{Zr}{a} + 2 \frac{Z^2 r^2}{a^2} \right) e^{-Zr/3a}$$

$$3p_z = \frac{2^{1/2}}{81\pi^{1/2}} \left(\frac{Z}{a} \right)^{5/2} \left(6 - \frac{Zr}{a} \right) r e^{-Zr/3a} \cos \theta$$

$$3p_x = \frac{2^{1/2}}{81\pi^{1/2}} \left(\frac{Z}{a} \right)^{5/2} \left(6 - \frac{Zr}{a} \right) r e^{-Zr/3a} \sin \theta \cos \phi$$

$$3p_y = \frac{2^{1/2}}{81\pi^{1/2}} \left(\frac{Z}{a} \right)^{5/2} \left(6 - \frac{Zr}{a} \right) r e^{-Zr/3a} \sin \theta \sin \phi$$

$$3d_z^2 = \frac{1}{81(6\pi)^{1/2}} \left(\frac{Z}{a} \right)^{7/2} r^2 e^{-Zr/3a} (3 \cos^2 \theta - 1)$$



https://en.wikipedia.org/wiki/Spherical_coordinate_system

<https://www.pearsonhighered.com/assets/samplechapter/0/3/2/1/0321803450.pdf>

Závěr I

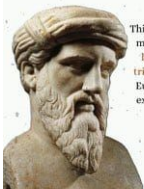
$$\vec{p} = m \frac{d^2}{dt^2} \vec{x}(t)$$

$$\left[-\frac{\hbar^2}{2m} \Delta + V(\vec{x}, t) \right] \Psi(\vec{x}, t) = i\hbar \frac{\partial \Psi(\vec{x}, t)}{\partial t}$$



PYTHAGOREAN THEOREM (GEOMETRY)

$$a^2 + b^2 = c^2$$



This theorem, proposed by ancient Greek mathematician Pythagoras, relates the lengths of the sides of a right-angled triangle. It is a fundamental principle in Euclidean geometry and has been used extensively in various fields, including physics, computer graphics, and engineering design.



FOURIER TRANSFORM (SIGNAL PROCESSING)

$$F(\omega) = \int_{-\infty}^{\infty} f(t)e^{-i\omega t} dt$$



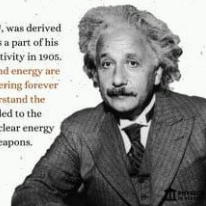
Developed by Joseph Fourier in the early 19th century, the Fourier Transform allows a mathematical function to be decomposed into its constituent frequencies. This technique is fundamental in the field of signal processing and has found applications in areas like image processing, audio signal processing, and solving partial differential equations.



EINSTEIN'S EQUATION (SPECIAL RELATIVITY)

$$E = mc^2$$

This equation, $E=mc^2$, was derived by Albert Einstein as a part of his theory of special relativity in 1905. It states that mass and energy are interchangeable, altering forever the way we understand the universe. It also led to the development of nuclear energy and nuclear weapons.



HEISENBERG'S UNCERTAINTY PRINCIPLE (QUANTUM MECHANICS)

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

Formulated by Werner Heisenberg in 1927, the uncertainty principle is one of the core tenets of quantum mechanics. It states that the more precisely the position of a particle is determined, the less precisely its momentum can be known, and vice versa. It underpins the inherently probabilistic nature of quantum mechanical phenomena.



DIRAC EQUATION (QUANTUM FIELD THEORY)

$$(i\hbar\gamma^\mu\partial_\mu - mc)\psi = 0$$



Proposed by Paul Dirac in 1928, this equation describes fermions—particles with half-integer spin—and led to the prediction of antimatter. It's critical in quantum field theory and has shaped our understanding of the fundamental nature of particles and the universe.



EULER'S IDENTITY (COMPLEX ANALYSIS)

$$e^{i\pi} + 1 = 0$$

Named after the Swiss mathematician Leonhard Euler, this identity is a special case of Euler's formula from complex analysis, and it beautifully combines five of the most important numbers in mathematics: 0, 1, π , e , and i . It's considered by many to be the most beautiful equation in mathematics due to its elegant linking of these fundamental numbers.



MAXWELL'S EQUATIONS (ELECTRODYNAMICS)

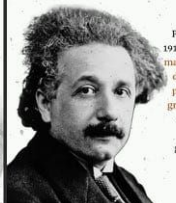
$$\begin{aligned} \nabla \cdot \mathbf{D} &= \rho \\ \nabla \cdot \mathbf{B} &= 0 \\ \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \times \mathbf{H} &= \mathbf{J} + \frac{\partial \mathbf{D}}{\partial t} \end{aligned}$$

Formulated by James Clerk Maxwell and Oliver Heaviside in 1862, these four equations describe how electric and magnetic fields interact. They underpin all of classical electrodynamics, optics, and electric circuits and paved the way for the advent of technologies such as radio, television, and radar.



EINSTEIN'S FIELD EQUATIONS (GENERAL RELATIVITY)

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + g_{\mu\nu}\Lambda = \frac{8\pi G}{c^4}T_{\mu\nu}$$



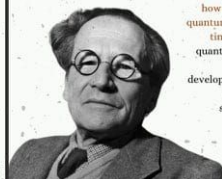
Published by Albert Einstein in 1915, these equations describe how matter and energy in the Universe distort spacetime to create the phenomenon we experience as gravity. They are fundamental to cosmology and astrophysics, predicting phenomena like gravitational waves and black holes.



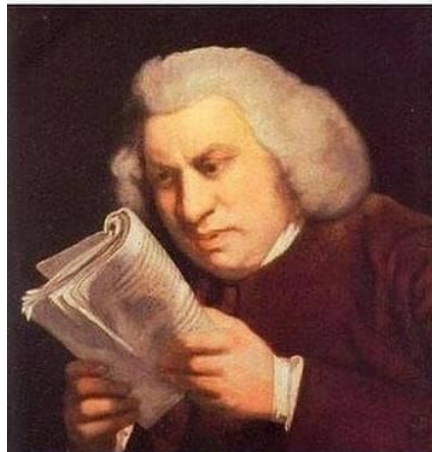
SCHRODINGER EQUATION (QUANTUM MECHANICS)

$$\hat{H}\Psi = i\hbar\frac{\partial\Psi}{\partial t}$$

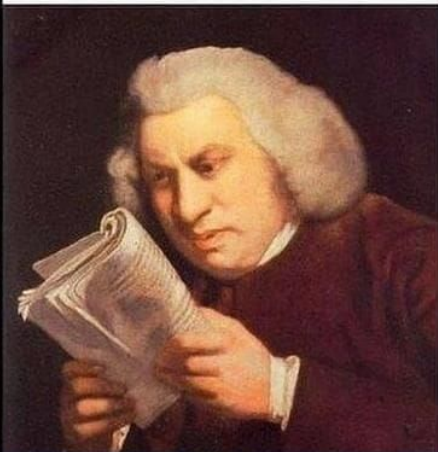
Proposed by Erwin Schrödinger in 1926, this equation describes how the quantum state of a quantum system changes with time. It forms the basis of quantum mechanics and has been crucial in the development of many modern technologies, from semiconductors to MRI scanners.



The first time you read about quantum mechanics:



The 1000th time you read about quantum mechanics:



Twitter Books @TwitterBooks

Last book that made you cry

4:52 PM · 1/27/20 · Sprinklr

23 Retweets 144 Likes

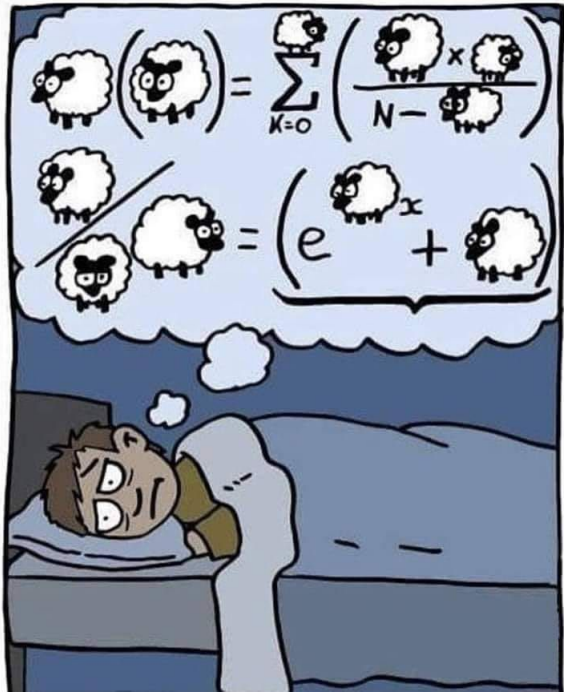
Replying to @TwitterBooks

University Physics with Modern Physics
14th Edition by Hugh D. Young, Roger A. Freedman

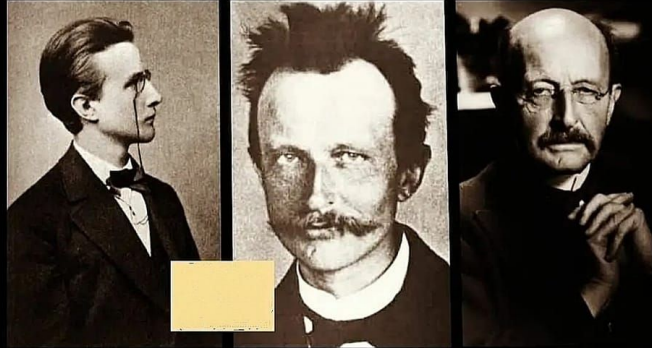
1 3

Roger Freedman @RogerFreed... · 5h

No doubt tears of joy.



Max Planck



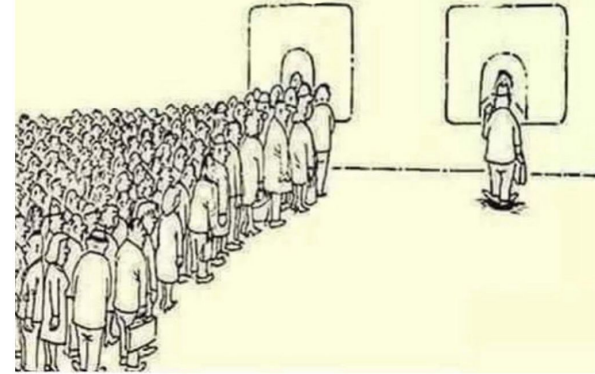
Before
discovery
of
Quantum
Physics

During
discovery
of
Quantum
Physics

After
discovery
of
Quantum
Physics

Students interested in
black hole, worm hole,
big bang, string theory,
time travel, etc.

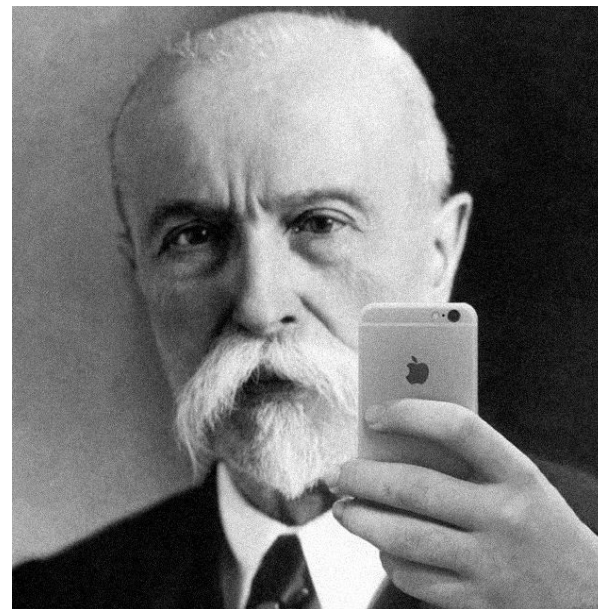
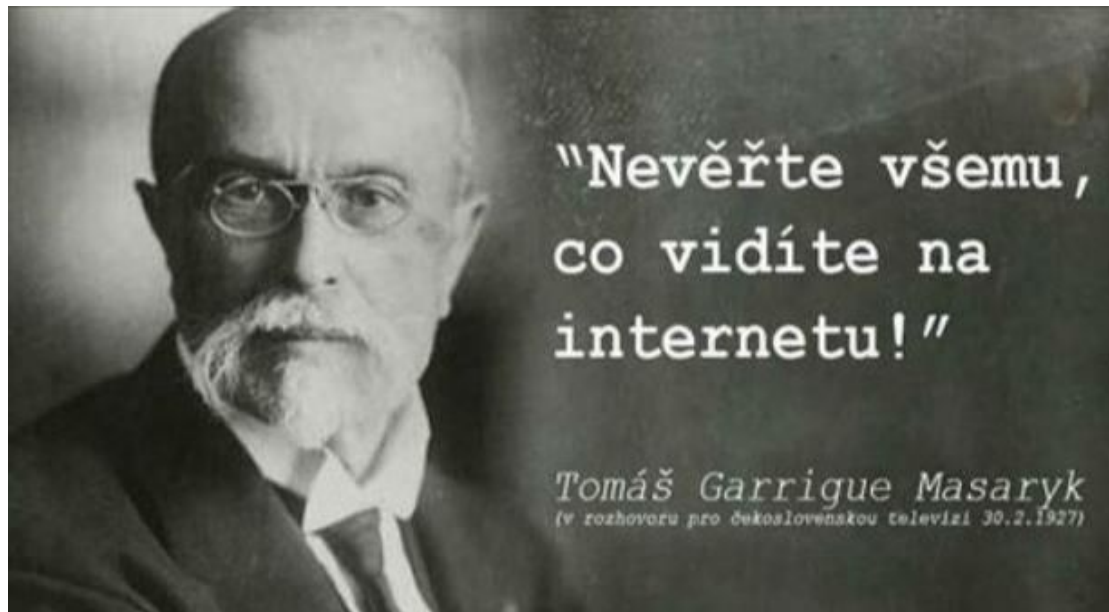
Those who
actually study
physics and math
at university level



wondersofphysics.com

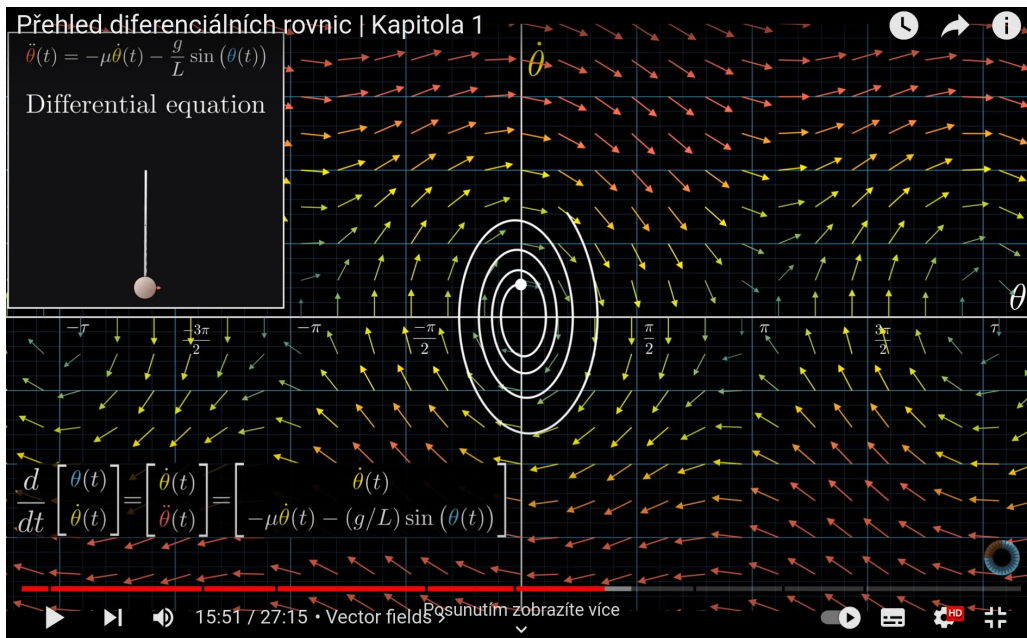


A gentle reminder...



Diferenciální rovnice @ 3B1B

https://www.youtube.com/watch?v=p_di4Zn4wz4&t=802s&ab_channel=3Blue1Brown

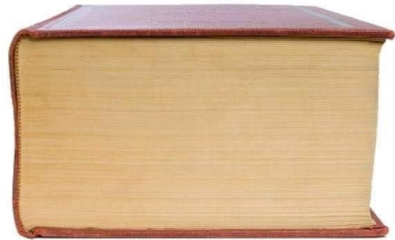


Backup

1927 :: https://en.wikipedia.org/wiki/Solvay_Conference

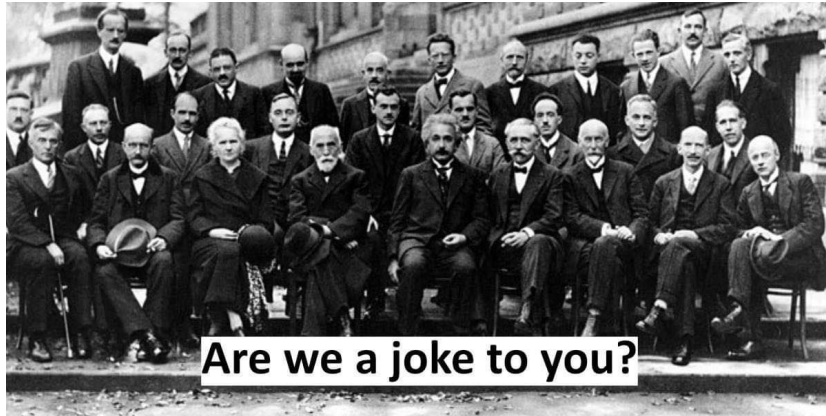
Midjourney

Physics



Sarcasm

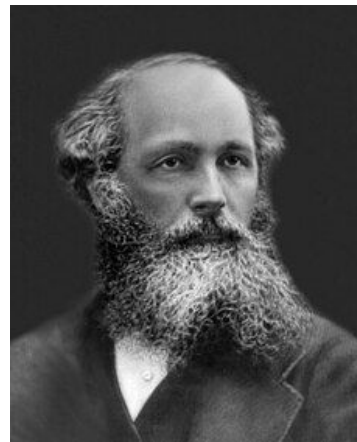
Physics if Newton slept under Coconut tree



Are we a joke to you?

Maxwellovy rovnice

- Rovnice popisující elektromagnetická pole.
- Popisují mj. i elektromagnetické vlny.
- Jsou invariantní vůči relativistické Lorentzově transformaci!



James Clerk Maxwell

$$\begin{aligned} \operatorname{div} \vec{B} &= 0 & \operatorname{div} \vec{E} &= \frac{\rho}{\epsilon_0} \\ \operatorname{rot} \vec{B} &= \mu_0 \left(\epsilon_0 \frac{\partial \vec{E}}{\partial t} + \vec{j} \right) & \operatorname{rot} \vec{E} &= -\frac{\partial \vec{B}}{\partial t} \end{aligned}$$

Maxwellovy rovnice

$$A^\mu = \left(\frac{1}{c} \phi, \vec{A} \right)$$

$$\vec{B} = \text{rot } \vec{A}$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\vec{E} = -\text{grad } \phi - \frac{\partial \vec{A}}{\partial t}$$

$$\partial_\alpha F^{\alpha\beta} = \vec{J}^\beta$$

Vákuum: $\partial_\alpha F^{\alpha\beta} = 0$

$$F^{\mu\nu} = \begin{pmatrix} 0 & -\frac{E_x}{c} & -\frac{E_y}{c} & -\frac{E_z}{c} \\ \frac{E_x}{c} & 0 & -B_z & B_y \\ \frac{E_y}{c} & B_z & 0 & -B_x \\ \frac{E_z}{c} & -B_y & B_x & 0 \end{pmatrix}$$

Pauliho rovnice

- Rozšíření nerelativistické Schrödingerovy rovnice pro částici se spinem $\frac{1}{2}$
- Vlnová funkce má 2 komponenty popisující 2 možné stavy (projekce) spinu.



Wolfgang Pauli

$$\begin{aligned} \text{Pauli: } H &= \frac{1}{2m} \left[(\vec{p} - q\vec{A})^2 - \hbar q \vec{\sigma} \cdot \vec{B} \right] + q\phi \\ &= \frac{1}{2m} \left[\vec{\sigma} \cdot (\vec{p} - q\vec{A}) \right]^2 + q\phi \end{aligned}$$

Klein – Gordonova rovnice – relativistická

- Relativistická rovnice popisující částici se spinem 0.
 - Inspirací je vztah $E^2 = m^2c^4 + p^2c^2$.
 - Spin je vnitřní moment hybnosti částice
 - Ale elektron má spin $\frac{1}{2}$...
 - A rovnice má některé nehezké vlastnosti...



Oskar Benjamin Klein

$$\hat{\vec{p}} = -i\hbar \nabla$$

$$\hat{\vec{p}}^2 = -\hbar^2 \Delta$$

$$E^2 = m^2c^4 + \vec{p}^2c^2$$

$$E \sim i\hbar \frac{\partial}{\partial t}$$

$$i^2 = -1$$

$$\Rightarrow \left[-\hbar^2 c^2 \Delta + m^2 c^4 \right] \psi = -\hbar^2 \frac{\partial^2}{\partial t^2} \psi$$

Klein – Gordonova rovnice – relativistická



Oskar Benjamin Klein

- Relativistická rovnice popisující částici se spinem 0.
 - Inspirací je vztah $E^2 = m^2c^4 + p^2c^4$.
 - Spin je vnitřní moment hybnosti částice
 - Ale elektron má spin $\frac{1}{2}$...
 - A rovnice má některé nehezké vlastnosti...

Klein - Gordon:
$$\frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2} - \Delta \psi + \frac{m^2 c^2}{\hbar^2} \psi = 0$$

$$\left[\frac{1}{c^2} \partial_t^2 - \Delta + \left(\frac{mc}{\hbar} \right)^2 \right] \psi = 0$$

$\hbar = c = 1$:
$$(\square + m^2) \psi = 0$$

Diracova rovnice – relativistická rovnice QM

- Ukazuje se, že popisuje částici se spinem $\frac{1}{2}$
 - tj. např. elektron
 - vlnová funkce zde má 4 komponenty
 - předpovídá existenci antičástic
- Demokratická v prostoru a čase, první derivace podle obou.

$$\left[\vec{\alpha} \cdot \hat{\vec{p}} c + \beta m c^2 \right] \psi(\vec{x}, t) = i \hbar \frac{\partial \psi(\vec{x}, t)}{\partial t}$$

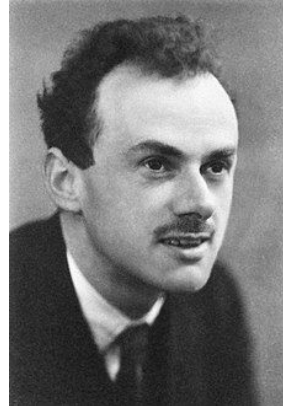
$$\left[-i \hbar c \vec{\alpha} \cdot \nabla + \beta m c \right] \psi = i \hbar \frac{\partial}{\partial t} \psi$$

$$\nabla \equiv \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

$\vec{\alpha}, \beta$... matrice

$$\psi \equiv \psi(\vec{x}, t) = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}$$

Diracova rovnice – relativistická rovnice QM



Paul A. M. Dirac

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- předpovídá existenci antičástic

$$D_{\text{Dirac}} \left(i\hbar \gamma^\mu \partial_\mu - mc \right) \psi = 0$$

$$\hbar = c^{-1} \quad (i \not{\partial} - m) \psi = 0$$

$$H = \vec{\alpha} \cdot \vec{p}c + \beta mc^2 \quad ; \quad H^2 = p^2 c^2 + m^2 c^4$$

$$\sqrt{a^2 + b^2} = a + b \quad :-)$$

$$\vec{\gamma} = \beta \vec{\alpha} \quad \gamma^0 = \beta$$

$$\gamma^\mu = (\gamma^0, \vec{\gamma}) \quad \alpha^\mu \gamma_\mu \equiv \not{\alpha}$$

$$\psi = \begin{pmatrix} \vdots \\ \vdots \end{pmatrix} \quad \psi: \mathbb{R}^{3,1} \mapsto \mathbb{C}^4$$

Princip stacionární akce

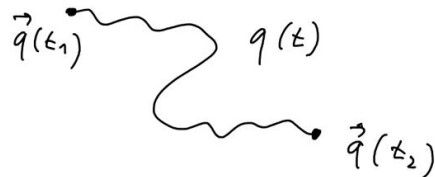
- Akce je integrál z Lagrangiánu
- Lagrangián je rozdílem kinetické a potenciální energie
 - ale je to funkce funkcí popisujících trajektorii:)
- Princip vede na Euler-Lagrangeovy rovnice, které vedou na Newtonovy.
 - vyhýbají se konceptu síly, pracují s potenciálem!

$$\delta \mathcal{S} = 0 \quad \mathcal{S} = \int_{t_1}^{t_2} L(q, \dot{q}, t) dt$$

$$\mathcal{S} = \int \mathcal{L}(\phi, \partial_\sigma \phi, t) d^4x$$

$$L = T - V$$

$$\mathcal{L} = \mathcal{T} - \mathcal{V}$$



Joseph-Louis Lagrange

Princip stacionární akce

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Joseph-Louis Lagrange

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = 0$$

$$\frac{\partial}{\partial r} \frac{\partial L}{\partial \dot{\phi}} - \frac{\partial L}{\partial \phi} = 0$$

Hamiltonovy rovnice

- Zavádějí hybnost jako nezávislou proměnnou
- Hamiltonián je funkcionál celkové energie
- Cesta k zobecnění do kvantové mechaniky! :-)

$$p \equiv \frac{\partial L}{\partial \dot{q}}$$

$$H = p\dot{q} - L \equiv H(p, q, t)$$

$$\frac{\partial H}{\partial q} = -\dot{p} \quad \frac{\partial H}{\partial p} = \dot{q}$$

$$H = T + V = T(p) + V(q)$$

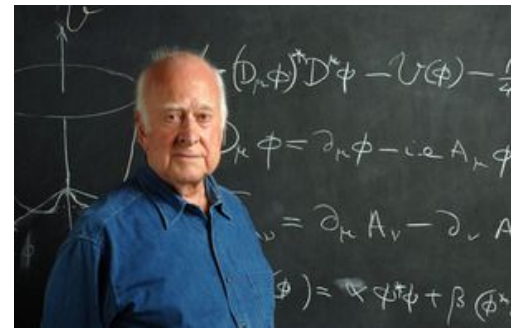
$$H = \frac{p^2}{2m} + V(q) \quad E_k \equiv T = \frac{1}{2} m v^2 = \frac{p^2}{2m}$$



Sir William Rowan Hamilton

Standardní Model

- Lagrangián Standardního Modelu
- Kinetické a interakční členy
- Higgsův potenciál
- Seldon Glashow, Steven Weinberg, David Politzer, David Gross, Frank Wilczek, Abdus Salam



$$\mathcal{L} = -\frac{1}{4} F^{\mu\nu} \cdot F_{\mu\nu}$$

$$+ i \bar{\psi} \not{D} \psi + h.c.$$

$$+ \bar{\Psi}_i \gamma_{ij} \Psi \phi + h.c.$$

$$+ (D_\mu \phi)^\dagger (D^\mu \phi)$$

$$+ \mu^2 \phi^\dagger \phi - \lambda (\phi^\dagger \phi)^2$$

